SYNTHETIC APERTURE RADAR IMAGING WITH DEEP NEURAL NETWORKS

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ABSTRACT

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Keywords: Synthetic aperture radar, inverse problems, computational imaging, deep learning, convolutional neural networks, plug-and-play priors, automatic target recognition.

Synthetic aperture radar (SAR) is a remote sensing imaging modality that has been in use since the 1960s. Conventional image formation in SAR is based on 2D inverse Fourier transform of the reflectivity field of the scene to be imaged. This conventional image formation technique is developed for a clean and complete data collection scenario. However, in reality, the collected data are only a reduced representation of the underlying scene due to hardware limitations and uncertainties in the data collection geometry, and hence suffer from reduction and phase errors. Therefore, many SAR image formation frameworks using regularization have been proposed over the years, in order to account for these limitations.

In this dissertation, we have focused on the SAR imaging problem, particularly image formation, phase error correction, and automatic target recognition (ATR), and developed three frameworks. The first framework tackles the SAR image formation problem. In this framework, SAR image formation is formulated as a regularized optimization problem, and using the plug-and-play (PnP) priors framework, we have incorporated deep learning-based priors into our formulation. Our second framework is an extension of the first one, which aims at joint image formation and phase error correction. Experimental results show the effectiveness of these two frameworks and

our proposed methods exceed the state-of-the-art image formation and phase error correction performances in the majority of the scenarios considered. The third proposed framework focuses on the ATR problem, and within this framework two ATR approaches are presented which perform the ATR task in the data domain rather than image domain. We have experimentally shown that the ATR task can be successfully performed in the data domain, and with further development, it might be possible to reach state-of-the-art performance.

Overall, we have shown that the performance in various SAR imaging tasks can be improved significantly using deep learning tools.

ÖZET

DERİN SİNİR AĞLARI İLE SENTETİK AÇIKLIKLI RADAR GÖRÜNTÜLEME

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ELEKTRONİK MÜHENDİSLİĞİ DOKTORA TEZİ, Aralık 2020

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Anahtar Kelimeler: Sentetik açıklıklı radar, ters problemler, hesaplamalı görüntüleme, derin öğrenme, evrişimsel sinir ağları, tak-çalıştır önseller, özdevimli hedef tanılama.

Sentetik açıklıklı radar (SAR), 1960'lardan beri kullanımda olan bir uzaktan görüntüleme yaklaşımıdır. Geleneksel SAR görüntü oluşturma yöntemi, görüntülenecek sahnenin yansıtma alanının 2B Fourier dönüşümü tabanlıdır. Bu geleneksel görüntü oluşturma yöntemi, temiz ve tam veri toplama senaryosu için geliştirilmiştir. Ancak, gerçekte toplanan veri, donanım kısıtları ve veri toplama geometrisindeki belirsizliklerden ötürü, altta yatan sahnenin ancak indirgenmiş bir temsilidir ve bunlar veri azaltımı ve faz hatalarına sebep olmaktadır. Bundan dolayı, bu kısıtları denkleştirme adına yıllar içinde düzenlileştirme kullanan birçok SAR görüntü oluşturma çerçeveleri önerilmiştir.

Bu tezde, SAR görüntüleme problemine odaklanılmıştır ve görüntü oluşturma, faz hatası düzeltimi ve özdevimli hedef tanılama (ATR) için üç çerçeve geliştirilmiştir. İlk çerçeve, SAR görüntüleme problemi odaklıdır. Bu çerçevede SAR görüntüleme, düzenlileştirmeli bir eniyileme problemi olarak kurgulanmıştır ve tak-çalıştır (PnP) önseller çerçevesi kullanılarak, derin öğrenme tabanlı önseller kurgulamaya dahil edilmiştir. İkinci çerçeve, önerilen ilk çerçevenin görüntü oluşturmanın yanında faz hatalarını da gidermeyi amaçlayan bir uzantısıdır. Deneysel sonuçlar, önerilen iki çerçevenin görüntü oluşturma ve faz hatası gidermede, düşünülen betiklerin çoğunluğunda en iyi başarımları elde ettiğini göstermektedir. Önerilen üçüncü

çerçeve ATR problemine odaklanmaktadır ve bu çerçevede ATR görevini görüntü uzayı yerine veri uzayında gerçekleştiren iki ATR yaklaşımı sunulmuştur. Deneysel sonuçlarla, ATR görevinin veri uzayında da başarılı bir şekilde gerçekleştirilebileceği gösterilmiş ve bu çerçevenin daha fazla geliştirilmesiyle, en iyi başarımlara ulaşılabilmesinin mümkün olduğu vurgulanmıştır.

Sonuç olarak bu tezdeki çalışmalar, farklı SAR görüntüleme görevlerinde, derin öğrenme teknikleri kullanılarak sonuçların önemli ölçüde iyileştirilebileceğini göstermektedir.

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to my wife and my son

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List of Abbreviations

ADMM Alternating Direction Method of Multipliers . xi, 4, 5, 33, 34, 35, 36, 37, 38, 39, 41, 43, 44, 51, 71, 72, 73
AE Autoencoder
AI Artificial Intelligence
AMP Approximate Message Passing
ATR Automatic Target Recognition iii, iv, v, vi, x, xi, 1, 5, 6, 7, 27, 28, 33, 84, 85, 89, 91, 92
BM3D Block-Matching and 3D Filtering
CNN Convolutional Neural Network. xi, xviii, xix, 4, 5, 28, 29, 30, 32, 36, 39, 42, 43, 44, 47, 52, 73, 84, 85, 91, 92, 94
CT Computed Tomography
DL Dictionary Learning xix, xxi, 33, 39, 43, 44, 52, 55, 62, 67, 68, 91
DNN Deep Neural Network
FBP Filtered Backprojection
FFT Fast Fourier Transform
FIR Finite Impulse Response
GMM Gaussian Mixture Model
GMTI Ground Moving Target Indication
GPU Graphics Processing Unit
GRU Gated Recurrent Unit
HQS Half Quadratic Splitting

IIR Infinite Impulse Response
ILSVRC ImageNet Large Scale Visual Recognition Challange
IMU Inertial Measurement Unit
$\bf IPRPHC$ Image-Phase Removed Phase History Classification xi, xi i, xiv, xxii, 5, 85, 86, 87, 88, 90, 92
ISTA Iterative Shrinkage-Thresholding Algorithm
JPL Jet Propulsion Laboratory
LSTM Long Short-Term Memory
MCA Multi-Channel Autofocus
MLP Multilayer Perceptron xi, 29, 30
$\bf MSTAR$ Moving and Stationary Target Acquisition and Recognition xiv, xviii, 27, 28, 85, 87, 89, 90
NASA National Aeronautics and Space Administration 2, 3
NCSR Nonlocally Constrained Sparse Representation
NLM Non-Local Means
\mathbf{NQR} Non-Quadratic Regularization
PFA Polar Format Algorithm
PGA Phase Gradient Autofocus
PHDC Phase History Domain Classification xi, xiv, xxii, 5, 84, 85, 86, 87, 88, 90, 92, 94
$ \mathbf{PnP} \ \text{Plug-and-Play} \ldots \ \text{iii, v, xi, 4, 5, 7, 33, 34, 35, 36, 39, 41, 43, 44, 55, 69, 91} $
ReLU Rectified Linear Unit
RNN Recurrent Neural Network xi, xviii, 29, 31, 32
ROI Region-of-interest
SAR Synthetic Aperture Radar iii, iv, v, vi, x, xi, xviii, 1, 2, 3, 4, 5, 6, 7, 8, 9, 11, 13, 14, 15, 16, 18, 19, 20, 21, 22, 24, 27, 28, 31, 33, 36, 37, 39, 40, 41, 47, 69, 70, 73, 84, 91, 92
SDA Sparsity-driven Autofocus x, xiv, 26, 27, 73, 75, 83, 92

SNR Signal-to-Noise Ratio xiv, xviii, xi	ix, xx, 2	26, 43, 44	45, 46,	48, 51,	54, 55, 5	58,
61, 75, 83, 91, 92, 93						
SSIM Structural Similarity xiv, xviii, x 75, 83, 91, 92, 93	xix, xx, 4	13, 44, 45	5, 46, 48,	51, 54,	55, 58, 6	31,
UWB Ultra-Wide Band						22

Chapter 1

Introduction

This dissertation presents new approaches to the problems of synthetic aperture radar (SAR) image reconstruction, phase error correction, and automatic target recognition (ATR). The purpose of this chapter is to: 1) introduce the problems addressed in this dissertation; 2) summarize the current state of SAR technology; 3) provide a concise description of the approach taken in this work by pointing out the main contributions; 4) present the outline of the dissertation.

1.1 Synthetic Aperture Radar (SAR) Imaging

Synthetic aperture radar (SAR) imaging is a remote sensing imaging modality that has been widely used in a variety of applications since 1960s. Its advantages over other remote sensing modalities, such as optical imaging, include; day and night capability, all-weather operation, ability of discriminating among different materials, penetrating through different covers and vegetation layers. Specifically, SAR is an active sensor which has its own illumination, i.e., it transmits a chirp signal and collects the returned signal from the area of interest, and hence can work at day or night, whereas optical imaging modalities require an external illumination, e.g., sunlight, which limits the operational time. Moreover, SAR works in the microwave regime, which allows its transmitted signals to penetrate through cloud or rain, as well as light foliage, hence it can work in any weather conditions and detect underground structures if they are buried shallowly in dry environments [4].

SAR sensors are carried on a platform (aircraft or satellite) which travels along a path transmitting microwave pulses towards the ground. Figure 1.1 shows an illustration of SAR data collection using an airborne platform. Some of the transmitted microwave energy is reflected from the ground back to the SAR sensor and received

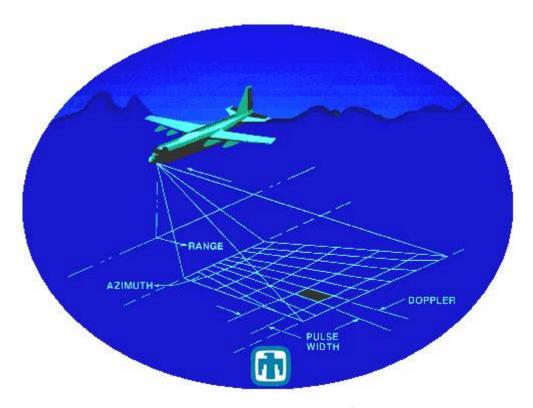


Figure 1.1 Simple illustration of SAR data collection. (Image obtained from the web site of Sandia National Laboratories.)

as a signal. This process is repeated for many aperture positions as the platform moves. The data used for imaging are obtained after a preprocessing of the received signal, involving mixing and filtering steps [5]. The SAR image formation problem is the problem of reconstruction of a spatial reflectivity field of the area of interest from the preprocessed SAR signals.

1.2 Current State of SAR Technology

The use of radar as an imaging sensor dates back as early as the 1940s. However, these early imaging radars, which are nowadays usually referred to as real aperture radars, had an important limitation: the poor resolution achievable with the operating wavelength, as the resolution is inversely proportional to the antenna or aperture size of the sensor. To overcome this limitation, the idea of synthesizing a very long antenna by moving a small one along a convenient path, which is generally attributed to Carl Wiley of Goodyear Aircraft Corporation, is developed in the 1950s [6]. Subsequently, a group of researchers at the University of Illinois carried out the experimental validation of the SAR concept [7]. The first operational airborne SAR system is considered to be the one developed at the University of Michigan, in 1957, which operated at the X-band. Over the following decades, many more SAR systems were built. JPL developed an L-band SAR sensor for NASA, which is



Figure 1.2 A TerraSAR-X image.

installed on a rocket in 1962, and on an airplane in 1966. NASA also upgraded the system developed at the University of Michigan, by adding L-band channel in 1973. SAR sensors have continued to be used in many more NASA missions, including the Apollo 17 lunar mission, SEASAT satellite, spaceborne imaging radar (SIR), and Magellan mission to image the planet Venus [8]. SEASAT is considered to be the first civilian application of SAR, which was oriented to oceanographic investigations. Although its operational lifetime was only a few months due to a short circuit in its power system, obtained results clearly demonstrated how impactful SAR systems were going to be for the area of remote sensing [6]. 1990s and onwards, many satellites with SAR sensors have been built: ALMAZ-1 (1991) from the Soviet Union (and later Russia); ERS-1 (1991), ERS-2 (1995), ENVISAT (2002), Sentinel-1A (2014) and Sentinel-1B (2016) from the European Space Agency (ESA); JERS-1 (1992), ALOS-1 (2006) and ALOS-2 (2014) from the Japan Aerospace Exploration Agency (JAXA); RADARSAT-1 (1995) and RADARSAT-2 (2007) from the Canadian Space Agency (CSA); TerraSAR-X (2007) and TanDEM-X (2010) from the German Space Agency (DLR). Figure 1.2 shows an image from the TerraSAR-X satellite.

Early developments in SAR research were mostly targeting the hardware-related issues and limitations, and considerable improvements were achieved. Back then, this research field was dominated by those working in physics, electromagnetics, and radar engineering. Only in the 1980s, a signal processing view of SAR has emerged [9–11]. Over the years, studies are geared towards imperfections of the collected data, e.g., noisy and/or incomplete data due to the limited observation time, or the environment not being very structured and cooperating. As a result,

the collected data would be a reduced representation of the underlying scene [12].

For many years (and still today, to a large extent), the standard approach for SAR image formation has been through a Fourier transform-based algorithm. This technique, however, is not designed with the limitations such as noisy or incomplete data in mind, and hence cannot overcome such limitations. Furthermore, within this approach, the knowledge about the underlying scene cannot be exploited, and the imaging cannot be reoriented to address the final objective better. The overall output quality is limited by the quality of the collected data.

Improvements in SAR imaging systems draw attention from many different research fields and SAR has started to find use in many applications including reconnaissance [13–15], change detection [16–18], oceanography [19–22], glaciology [23, 24], forestry [25, 26], earthquake monitoring [27–29] and ground moving target indication (GMTI) [15, 30–32]. Most of these applications, if not all, can benefit from automated processing techniques in extracting information from a SAR image for an accurate and efficient interpretation of the scene. Therefore, new processing techniques which are geared towards the final objectives of the mission, and which are robust to reduced data domains are required for the SAR systems [12].

1.3 Contributions of this Dissertation

The first contribution of this dissertation is the formulation of a SAR image reconstruction algorithm using the Plug-and-Play (PnP) priors framework [33], namely, PnP-CNN-SAR. This is the first SAR image formation approach that combines a physics-based forward model with a learning-based prior in a principled and computationally feasible way within the PnP framework. In this framework, the image formation problem is solved through minimization of a regularization-based objective function. In particular, the minimization problem is divided into sub-problems using the Alternating Direction Method of Multipliers (ADMM) algorithm [34], resulting in separating the data-fidelity term and the regularization term of the objective function, to solve sub-problems iteratively. PnP priors framework [33] has shown that, after this separation, optimization of the regularization term, i.e., the prior term, is equivalent to a simple denoising of an intermediate solution under white Gaussian noise. Hence, the regularizer can be replaced with an off-the-shelf denoiser. In our framework, we have replaced the regularizer with a convolutional neural network (CNN) denoiser. We have demonstrated the effectiveness of our framework, i.e., PnP-CNN-SAR, in a variety of scenarios, i.e., for different levels of data limitations and noise for both synthetic and real SAR scenes.

The second contribution of this dissertation is the extension of our first contribution to jointly address phase errors along with image reconstruction, namely, PnP-CNN-SAR-AF. This framework also utilizes the PnP priors using the ADMM algorithm and deep priors, however there are two additional steps in each ADMM iteration, namely, phase error estimation and phase error correction. We have demonstrated the effectiveness of PnP-CNN-SAR-AF in a variety of scenarios, i.e., for different phase error levels for real SAR scenes, improving upon the performance of existing methods.

The third contribution of this dissertation is a new SAR ATR framework that works in the phase history domain. Within this framework, we have introduced two methodologies, namely, phase history domain classification (PHDC) and image-phase removed phase history classification (IPRPHC). PHDC purely works in the phase history domain, i.e., the classification task is performed without image reconstruction. IPRPHC, on the other hand, involves certain steps, particularly, image reconstruction, image-phase removal, and phase history generation from the phase-removed image, before the classification task is performed. Two well-known CNNs, i.e., AlexNet [35] and VGG16 [36], are used for the classification task for both methodologies. IPRPHC has reached a performance level similar to that of state-of-the-art SAR ATR methods while the performance of the PHDC was not competitive enough. Nevertheless, our results suggests that SAR ATR in the phase history domain can be an important research direction.

Overall, this dissertation proposes several ways to utilize deep neural networks (DNNs), particularly CNNs, in order to improve the performance of SAR image reconstruction and automatic target recognition. The ideas proposed here can be extended to build fully-automated end-to-end SAR image reconstruction and ATR frameworks utilizing CNNs.

1.4 Organization

This dissertation is organized as follows. Chapter 2 includes the review of the background knowledge that the ideas proposed here are built upon, i.e., Section 2.1 presents a review of SAR imaging and describes the problems to be addressed in detail, the deep learning background is provided in Section 2.2 along with a brief survey of its utilization in the SAR imaging literature, and Section 2.3 is dedicated to the PnP priors framework. PnP-CNN-SAR is introduced in Chapter 3 and experimental results on both synthetic and real SAR scenes are provided. In Chapter 4, the framework for joint image reconstruction and phase error correction, i.e., PnP-CNN-SAR-AF, is described and its experimental results for real SAR scenes are

presented. Two methodologies proposed for SAR ATR in the phase history domain and their experimental results are presented in Chapter 5. Chapter 6 includes the summary of the contributions and obtained results, as well as potential future work suggestions.

Chapter 2

Preliminaries

In this chapter, we provide background information on three topics, namely, SAR imaging, deep learning, and PnP priors. The discussion on SAR imaging includes the basic principles, mathematical formulation of the data collection and preprocessing in SAR, the autofocus problem and phase errors in SAR imaging, and a review of SAR ATR. The deep learning discussion gives the required background knowledge, and then the use of DNNs for the various SAR tasks is presented. The final part of this chapter is dedicated to the PnP priors framework, which is utilized in our first two contributions.

2.1 SAR Background

Radars use a basic principle known as echo ranging. Echo ranging is the concept that is used to measure the distance of an object by transmitting an echo signal to the object and *listening* the reflection of the echo signal. As the echo signal's reflection will travel twice the distance of the object, one can easily estimate the distance if the propagation speed of the echo signal is known. The distance then will be half of the round trip flight time of the echo signal multiplied by its propagation speed. This principle is implemented in radar systems by transmitting high-bandwidth pulses and using pulse compression techniques [5,37]. Hence, radars can distinguish objects that are at different distances. Imaging radars, however, have to be able to distinguish objects that are at the same distance but at different directions as well, in order to produce a 2D visualization of the scene, i.e., having resolving power only in the range direction is not sufficient. In theory, this can be achieved by transmitting a narrow beam to illuminate a narrow strip of the ground. However, this is not as easy as it sounds, since how narrow the transmitted beam can be, or analogously how small the cross-range resolution can be depends on the antenna aperture size

through the following equation:

$$\rho = \frac{\lambda_w R}{w} \tag{2.1}$$

where λ_w is the wavelength of the illuminating source, R is the target range, and w is the width of the antenna aperture. Consider the case where the wavelength of the source and the distance of the target are 0.03 m and 50 km, respectively. This is a typical wavelength for an X-band radar. Suppose that we want to reconstruct a radar image for which the resolution is 1 meter. In order to achieve this resolution level, we would need a physical antenna with the width of 1500 m, which is of course impractical to carry on an aircraft or a satellite.

SAR solves this problem by sending multiple pulses from a number of observation points, and then focusing the received information coherently to obtain a high-resolution 2-D description of the scene. Hence it synthesizes the effect of a large antenna, using multiple observations from a small antenna [12].

SAR imaging systems can operate in two modes, namely, stripmap-mode SAR, and spotlight-mode SAR. In stripmap-mode SAR, the antenna remains fixed with respect to the radar platform so that the antenna beam sweeps out a strip on the ground. In spotlight-mode, the antenna is steered to continuously illuminate a single spot of terrain. This study focuses on the spotlight-mode SAR. Spotlight-mode SAR is able to provide higher resolution at the expense of spatial coverage, as by steering the antenna, the same terrain portion can be observed through a wider range of angles while other areas within a given accessibility swath of the SAR cannot be illuminated. Figure 2.1 shows the geometry for data collection in a spotlight-mode SAR.

2.1.1 Spotlight-Mode SAR Observation Model

This section provides the preprocessing of the received signals in the spotlight-mode SAR, based on the tomographic derivation in [10], and very closely follows the development in [12]. Let f(x,y) be the complex reflectivity density of the ground patch, and assume that it is constant over the range of frequencies and the range of viewing angles θ employed by the radar. Note that this is an approximation, and there are cases where the dependence of the reflectivity on frequency or aspect angle is important and must be taken into account [38].

In most SAR applications, the transmitted signal is a linear FM chirp signal, which

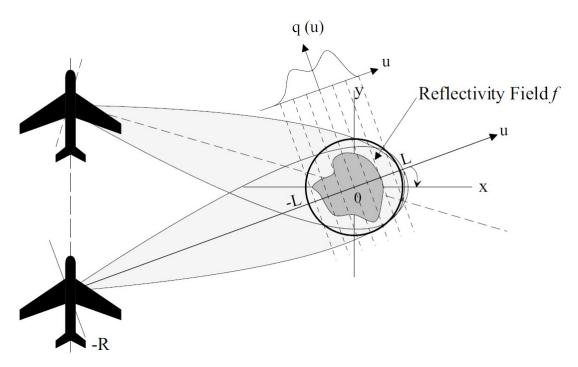


Figure 2.1 Spotlight-mode SAR imaging geometry [1].

has the following form:

$$s(t) = \begin{cases} e^{j(\omega_0 t + \alpha t^2)}, & |t| \le \frac{T_p}{2} \\ 0, & \text{otherwise} \end{cases}$$
 (2.2)

where ω_0 is the carrier frequency and 2α is the chirp rate. Assume the radar transmits the real part of such a signal, $\Re\{s(t)\}$. The return signal $\nu_{(\theta,x_0,y_0)}(t)$ from a differential area centered on the point (x_0,y_0) at a distance R_0 will be:

$$\nu_{(\theta,x_0,y_0)}(t) = |f(x_0,y_0)|\cos\left(\omega_0\left(t - \frac{2R_0}{c}\right) + \alpha\left(t - \frac{2R_0}{c}\right)^2 + \angle f(x_0,y_0)\right)dx dy$$
(2.3)

where c is the speed of the light, and $2R_0/c$ accounts for the two-way travel time from radar to target. Here, the effect of propagation attenuation is neglected, as it can be compensated for later. The complex-valued nature of f(x,y) captures both amplitude scaling and phase shifting of the transmitted waveform by the scatterers. The amplitude scaling is due to that only a fraction of the radiated energy is reflected back to the receiver. The phase shift of the reflected signal, however, could be the result of various factors, the most prominent being the shift at the air/target interface due to the difference between the dielectric constants of air and the target material. The phase shift is also due to the tendency of the RF radiation to creep around target surfaces and its ability to penetrate soft objects and be reflected from within [10]. For most SAR scenes the phase of the reflectivity at a certain location

can be modeled to be random, with a uniform probability density, and uncorrelated with the phase at other locations [11].

The return signal can be more simply written as

$$\nu_{(\theta,x_0,y_0)}(t) = \Re\left\{f(x_0,y_0)s\left(t - \frac{2R_0}{c}\right)\right\}dx \ dy. \tag{2.4}$$

Now let us consider the return from a continuum of scatterers which are at the same distance to the radar. The return from such scatterers will be received by the radar at the same time. Let R be the distance from the radar to the center of the scene, and L be the radius of the ground region of interest, as shown in Figure 2.1. Points in the ground patch equidistant from the radar lie on an arc, but for a typical system $R \gg L$, so that this arc is nearly a straight line. This inequality is related to two conditions that must be satisfied, so that we can assume points at the same range lie on a line (i.e. so that curvature of the wavefront can be neglected). First, the range error due to this assumption for any point in the ground patch must be less than a resolution cell:

$$\frac{L^2}{2R} < \frac{c}{2B} = \rho_x,\tag{2.5}$$

where ρ_x is the range resolution, and B is the bandwidth of the transmitted waveform. Second, the range error due to this assumption at a particular point must not vary much through the aperture:

$$\frac{L^2 sin(2\theta_{max})}{R} \ll \frac{c}{2\omega_0},\tag{2.6}$$

where θ_{max} is the maximum look angle. The derivations of these conditions can be found in [10]. We will assume that the combined return from such an "equidistant" set of scatterers is the sum of the returns that would be received from each individual scatterer. This is a common and reasonable assumption, as discussed in [39]. Let us take $q_{\theta}(u)$ to be such a sum of reflectivities (i.e. a line integral) at distance R+u to the radar, from observation angle θ . Then, we can write the relationship between the projection $q_{\theta}(u)$ and f(x,y) as [40]:

$$q_{\theta}(u) = \iint\limits_{x^2 + u^2 < L^2} \delta(u - x\cos\theta - y\sin\theta) f(x, y) dx \ dy \tag{2.7}$$

This is the standard Radon transform. With this definition, the return signal from a differential line of scatterers normal to the u axis at $u = u_0$ is given by

$$\nu_{\theta,u_0}(t) = \Re\left\{q_{\theta}(u_0)s\left(t - \frac{2(R + u_0)}{c}\right)\right\}du.$$
 (2.8)

This is the contribution to the received signal of all scatterers at range $R+u_0$. Then the return from the entire ground patch (which is what the sensor actually receives) at observation angle θ is given by the integral of ν_{θ,u_0} over u

$$\nu_{\theta}(t) = \Re \left\{ \int_{-L}^{L} q_{\theta}(u) s \left(t - \frac{2(R+u_0)}{c} \right) du \right\}. \tag{2.9}$$

Taking into account that $s(\cdot)$ is a chirp pulse, we have

$$\nu_{\theta}(t) = \Re\left\{ \int_{-L}^{L} q_{\theta}(u) \exp\left\{ j \left[\omega_{0} \left(t - \frac{2(R+u)}{c} \right) + \alpha \left(t - \frac{2(R+u)}{c} \right)^{2} \right] \right\} du \right\}$$
 (2.10)

on the interval

$$-\frac{T_p}{2} + \frac{2(R+L)}{c} \le t \le \frac{T_p}{2} + \frac{2(R-L)}{c}.$$
 (2.11)

Letting $\tau_0 = 2R/c$ be the round-trip delay to the center of the ground patch and mixing with $\nu_{\theta}(t)$ with the reference chirp

$$\exp\left[-j(\omega_0(t-\tau_0) + \alpha(t-\tau_0)^2)\right]$$
 (2.12)

and then low-pass filtering yields the complex signal¹

$$\bar{r}_{\theta}(t) = \int_{-L}^{L} q_{\theta}(u) \exp\left\{j\frac{4\alpha u^2}{c^2}\right\} \cdot \exp\left\{-j\frac{2}{c}(\omega_0 + 2\alpha(t - \tau_0))u\right\} du. \tag{2.13}$$

So $\bar{r}_{\theta}(t)$ is the demodulated observation signal at platform position θ , as a function of time. In practice, the mixing operation described above is carried out by multiplying $\nu_{\theta}(t)$ with the in phase and quadrature components of the reference chirp signal, i.e.,

$$s_I(t) = \cos(\omega(t - \tau_0) + \alpha(t - \tau_0)^2)$$
 (2.14)

$$s_O(t) = -\sin(\omega(t - \tau_0) + \alpha(t - \tau_0)^2),$$
 (2.15)

separately. Also note, we assume here that τ_0 is known. In practice it is only known imperfectly and this makes it necessary to have a post-processing technique in SARs known as autofocus or automatic phase-error correction [5]. The autofocus problem will be discussed in Section 2.1.4.

We will assume that the effect of the quadratic phase term $\exp\{j4\alpha u^2/c^2\}$ in (2.13) can be neglected. This is a reasonable approximation for most situations, since usually it is true that $4\alpha u^2 \ll c^2$. A more detailed analysis of this approximation

¹Note that a constant factor of 1/2 is neglected here.

can be found in [5]. After this approximation, the observed signal is given by:

$$r_{\theta}(t) = \int_{-L}^{L} q_{\theta}(u) \exp\left\{-j\frac{2}{c}(\omega_0 + 2\alpha(t - \tau_0))u\right\} du$$
$$= \int_{-L}^{L} \exp\left\{-j\Omega(t)u\right\} du. \tag{2.16}$$

This signal can be identified as the Fourier transform of the projection $q_{\theta}(u)$ where the spatial frequency variable is $\Omega(t) = \frac{2}{c}(\omega_0 + 2\alpha(t - \tau_0))$. Note $\Omega(t)$ is limited to a finite spatial frequency interval, because the observation duration t is limited, and the chirp rate α is finite (equivalently s(t) is narrow-band). Also $\Omega(t)$ is offset from the origin of the spatial frequency plane due to ω_0 . In summary, at least within the time interval considered, the processed return signal $r_{\theta}(t)$ carries band-pass information related to a particular line integral of the reflectivity field.

To derive the relationship between the field f(x,y) and the demodulated observed signal $r_{\theta}(t)$, let us substitute (2.7) in the observation relationship (2.16), to obtain

$$r_{\theta}(t) = \int\limits_{|u| \le L} \iint\limits_{x^2 + y^2 \le L^2} \delta(u - x\cos\theta - y\sin\theta) f(x, y) \exp\{-j\Omega(t)\} dx \ dy \ du$$

$$= \iint\limits_{x^2 + y^2 \le L^2} f(x, y) \exp\{-j\Omega(t)(x\cos\theta + y\sin\theta)\} dx \ dy$$
(2.17)

Hence, $r_{\theta}(t)$ is a finite (i.e. band-limited) slice at angle θ from the 2D Fourier transform of the filed f(x,y). Here $\Omega(t)$ serves as the radial spatial frequency. So, there are two interpretations of $r_{\theta}(t)$: a 1D Fourier transform of the projections (based on (2.16)), and a slice through the 2D Fourier transform of the field (based on (2.17)). This equivalence is essentially a band-limited version of the projection slice theorem [40] from computed tomography (CT). The data $r_{\theta}(t)$ from all observation angles are usually called the phase histories and lie on a polar grid in the 2D frequency domain as shown in Figure 2.2.

Now let us derive the discrete observation model. (2.17) can be compactly written as $r_{\theta}(t) = (\mathcal{H}_{\theta}f(x,y))(t)$, where \mathcal{H}_{θ} is the continuous observation kernel. In practice, the observations at the *i*-th observation angle θ_i are samples $r_{\theta_i}(t_j)$ of the continuous received signal $r_{\theta_i}(t)$ at sampling times t_j . This sampling in the time domain results in sampling of the spectrum of the underlying reflectivity field. Sampling places a limit on the maximum allowable scene size that can be imaged without aliasing (in the spatial domain) [12].

Let \mathbf{r}_{θ_i} be the vector of these observed samples, \mathbf{H}_{θ_i} be a discretized approximation to the continuous observation kernel \mathcal{H}_{θ_i} and \mathbf{f} be a complex-valued vector representing

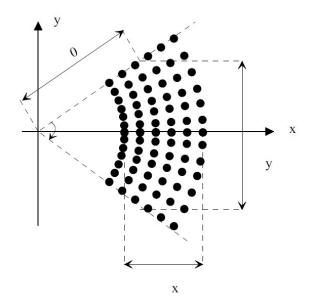


Figure 2.2 Graphical representation of an annulus segment containing known samples of the Fourier transform of the reflectivity density.

the unknown sampled reflectivity image. Then, overall, we can write:

$$\underbrace{\begin{bmatrix} \mathbf{r}_{\theta_1} \\ \mathbf{r}_{\theta_2} \\ \vdots \\ \mathbf{r}_{\theta_M} \end{bmatrix}}_{\mathbf{r}} = \underbrace{\begin{bmatrix} \mathbf{H}_{\theta_1} \\ \mathbf{H}_{\theta_2} \\ \vdots \\ \mathbf{H}_{\theta_M} \end{bmatrix}}_{\mathbf{H}} \mathbf{f} \tag{2.18}$$

where M is the total number of angular observation points. The data $\mathbf{r} \in \mathbb{C}^{KM \times 1}$ are the column-stacked sampled phase histories where K is the number of range positions. If we consider the presence of measurement noise, the observation model becomes

$$\mathbf{y} = \mathbf{H}\mathbf{f} + \mathbf{n} \tag{2.19}$$

where \mathbf{n} stands for measurement noise and \mathbf{y} is the noisy observations.

2.1.2 Range and Azimuth Resolution

Now, let us try to find the relationship between the achievable resolution of the SAR image and the dimensions of the annular region shown in Figure 2.2, which is specified by system parameters. We can motivate a definition of resolution in the image domain with the assumption that this annulus can be approximated by a rectangle of width $\Delta\Omega_x$ and height $\Delta\Omega_y$. Let us consider a point reflector in the scene. If we compute the Fourier transform of the scene limited to this rectangular region, and then compute an inverse Fourier transform, we would obtain a two

dimensional sinc function. The resolution of the formed image gets better as the mainlobe of this sinc gets narrower, which can be achieved by a wider the support of the rectangular region. More precisely, the first zero crossings of the sinc occur at $2\pi/\Delta\Omega_x$ and $2\pi/\Delta\Omega_y$. Hence, the resolution of two point reflectors having equal reflectivity requires that the reflectors be separated by more than $\rho_x = 2\pi/\Delta\Omega_x$ in the x (range) dimension and $\rho_y = 2\pi/\Delta\Omega_y$ in the y (cross-range or azimuth) dimension.

Let us first consider the range resolution. Assume that the width of the rectangle $2\pi/\Delta\Omega_x$ is equal to the radial width of the annular region, which is essentially the spatial frequency bandwidth of each return. We can determine the lower and upper limits of the radial extent by substituting the limits for the observation time t from (2.11) into the definition of $\Omega(t)$ to find the lowest and highest spatial frequencies:

$$\Omega_{x_l} = \frac{2}{c} \left(\omega_0 - \alpha T_p + \frac{4\alpha L}{c} \right)
\Omega_{x_h} = \frac{2}{c} \left(\omega_0 + \alpha T_p - \frac{4\alpha L}{c} \right).$$
(2.20)

For a typical SAR, we can assume $T_p \gg 4L/c$. Hence, we can conclude that

$$\Delta\Omega_x = \Omega_{x_h} - \Omega_{x_l} \approx \frac{4\alpha T_p}{c} = \frac{4\pi B}{c}$$
 (2.21)

where we have used the fact that the bandwidth of the transmitted pulse (in Hz) is given by $B = \alpha T_p/\pi$.

Now let us consider the cross-range (azimuth) resolution, which will be determined by $\Delta\Omega_y$. From Figure 2.2:

$$\sin\left(\frac{\Delta\theta}{2}\right) \approx \frac{\Delta\Omega_y/2}{\Omega_0} \tag{2.22}$$

where, in our case $\Omega_0 = 2\omega_0/c$. Hence, for $\Delta\theta \ll 1$ we have

$$\Delta\Omega_y \approx \frac{2\omega_0 \Delta\theta}{c}.\tag{2.23}$$

Lastly, since the wavelength of the transmitted pulse is given by $\lambda = 2\pi c/\omega_0$, we can deduce the following range and cross-range resolution relationships for the system:

$$\rho_x \approx \frac{c}{2B} \tag{2.24}$$

$$\rho_y \approx \frac{\pi c}{\omega_0 \Delta \theta} = \frac{\lambda}{2\Delta \theta}.$$
 (2.25)

In conclusion, the range resolution ρ_x depends on the bandwidth of the transmitted pulse, while the cross-range resolution is determined by angular diversity of the observations and carrier frequency (and therefore equivalently wavelength) of the transmitted chirp signal.

2.1.3 SAR Image Reconstruction Methods

The problem of SAR image reconstruction is to obtain an estimate of the reflectivity density f(x,y) based on the observed, pre-processed SAR data. Since the phase history data constitute a band-limited two dimensional spatial Fourier transform of the reflectivity density, the standard approach to tackle the SAR image reconstruction problem has been based on two-dimensional fast Fourier transform (FFT). These 2D FFT based methods are therefore termed as *conventional* methods. However, these 2D FFT-based reconstruction methods are ideal only when perfect data are available throughout the spatial frequency domain, which is not the case in a practical measurement scenario, primarily due to finite bandwidth of the transmitted signal and the finite range of look angles. Therefore, the problem in (2.19) is ill-posed [41], and can be considered as ill-conditioned to a certain extent depending on the specifics of the observation scenario. Consequently, (2.19) can only be solved satisfactorily by incorporating some sort of regularization into the inversion process. In this section, we will describe two well-known conventional methods, namely, polar format algorithm (PFA) and filtered backprojection (FBP), as well as regularization-based image reconstruction methods.

2.1.3.1 Conventional Reconstruction Methods

Conventional methods are based on the inverse operator for the case when perfect data are available throughout the spatial frequency domain. These methods have no explicit mechanism to counter any imperfection in the data. Although there are algorithmic differences between the two methods, the reconstructions they produce are similar.

Polar Format Algorithm In polar format algorithm (PFA), the known data samples are first interpolated to a Cartesian grid, assuming unknown samples to be zero. Then, an inverse 2-D FFT is employed and the magnitude of the reconstructed complex image is displayed for viewing. To reduce sidelobe levels, the data can be windowed before FFT processing. This is the most common SAR image reconstruction algorithm.

Filtered Backprojection Filtered backprojection algorithm (FBP) [10,40,42] is suggested by the tomographic formulation of SAR [10]. The derivation of FBP is through the 2D inverse Fourier transform in polar coordinates. The radial slices in the frequency domain are then recognized as 1-D Fourier transforms of the projections of the field at the corresponding angle, by virtue of the projection slice theorem [40]. This way the double Fourier integral is reduced to two sequential operations: first the data at each observation angle are filtered by a ramp filter, and then the results are backprojected to obtain a reconstruction. FBP is the algorithm that is currently used in commercial CT scanners.

2.1.3.2 Regularization-Based Reconstruction Methods

In image reconstruction and restoration problems, the goal is to find an estimate of a 2D field from its indirect observations. Hence, we can view image reconstruction and restoration problems as general observation problems which we meet in most situations of engineering interest. We will concentrate on problems where the mathematical relationship between the measurements and the field is governed by a linear integral equation, i.e., as in (2.19). This simple observation model that we derived for SAR imaging is analogous to those of many other engineering problems. At first look, it might seem easy to find an estimate $\hat{\mathbf{f}}$ of \mathbf{f} with a simple matrix inversion, however that certainly is not possible in general, e.g., the matrix may not even be square. There are four main issues that this approach can not handle [43]:

- 1. Due to the observation noise, there may not exist any **f** which solves this equation exactly.
- 2. There may be more than one \mathbf{f} which satisfy these equations, hence the solution may not be unique.
- 3. The estimate $\hat{\mathbf{f}}$ is desired to remain stable despite the perturbations in the observations.
- 4. It is desirable to include any a priori information about \mathbf{f} in the inversion process, however this approach is only data-driven.

Least-squares solution can overcome the first problem mentioned above. The solution is the best fit to the observed data in the least-squares sense:

$$\hat{\mathbf{f}}_{\mathsf{LS}} = \underset{\mathbf{f}}{\operatorname{argmin}} \|\mathbf{y} - \mathbf{H}\mathbf{f}\|_{2}^{2} \tag{2.26}$$

where $\|\cdot\|_2$ denotes the ℓ_2 norm. If **H** has full column rank, the estimate is unique

and is obtained as

$$(\mathbf{H}^H \mathbf{H})\hat{\mathbf{f}}_{LS} = \mathbf{H}^H \mathbf{y}. \tag{2.27}$$

However, when the null-space of \mathbf{H} is not empty, the least-squares solution is not unique. A common approach to address this problem is to choose the field with minimum norm, among the set of least-squares solutions, as the estimate of the true field \mathbf{f} :

$$\hat{\mathbf{f}}^{+} = \underset{\mathbf{f}_{LS}}{\operatorname{argmin}} \left\| \hat{\mathbf{f}}_{LS} \right\|_{2} \tag{2.28}$$

which is called as the generalized solution. Although generalized solution provides, a simple, reasonable way to deal with the first two issues mentioned in the list above, it does not directly address the third and fourth issues. Particularly, if the model matrix **H** is ill-conditioned, i.e., the ratio of the largest eigenvalue to the smallest is very large, small changes in the data lead to large changes in the solution.

These issues, which cannot be resolved by the generalized solution, can be addressed using regularization. Regularization allows us to include any prior information to stabilize the solution in the presence of noisy data, and allow reasonable estimates.

Tikhonov Regularization One of the most common regularization methods is Tikhonov regularization [44, 45]. In this method, the prior information on the field is incorporated by including an additional term to the least-squares cost function:

$$\hat{\mathbf{f}}_{\mathsf{Tik}} = \underset{\mathbf{f}}{\operatorname{argmin}} \|\mathbf{y} - \mathbf{H}\mathbf{f}\|_{2}^{2} + \lambda^{2} \|\mathbf{P}\mathbf{f}\|_{2}^{2}$$
(2.29)

where **P** is a matrix, and λ is a scalar. The first term in (2.29) is a data-fidelity term. The second term, through which the prior information about **f** is incorporated, is called the regularizer. The parameter λ determines the weight of the prior knowledge in the estimation process.

The choice of the matrix \mathbf{P} is determined by which information about the field we want to incorporate and the simplest choice would be an identity matrix. In this case, large values in the reconstruction would be penalized by the regularizer. If \mathbf{P} is chosen as 2D derivative (gradient) operator, then its effect will be to penalize roughness on the solution, which essentially enforces the final reconstruction to be smooth.

By taking the gradient of (2.29) with respect to \mathbf{f} and equate it to zero, we reach the following set of linear equations as the solution for the Tikhonov regularization:

$$(\mathbf{H}^H \mathbf{H} + \lambda^2 \mathbf{P}^H \mathbf{P}) \hat{\mathbf{f}}_{\mathsf{Tik}} = \mathbf{H}^H \mathbf{y}$$
 (2.30)

When the null spaces of \mathbf{H} and \mathbf{P} are distinct, there exists a unique, closed-form solution to (2.30).

Non-Quadratic Regularization The cost function of Tikhonov regularization (2.29) is a quadratic function of **f**. Hence, (2.30) leads to a linear function of **f**, which in turn corresponds to a linear processing of the data **y** for image restoration or reconstruction. Although it is desirable to have such a linear processing as it ensures computational efficiency, it fails to obtain more powerful results that are only possible if nonlinear methods are allowed [43], e.g., in many imaging applications the data are expected to be sparse as themselves, or in a transform domain, and sparsity cannot be effectively enforced within a linear framework. Thus, let us consider more general problems of the following form:

$$\hat{\mathbf{f}}_{NQ} = \underset{\mathbf{f}}{\operatorname{argmin}} \|\mathbf{y} - \mathbf{H}\mathbf{f}\|_{2}^{2} + \lambda^{2} \sum_{i=1}^{M} \psi((\mathbf{P}\mathbf{f})_{i})$$
(2.31)

where M is the length of the vector \mathbf{Pf} , and $(\mathbf{Pf})_i$ denotes its i-th element. Note that when $\psi(x) = x^2$, (2.31) reduces to the Tikhonov cost function in (2.29), however $\psi(x)$ is in general non-quadratic. Many well-known regularization approaches fall into the category of non-quadratic regularization, e.g., maximum entropy [46], and the total variation [47,48] methods. Note that unlike Tikhonov regularization, (2.31) does not lead to a closed-form solution in general, hence numerical methods must be used to find $\hat{\mathbf{f}}_{NQ}$. Half quadratic splitting (HQS) [49,50] is one such method that can be used to solve (2.31).

The sparsity-promoting formulation for the case that the field \mathbf{f} is expected to be sparse would be in the form

$$\hat{\mathbf{f}}_{\mathsf{Sparse}} = \underset{\mathbf{f}}{\operatorname{argmin}} \|\mathbf{y} - \mathbf{H}\mathbf{f}\|_{2}^{2} + \lambda^{2} \|\mathbf{f}\|_{p}^{p}$$
(2.32)

with $p \leq 1$ being a common choice. See Figure 2.3 for an illustration for the case of a scalar, real-valued version of $\|\mathbf{f}\|_p^p$, i.e., $|f|^p$ for various choices of p. When we view these plots as penalty functions, we deduce that as the value of p gets smaller, the relative penalty on large values of f reduces. In the SAR imaging problem, this effect helps the preservation and enhancement of strong scatterers in the scene, while still suppressing artifacts. On the other hand, small p values are more punishing for the smaller values of f, hence enforcing sparsity. However, for the values of p < 1, the penalty functions become concave, thus minimization of objective functions containing such terms can be challenging. Therefore the choice of p = 1 has become quite popular as it can enforce sparsity while also generating a

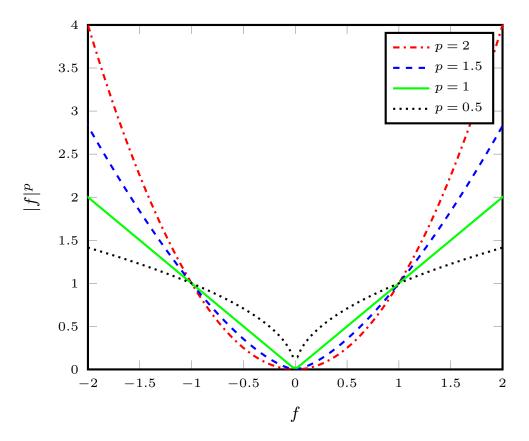


Figure 2.3 Behavior of the function $|f|^p$ for various choices of p.

convex penalty function.

The formulation for the total variation regularization is given as:

$$\hat{\mathbf{f}}_{\mathsf{TV}} = \underset{\mathbf{f}}{\operatorname{argmin}} \|\mathbf{y} - \mathbf{H}\mathbf{f}\|_{2}^{2} + \lambda^{2} \|\nabla \mathbf{f}\|_{1}$$
 (2.33)

where ∇ is the discrete gradient operator. The regularizer in (2.33) preserves strong edges and produces improved reconstruction quality in piece-wise smooth regions.

Feature-Enhanced SAR Imaging A feature-enhanced SAR image reconstruction approach based on non-quadratic regularization was proposed in [1, 12]. The overall SAR image formation in this case is modeled by the following optimization problem:

$$\hat{\mathbf{f}}_{\mathsf{FE}} = \underset{\mathbf{f}}{\operatorname{argmin}} \|\mathbf{y} - \mathbf{H}\mathbf{f}\|_{2}^{2} + \lambda_{1} \|\mathbf{f}\|_{p}^{p} + \lambda_{2} \|\nabla |\mathbf{f}|\|_{p}^{p}$$
(2.34)

where ∇ is the 2D discrete gradient operator. $|\mathbf{f}|$ denotes the vector of magnitudes of the complex-valued vector \mathbf{f} . The scalar parameters λ_1 and λ_2 determine the weights of the corresponding terms in the estimation process. The first term in (2.34) is the data-fidelity term that ensures the estimate $\hat{\mathbf{f}}_{\mathsf{FE}}$ is consistent with the data. The second and third terms in (2.34) incorporate prior information regarding

both the behavior of the field \mathbf{f} , and the nature of the features of interest in the resulting reconstructions. These terms are aimed at enhancing point-based and region-based features respectively. The relative magnitudes of the parameters λ_1 and λ_2 determine the relative emphasis on these two types of features. Note for the region-based feature enhancement that the smoothness is imposed on the magnitude of the reflectivity field, as the correlation in a homogeneous region of \mathbf{f} in SAR is due to the similarity of backscatter power, which is better represented in the magnitude of \mathbf{f} than its real and imaginary parts [12].

2.1.4 SAR Autofocus Problem & Phase Errors

In SAR systems, the demodulation time τ_0 at every aperture position², which is the time required for the signal transmitted by the SAR sensor to propagate from the SAR platform to the field and return, is required to be known precisely, in order to obtain the data used for imaging from the returned signals. The inexact knowledge of the demodulation time causes phase errors in the SAR data which result in defocusing of the reconstructed images [5]. The inexact measurement of the distance between the SAR platform and the scene due to SAR platform position uncertainties, or random delays in the signal due to propagation in atmospheric turbulence are among the most common causes of demodulation time errors. As these errors cause defocusing in the reconstructed image, this problem is known as SAR autofocus problem, and phase error correction methods are also called autofocus methods. In addition to the uncertainties related to the SAR platform, moving targets on the ground can also cause phase errors. However, this kind of phase errors would affect the reconstructed image locally, i.e., only around the moving target, hence these phase errors cause space-variant defocusing.

In current SAR systems, inertial measurement units (IMUs) are used to measure the distance R between the SAR platform and the patch center. However, even with high quality IMUs, the measurement of R might not be within the maximum tolerated error margin. Inexact measurements of R cause demodulation time errors, which in turn results in phase errors in the SAR data obtained after the preprocessing of the returned signal. Conventional approaches to tackle this problem have been to increase the accuracy of IMUs and postprocessing the reconstructed SAR image to remove phase errors. These postprocessing methods are generally termed as autofocusing techniques, and they have advantages over improving IMUs' accuracy. Improving accuracy of the IMU systems can only help reducing the effects of demodulation time errors caused by platform position uncertainties, however it

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²Recall that in Section 2.1.1, the demodulation time was defined as $\tau_0 = 2R/c$ where R is the distance from the SAR platform to the patch center.

would be of no help against phase errors caused by atmospheric turbulence. Autofocusing techniques can remove the effects of demodulation errors independent of the error source, while also eliminating the significant hardware costs associated with ultra-high-accuracy navigation systems [5].

In Section 2.1.1, we have derived the SAR observation model without considering demodulation time errors, i.e., with the assumption that the demodulation time at each aperture position is known exactly. Demodulation time errors can be modeled as constant phase errors on each range compressed pulse. Hence, if we let ϵ denote the demodulation time error, during preprocessing, the received signals are multiplied with

$$s_I(t) = \cos(\omega(t - \tau_0 + \epsilon) + \alpha(t - \tau_0 + \epsilon)^2)$$
(2.35)

$$s_Q(t) = -\sin(\omega(t - \tau_0 + \epsilon) + \alpha(t - \tau_0 + \epsilon)^2), \tag{2.36}$$

instead of the expressions in (2.14) and $(2.15)^3$. Then, the output of the preprocessing step becomes

$$Z_{\epsilon}(\Omega(t)) = r_{\theta_{\epsilon}}(t) = \exp\left\{(-j\epsilon^{2}\alpha)(j\frac{\epsilon c}{2}\Omega(t))\right\} \int_{|u| \le L} q_{\theta}(u) \exp\left\{-j\Omega(t)u\right\} du. \quad (2.37)$$

Thus, the corrupted and error-free phase history expressions are related as

$$Z_{\epsilon}(\Omega(t)) = \exp\{(-j\epsilon^2\alpha)(j\frac{\epsilon c}{2}\Omega(t))\}Z(\Omega(t)). \tag{2.38}$$

 $\epsilon^2 \alpha \ll 1$, and hence we can assume $\exp\{-j\epsilon^2 \alpha\} \approx 1$. Then (2.38) becomes

$$Z_{\epsilon}(\Omega(t)) = \exp\left\{j\frac{\epsilon c}{2}\Omega(t)\right\}Z(\Omega(t)). \tag{2.39}$$

Hence, $Z_{\epsilon}(\Omega(t))$ is simply $Z(\Omega(t))$ altered by a linear phase term. Substituting $\Omega(t) = \frac{2}{c}(\omega_0 + 2\alpha(t - \tau_0))$ into (2.39), we get

$$Z_{\epsilon}(\Omega(t)) = \exp\{(j\epsilon\omega_0)(j\epsilon(2\alpha(t-\tau_0)))\}Z(\Omega(t)). \tag{2.40}$$

The value of the term $2\alpha(t-\tau_0)$ is in general much smaller compared to the value of ω_0 , and hence can be neglected. Thus we would obtain

$$Z_{\epsilon}(\Omega(t)) = \exp\{j\phi\}Z(\Omega(t)) \tag{2.41}$$

where $\phi = \epsilon \omega_0$ is the phase error. Note that ϕ is different at every aperture position,

³Note that the development in this section closely follows the one in [51].

therefore its effect on the reconstructed image is along the cross-range. The implication of such an error in the image domain is the convolution of each range line of the image with a 1D blurring kernel in the cross-range direction. Hence, such phase errors cause defocusing of the image in the cross-range direction [52].

Although most phase errors encountered are 1D cross-range varying functions, it is also possible to encounter both range and cross-range varying 2D phase errors. For example, low frequency UWB SAR systems may suffer from severe propagation effects that can appear through the ionosphere, including Faraday rotation, dispersion, and scintillation. Dispersion imposes an unknown phase error on the transmitted chirp. On a single pulse basis, dispersion would cause the ideal impulse response function to defocus in the range direction. Furthermore, pulse-to-pulse variations in the dispersion and propagation delay lead to defocus in the cross-range direction. Because these unknown phase errors change over the synthetic aperture, the 2D phase history becomes corrupted by a fully coupled 2D phase error [53]. In principle, 2D phase errors can be handled in two sub-categories, namely, separable and non-separable errors, nevertheless 2D separable phase errors are not common in practice.

2.1.4.1 2D Non-separable Phase Errors

In the presence of 2D non-separable phase errors, all sample points of the phase history data are perturbed with different and potentially independent phase errors. Let $\mathbf{R} \in \mathbb{C}^{K \times M}$ denote the sampled phase history data⁴, and Φ_{2D-ns} be a 2D non-separable phase error function. The relationship between the phase-corrupted and error-free phase histories are as follows:

$$\mathbf{R}_{\epsilon}(k,m) = e^{j\mathbf{\Phi}_{2D-ns}(k,m)}\mathbf{R}(k,m)$$
(2.42)

where \mathbf{R}_{ϵ} denotes the phase-corrupted phase history data, and $k \in 1, 2, ..., K$ and $m \in 1, 2, ..., M$ denote range and cross-range sample positions, respectively. To express this relationship in terms of the observation model, first we define the vector ϕ_{2D-ns} as

$$\phi_{2D-ns} = [\phi_{2D-ns}(1), \phi_{2D-ns}(2), \dots, \phi_{2D-ns}(KM)]^T$$
(2.43)

which is created by concatenating the columns of the phase error matrix Φ_{2D-ns} under each other. Using the corresponding vector forms, the relationship in (2.42)

⁴Recall that the column-stacked vector version of sampled phase histories were denoted by $\mathbf{r} \in \mathbb{C}^{KM \times 1}$, in Section 2.1.1.

becomes

$$\mathbf{r}_{\epsilon} = \mathbf{D}_{2D-ns}\mathbf{r} \tag{2.44}$$

where $\mathbf{D}_{2D-ns} \in \mathbb{C}^{KM \times KM}$ is a diagonal matrix in the form:

$$\mathbf{D}_{2D-ns} = \text{diag}\{e^{\phi_{2D-ns}}\} \tag{2.45}$$

In terms of observation model matrices, the relationship in (2.44) is as follows:

$$\mathbf{H}(\phi_{2D-ns})\mathbf{f} = \mathbf{D}_{2D-ns}\mathbf{H}\mathbf{f} \tag{2.46}$$

where **H** is the initially assumed model matrix by the imaging system, and $\mathbf{H}(\phi_{2D-ns})$ is the model matrix that takes the phase errors into account.

2.1.4.2 2D Separable Phase Errors

A 2D separable phase error function is composed of range varying and cross-range varying 1D phase error functions as follows:

$$\mathbf{\Phi}_{2D-s}(k,m) = \boldsymbol{\xi}(k) + \boldsymbol{\gamma}(m). \tag{2.47}$$

Here, $\boldsymbol{\xi} \in \mathbb{C}^{K \times 1}$ represents the range varying phase error, and $\boldsymbol{\gamma} \in \mathbb{C}^{M \times 1}$ represents the cross-range varying phase error. The vector $\boldsymbol{\phi}_{2D-s} \in \mathbb{C}^{KM \times 1}$ for 2D separable phase errors is obtained by concatenating the columns of $\boldsymbol{\Phi}_{2D-s}$ as follows:

$$\phi_{2D-s} = \left[\underbrace{\xi(1) + \gamma(1)}_{\phi_{2D-s}(1)}, \dots, \underbrace{\xi(K) + \gamma(1)}_{\phi_{2D-s}(K)}, \underbrace{\xi(2) + \gamma(1)}_{\phi_{2D-s}(K+1)}, \dots, \underbrace{\xi(K) + \gamma(M)}_{\phi_{2D-s}(K)}, \dots, \underbrace{\xi(K) + \gamma(M)}_{\phi_{2D-s}(KM)}\right]^{T}$$
(2.48)

A 2D separable phase error function affects the observation model matrix in the as follows:

$$\mathbf{H}(\phi_{2D-s})\mathbf{f} = \mathbf{D}_{2D-s}\mathbf{H}\mathbf{f} \tag{2.49}$$

where \mathbf{D}_{2D-s} is given as

$$\phi_{2D-s} = [\phi_{2D-s}(1), \phi_{2D-s}(2), \dots, \phi_{2D-s}(KM)]^T.$$
 (2.50)

2.1.4.3 1D Phase Errors

In the case of 1D phase errors, for a particular cross-range position the phase error is the same at all range positions. Let $\phi_{1D} \in \mathbb{C}^M$ be a vector whose elements are cross-range varying phase errors for every aperture position m = 1, 2, ..., M:

$$\phi_{1D} = [\phi_{1D}(1), \phi_{1D}(2), \dots, \phi_{1D}(M)]^T.$$
 (2.51)

Then, the relationship between the error-free and the phase-corrupted data can be expressed as:

$$\mathbf{H}(\phi_{1D})\mathbf{f} = \mathbf{D}_{1D}\mathbf{H}\mathbf{f} \tag{2.52}$$

where \mathbf{D}_{1D} is given as

$$\mathbf{D}_{1D} = \operatorname{diag} \left\{ \underbrace{\left[\underbrace{e^{j\phi_{1D}(1)}, \dots, e^{j\phi_{1D}(1)}}_{K}, \underbrace{e^{j\phi_{1D}(2)}, \dots, e^{j\phi_{1D}(2)}}_{K}, \dots, \underbrace{e^{j\phi_{1D}(M)}, \dots, e^{j\phi_{1D}(M)}}_{K} \right] \right\}. \tag{2.53}$$

Note that, in the case of 1D phase errors, there are M unknowns, while in 2D separable and non-separable phase error cases, there are M+K and MK unknowns, respectively. Hence, 2D non-separable phase error correction is a more challenging task compared to the others.

2.1.5 Existing Autofocus Methods

SAR autofocus has been an active area of research over the years [52,54–104]. In early studies, researchers have focused on 1D phase errors, and tackling 2D separable and non-separable phase errors is a relatively recent research area [52,75–77,79,96]. Most of the existing autofocusing algorithms perform postprocessing, i.e., they try to eliminate the effects of phase errors from conventionally reconstructed images. Recently, algorithms for joint image reconstruction and phase error correction have also been proposed [52,75,76,79,81].

2.1.5.1 Conventional Methods

Two methods, namely, inverse filtering and subaperture-based methods, are considered as conventional methods, as they are the earliest attempts to tackle phase errors, and have simple formulations with easy implementations.

Inverse Filtering In inverse filtering, the amount of defocus on a single point target is used to estimate phase errors. As mentioned before, the effect of 1D phase errors in the image domain can be viewed as convolution of each range line of the image with a 1D blurring kernel in the cross-range direction. This effect can be expressed as

$$\tilde{\mathbf{F}}(a,b) = \tilde{\mathbf{h}}(b) \circledast \mathbf{F}(a,b) \tag{2.54}$$

where

$$\tilde{\mathbf{h}}(b) = \mathfrak{F}_m^{-1} e^{j\phi(m)}. \tag{2.55}$$

Here, $\tilde{\mathbf{F}}$ denotes the defocused version of the image \mathbf{F}^5 , \mathfrak{F}^{-1} denotes inverse Fourier transform, a and b are range and cross-range image domain indices, respectively, \circledast denotes circular convolution operation and, m is the cross-range index in the frequency domain. In inverse filtering approach, it is assumed that a single point target can be isolated in the defocused image. This technique estimates phase errors by finding such an isolated strong point target in the defocused image and then using the defocus information on that point target.

Subaperture-based Methods In these methods, the data from subapertures are used to estimate phase errors. The main assumption of these methods is that the phase error function is a polynomial function of the aperture position. For example, in case of quadratic phase error, the aperture is divided into two subapertures, then in each subaperture, the phase error is approximated by a linear function. Since a linear phase error function only shifts the image proportional to its slope, the two low-resolution defocused images reconstructed from the two subaperture data are shifted versions of the original image in reverse directions [51].

2.1.5.2 Phase Gradient Autofocus (PGA)

In phase gradient autofocus (PGA) [57,61], the phase error function is estimated by averaging across many range lines, as every target in the image is corrupted by the same blur function. This averaging operation is performed using maximum likelihood estimation formulation. Unlike subaperture-based methods, PGA is a non-parametric algorithm. The algorithm aims to isolate a number of single targets in the image to estimate phase error from those targets. Single targets are isolated via center shifting and windowing operations. PGA selects the strongest target on each range line and circularly shifts it to the scene center, as using the targets with strong reflectivities provides a much better phase error estimation than using the targets with weak reflectivities. This shifting operation produces a new image. In

 $^{{}^{5}\}mathbf{F}$ is the matrix version of \mathbf{f} , i.e., another interpretation is that \mathbf{f} is a column-stacked vector generated from \mathbf{F} .

this new image, all of the targets which will be used in the estimation process, lie in the center of the cross-range dimension. In the next step, a windowing operation is performed to preserve the information contained in the blur footprints of the center-shifted targets and reject information from all other surrounding targets with weak reflectivities. After center-shifting, the necessary information, contained in the support of the blur footprint, is extracted through windowing. In this stage, determining the window width is crucial, as choosing a width smaller than the blur footstep would prohibit capturing all of the necessary information, while choosing the width larger than the blur footprint would increase noise levels.

2.1.5.3 Multi-Channel Autofocus (MCA)

Multi-channel autofocus MCA [74] employs a non-iterative algorithm which finds the focused image using a basis formed from the defocused image, relying on a condition on the image support to obtain a unique solution. In particular, MCA estimates 1D phase error functions by directly solving a set of linear equations obtained through an assumption that there are zero-reflectivity regions in the scene to be imaged. When this is not precisely satisfied, presence of a low-return region is exploited, and the phase error is estimated by minimizing the energy of the low-return region. When the desired conditions are satisfied, MCA performs very well. However, in scenarios involving low-quality data, e.g., due to low SNR, the performance of MCA degrades. A number of modifications to MCA have been proposed, including the incorporation of sharpness metric optimization into the framework [74], and the use of a semi-definite relaxation based optimization procedure [83] for better phase error estimation performance [51].

2.1.5.4 Sparsity-driven Autofocus (SDA)

In sparsity-driven autofocus (SDA) [52], phase error correction is performed jointly with image reconstruction, rather than as a postprocessing step. The overall process is formulated as a regularized optimization with the cost function

$$J(\mathbf{f}, \boldsymbol{\phi}) = \|\mathbf{y} - \mathbf{H}(\boldsymbol{\phi})\mathbf{f}\|_{2}^{2} + \lambda \|\mathbf{f}\|_{1}$$
(2.56)

where λ is the regularization parameter, which determines the emphasis on the sparsity-promoting ℓ_1 -norm prior, and ϕ is the correct phase of the phase history data. Note that, unlike (2.19), in SDA the model matrix \mathbf{H} depends on ϕ , rather than being constant. This optimization problem is solved through a coordinate descent based numerical iterative algorithm. This algorithm jointly minimizes the cost function with respect to \mathbf{f} and ϕ . Particularly in each iteration, there are image formation, phase error estimation, and phase error correction steps. In image

formation step, ϕ is kept constant and \mathbf{f} is estimated, then in the phase error estimation step, ϕ is updated using the new estimate of \mathbf{f} , and finally in the phase error correction step, the model matrix \mathbf{H} is updated according to the updated ϕ . These three steps are performed iteratively until convergence, or stopping criteria are satisfied. Algorithm 1 summarizes the SDA process. The success of SDA is partially attributed to the use of non-quadratic regularization, as it helps overall algorithm to exhibit robustness to small perturbations on the observation model matrix [105]. In SDA, separate algorithms have been proposed for the cases of 1D phase errors, 2D separable phase errors, and 2D non-separable phase errors.

Algorithm 1: Sparsity-driven autofocus algorithm.

```
Require: \mathbf{H}(\boldsymbol{\phi}^{(0)}), \epsilon, N.

\mathbf{f}^{(0)} \leftarrow \mathbf{H}(\boldsymbol{\phi}^{(0)})^H \mathbf{y}

while \frac{\|\mathbf{f}^{(n+1)} - \mathbf{f}^{(n)}\|}{\|\mathbf{f}^{(n)}\|} \ge \epsilon and n \le N do

\hat{\mathbf{f}}^{(n+1)} = \operatorname{argmin}_{\mathbf{f}} J(\mathbf{f}, \boldsymbol{\phi}^{(n)})

\hat{\boldsymbol{\phi}}^{(n+1)} = \operatorname{argmin}_{\boldsymbol{\phi}} J(\mathbf{f}^{(n+1)}, \boldsymbol{\phi})

Update \mathbf{H}(\boldsymbol{\phi}^{(n+1)}) using \boldsymbol{\phi}^{(n+1)} and \mathbf{H}(\boldsymbol{\phi}^{(n)}).

n \leftarrow n+1

end while
```

2.1.6 SAR ATR

Automatic target recognition (ATR) is an important application of SAR imaging. In ATR, the task is to automatically identify the regions-of-interest (ROIs) that contain targets and then to determine which classes targets belong to. ATR tasks are considered as highly challenging, as SAR images are extremely sensitive to target orientation. Figure 2.4 shows the conventionally reconstructed images of T72 tank in the MSTAR dataset [2] for different orientations relative to the imaging platform.

SAR ATR has been an active and popular research area for decades [106–126]. The process of SAR ATR generally includes four steps; detection, discrimination, feature extraction, and classification [119]. In the detection step, potential ROIs are located, and then falsely detected ROIs are discarded in the discrimination step. In the feature extraction step, distinctive features are extracted from the detected ROIs, and then finally classification task is performed using these extracted features in the last step. In SAR ATR literature, the emphasis is mostly on last two steps with the assumption that the ROIs including targets are already detected and discriminated. In early works, the classification task was done without feature extraction, i.e., by directly using reconstructed images, with pre-defined classifiers. Later studies first introduced feature extraction with hand-picked features, and then learned classi-

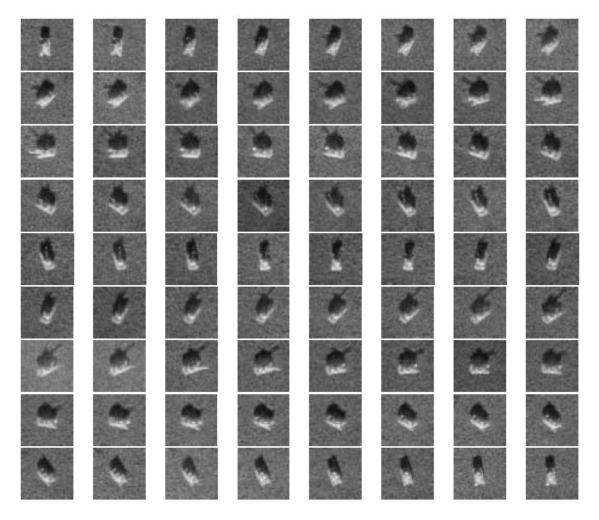


Figure 2.4 SAR images of T72 tank in the MSTAR dataset [2] for different orientations relative to the radar platform. From top left to bottom right, each image corresponds to an orientation angle from 0° to 355° , with 5° increments.

fiers, e.g., using dictionary learning, rather than pre-defined ones. Recent studies combined these two steps using convolutional neural networks (CNNs). The main advantage of using a CNN to learn features rather than using hand-picked features is that the CNN can learn more useful patterns whose existences are not necessarily apparent in the image.

2.2 Deep Learning Background

This section provides background information on deep learning and presents a brief review on the use of deep learning tools within SAR imaging and SAR ATR frameworks.

2.2.1 Brief History of Deep Learning

Deep learning can be considered as one of the solutions to a variety of artificial intelligence (AI) problems. Although it has become extremely popular only in the last decade, deep learning has a long history that dates back to 1940s. However, it has been called with different names at different time windows, i.e., cybernetics in 1940s-1960s, connectionism in the 1980s to early 1990s, artificial neural networks in the 1990s-200s, and only recently, the name 'deep learning' has become widespread [127].

The idea of building a machine that can think is at least over two millennia old and is usually attributed to ancient Greeks [127]. The earliest primitive attempts to build such machines dates around the medieval times, e.g., polymaths such as al-Jazari [128] and Leonardo da Vinci [129] designed machines that are regarded as the ancestors of modern robots.

Although the ideas that evolved into deep learning has emerged around the time that first computers were built, these ideas did not draw much attention until recently due to two main reasons: 1) Computers did not have enough computational resources that would satisfy the needs to train deep neural networks; and 2) deep neural networks need large amounts of data to train, which has only been available after the recent digitization of the society [127].

2.2.2 Deep Neural Network Layout Types

In this section, we will describe four common deep neural network layouts, namely, multilayer perceptron (MLP), autoencoder (AE), convolutional neural network (CNN), and recurrent neural network (RNN).

2.2.2.1 Multilayer Perceptron (MLP)

The perceptron, which was capable of binary classification, was introduced in 1958 [130]. MLPs are extensions of perceptrons which are designed for more complicated tasks. An MLP consists of an input layer, one or more hidden layers, and an output layer. The nodes in the layers of MLPs, apart from the input layer, are made of perceptrons with nonlinear activation functions, and are often called as neurons. Figure 2.5 shows an MLP with an input layer with dimension of 3, an output layer with dimension of 2, and two hidden layers with 4 and 3 neurons, respectively. MLPs are also called as feedforward neural networks, as information only flows from input layer to output layer, i.e., there are no feedback connections in which the output is fed back into the network. MLPs are considered to be main or 'vanilla' neural

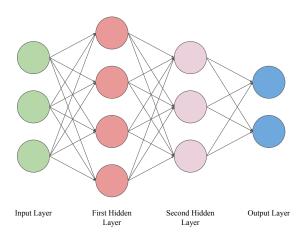


Figure 2.5 An illustration of a multilayer perceptron with two hidden layers.

network as other layouts are derived from MLPs.

2.2.2.2 Autoencoder (AE)

Autoencoders (AE)s are neural networks that try to mimic their input at their output. AEs can be considered as consisting of two parts, i.e., encoder and decoder. The encoder aims to learn a mapping from the input to a transform domain, which typically has a lower dimensionality than the input, and the task of the decoder is to reconstruct the input at the output from the transformed version of the input. However, AEs are not generally used to copy the input at the output, rather they are designed to be unable to learn to copy perfectly. They are usually restricted in ways to ensure that they will only copy approximately, and only copy those inputs which resembles the training data [127]. These restrictions usually lead AEs to learn useful properties of the data. Figure 2.6 shows an illustration of an AE with one hidden layer. AEs can be used in many tasks, e.g., dimensionality reduction, feature learning, manifold learning, information retrieval, and denoising.

2.2.2.3 Convolutional Neural Network (CNN)

Convolutional neural networks (CNNs) are a special kind of MLPs which are suitable for data that has grid-like topology, e.g., images. The earliest CNN-like design was Neocognitron [131] which was built for shift-invariant pattern recognition in 1980s. CNNs are neural networks in which at least in one layer, the matrix multiplication is replaced with the convolution operation. Unlike MLPs, in CNNs, nodes in a layer are only connected to a subset of nodes from the previous and the next layers, hence CNNs have sparse connectivity. Also, within convolutional layers, same parameters are used for each node. Sparse interactions and parameter sharing help with both efficient memory use and translational and shift-invariance. Figure 2.7 shows a CNN

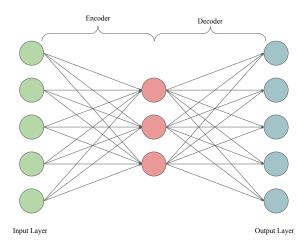


Figure 2.6 An illustration of an autoencoder with a single hidden layer.

with the convolutional kernel size of 3 at every layer. Note that each node is only connected to 3 nodes from the previous layer and 3 nodes from the next layer (except those at the edges) rather than all nodes.

2.2.2.4 Recurrent Neural Network (RNN)

Recurrent neural networks (RNNs) [132] are neural networks that are designed for processing sequential data, e.g., speech signals, videos, stock market values, texts, DNA sequences etc. Unlike previous layouts, RNNs have feedback connections from one or more layers to other (same or previous) layer(s). Figure 2.8 shows an illustration of an RNN in which the value of the hidden node is fed back to itself. RNNs can have finite impulse response (FIR) or infinite impulse response (IIR). FIR RNNs are in the form of directed acyclic graph while IIR RNNs are in the form of directed cyclic graphs. Finite impulse RNNs can be unrolled and replaced with a feedforward neural network [133]. Long short-term memory (LSTM) networks [134] and gated recurrent units (GRUs) [135] are two common layouts of RNNs, both of which can be FIR or IIR.

2.2.3 Deep Learning in SAR Literature

After the resurgence of deep learning in 2006 [136], and especially since the success of AlexNet [35] in ILSVRC 2012, DNNs have enjoyed a growing amount of attention coming from many different research areas, most notably, computer vision and signal processing. DNNs have been used for many image processing tasks in the last decade, including classification, reconstruction, restoration, denoising, deblurring, and super-resolution. Naturally, SAR imaging has also received its share

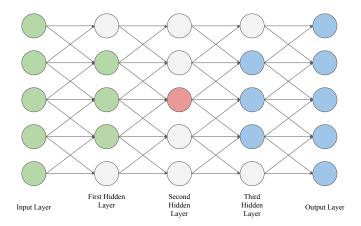


Figure 2.7 An illustration of a CNN with the convolutional kernel size of 3. All the nodes whose values are used to determine the value of the red node in the second hidden layer are shown in green, while all the nodes whose values are determined using the value of the red node are shown in blue.

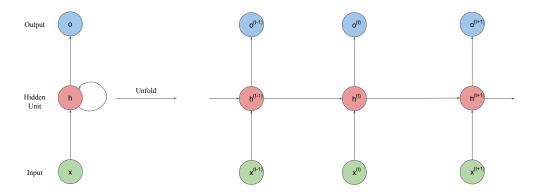


Figure 2.8 An illustration of an RNN with a single input, a single hidden unit and a single output.

from this ever-growing interest. Much like early⁶ deep learning works were mostly concentrated on image classification tasks, initial works on SAR imaging using deep learning tools were on ATR [117, 121–123, 125, 126, 137–140] and other classification or detection applications [141–145]. However, only a few works focused on image formation [146–149].

2.3 Plug-and-Play (PnP) Priors

As we have discussed before, the discrete observation model given in (2.19), like many inverse problems in imaging, is ill-posed, and can only be solved satisfactorily by incorporating regularization into the inversion process. Regularized cost functions typically involve two terms: a data fidelity term to ensure the final image is consistent with the measured data, and a regularizer (prior) term that promotes solutions with desirable properties [150,151]. Regularization-based image reconstruction methods can alleviate problems caused by incomplete data or sparse apertures. Also, they produce images with increased resolution, reduced sidelobes, and reduced speckle by incorporating prior information about the scene. In the past, many regularization-based image reconstruction methods have been proposed including non-quadratic regularization (NQR) based [1,75], sparsity-driven [87,152–154] and dictionary learning (DL) based [3] SAR imaging.

It is not an easy task to minimize a regularized cost function, as the two terms are usually quite different, and it is unlikely to find a single optimization approach that will be appropriate for both. Using a proximal algorithm [155] such as iterative shrinkage thresholding algorithm (ISTA) [156] and alternative direction method of multipliers (ADMM) [34,157] is the natural choice to tackle this issue, as they convert the original optimization problem into a series of smaller optimization problems, hence each term can be minimized separately using an appropriate minimization method. Proximal methods have also found use in SAR imaging tasks [89, 158–160]. The key ingredient in these methods is the proximal operator which involves an optimization problem that is shown to be equivalent to a simple denoising of intermediate solutions under Gaussian white noise [33, 151].

Recently, a new framework, called plug-and-play (PnP) priors has been introduced [33] which replaces the proximal operator by a suitable denoising method. This framework has gained great attention from the computational imaging community [151,161–185] since its emergence. Different denoisers have been adopted within the PnP framework, such as BM3D [186] (as in [33,151,162–164,167,168,170,179]),

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⁶Early within this last resurgence.

non-local means (NLM, as in [162, 164, 173, 176, 182]), NCSR [187] (as in [161]), Gaussian mixture models (GMM, as in [165, 169, 172]), and deep learning based ones [166, 174, 175, 178, 181, 183, 185]. Along with ADMM, and ISTA, various other proximal algorithms, including approximate message passing (AMP) [188], half quadratic splitting (HQS) [49, 50], primal-dual splitting [189], and consensus equilibrium [171] have been used within PnP. This framework has been applied to many imaging problems including reconstruction [33, 151, 164, 174, 175, 179, 180, 183, 185], denoising [166, 176], restoration [162, 165, 167, 170, 172, 173], deblurring [165, 166, 181], super-resolution [161, 162, 173, 181, 182], phase retrieval [163, 185], and hyperspectral sharpening [169].

PnP priors with a variable splitting algorithm such as ADMM enables the integration of data fidelity based on a physical observation model with any explicit or implicit regularizing prior. Next, we will first describe the ADMM algorithm, and then formulate the PnP framework using ADMM.

2.3.1 Alternating Direction Method of Multipliers (ADMM)

Consider the unconstrained optimization problem

$$\hat{\mathbf{x}} = \underset{\mathbf{x}}{\operatorname{argmin}} f(\mathbf{x}) + \lambda g(\mathbf{x}) \tag{2.57}$$

The idea of ADMM is to convert (2.57) into the following constrained problem by variable splitting:

$$\{\hat{\mathbf{x}}, \hat{\mathbf{v}}\} = \underset{\mathbf{x}, \mathbf{v}}{\operatorname{argmin}} f(\mathbf{x}) + \lambda g(\mathbf{v}) \text{ subject to } \mathbf{x} = \mathbf{v}$$
 (2.58)

Next, consider the augmented Lagrangian function corresponding to (2.58):

$$\mathcal{L}(\mathbf{x}, \mathbf{v}, \mathbf{u}) = f(\mathbf{x}) + \lambda g(\mathbf{v}) + \mathbf{u}^{T}(\mathbf{x} - \mathbf{v}) + \frac{\rho}{2} \|\mathbf{x} - \mathbf{v}\|_{2}^{2}$$
(2.59)

where ρ is the penalty parameter of the ADMM.

The minimizer of (2.58) is the saddle point of \mathcal{L} , which can be found by solving a sequence of sub-problems, i.e., the ADMM iterations, which, in *scaled form*, are

given by [34]:

$$\mathbf{x}^{(k+1)} = \underset{\mathbf{x} \in \mathbb{R}^n}{\operatorname{argmin}} f(\mathbf{x}) + \frac{\rho}{2} \left\| \mathbf{x} - \tilde{\mathbf{x}}^{(k)} \right\|_2^2$$
 (2.60)

$$\mathbf{v}^{(k+1)} = \underset{\mathbf{v} \in \mathbb{R}^n}{\operatorname{argmin}} \lambda g(\mathbf{v}) + \frac{\rho}{2} \left\| \mathbf{v} - \tilde{\mathbf{v}}^{(k)} \right\|_2^2$$
 (2.61)

$$\bar{\mathbf{u}}^{(k+1)} = \bar{\mathbf{u}}^{(k)} + (\mathbf{x}^{(k+1)} - \mathbf{v}^{(k+1)})$$
 (2.62)

where $\bar{\mathbf{u}}^{(k)} \triangleq (1/\rho)\mathbf{u}^{(k)}$ is the scaled Lagrange multiplier, $\tilde{\mathbf{x}}^{(k)} \triangleq \mathbf{v}^{(k)} - \bar{\mathbf{u}}^{(k)}$ and $\tilde{\mathbf{v}}^{(k)} \triangleq \mathbf{x}^{(k+1)} + \bar{\mathbf{u}}^{(k)}$. Under mild conditions, e.g., when both f and g are closed, proper and convex, and if a saddle point of \mathcal{L} exists, it can be shown that the iterates (2.60)-(2.62) converge to the solution of (2.58).

2.3.2 PnP ADMM

Note that ADMM iterations (2.60)-(2.62) have a modular structure. (2.60) can be regarded as an inversion step as it involves the forward imaging model $f(\mathbf{x})$ whereas (2.61) can be considered as a denoising step as it is an image domain operation involving the prior $g(\mathbf{v})$. To show that, let us define $\sigma \triangleq \sqrt{\lambda/\rho}$, then (2.61) becomes

$$\mathbf{v}^{(k+1)} = \operatorname*{argmin}_{\mathbf{v} \in \mathbb{R}^n} g(\mathbf{v}) + \frac{1}{2\sigma^2} \left\| \mathbf{v} - \tilde{\mathbf{v}}^{(k)} \right\|_2^2$$
 (2.63)

Treating $\tilde{\mathbf{v}}^{(k)}$ as the noisy image, (2.63) minimizes the residue between $\tilde{\mathbf{v}}^{(k)}$ and clean image \mathbf{v} using the prior $g(\mathbf{v})$. For example, if $g(\mathbf{v}) = ||\mathbf{v}||_{\mathsf{TV}}$, i.e., the total variation norm, then (2.63) is the standard total variation denoising problem.

Venkatakrishnan et al. [33] proposed a variant of the ADMM algorithm by suggesting that one does not need to specify g before running the ADMM. Instead, they replaced (2.61) by an off-the-shelf image denoising algorithm, denoted by \mathcal{D}_{σ} , to yield

$$\mathbf{v}^{(k+1)} = \mathcal{D}_{\sigma}(\tilde{\mathbf{v}}^{(k)}) \tag{2.64}$$

which they called Plug-and-Play (PnP) ADMM. The convergence guarantee of the ADMM does not necessarily hold in general, for the PnP-ADMM. Accordingly, convergence is a topic of ongoing research [162, 177, 178, 184].

Chapter 3

PnP-CNN-SAR for Image Reconstruction

Previously, we have discussed that the discrete observation model given in (2.19) is ill-posed [41], and can be considered as ill-conditioned to a certain extent depending on the specifics of the observation scenario. Hence, we have concluded that it can only be solved satisfactorily using regularization. In this chapter, we introduce the PnP-CNN-SAR framework for SAR image reconstruction. In this framework, the problem of SAR image reconstruction is formulated as a regularized optimization problem, then using the PnP priors framework with ADMM, the objective function is divided into subproblems and the proximal operator is replaced with a CNN-based prior. The final portion of this chapter contains experimental results, demonstrating the effectiveness of the proposed PnP-CNN-SAR framework. This is a joint work with Ammar Saleem, another Ph.D. student in our group.

3.1 Objective Function

Consider the discrete observation model introduced in (2.19):

$$y = Hf + n \tag{3.1}$$

where $\mathbf{y} \in \mathbb{C}^{M \times 1}$ is the complex valued column-stacked observation vector from which we desire to estimate the underlying column-stacked SAR image $\mathbf{f} \in \mathbb{C}^{N \times 1}$ where $M \leq N^1$, and $\mathbf{H} \in \mathbb{C}^{M \times N}$ is the Fourier transform-based forward model. Let

¹Note that, in Chapter 2, we have used K and M to denote numbers of sampled observation points in range and cross-range directions, respectively. However, in this chapter, we only use column-stacked vectors for both phase histories and underlying SAR image, i.e., \mathbf{y} and \mathbf{f} , throughout the formulation of the proposed

 \mathcal{R} denote the regularization function we impose on the reflectivity field. Then the regularized objective function becomes

$$\hat{\mathbf{f}} = \underset{\mathbf{f}}{\operatorname{argmin}} \|\mathbf{y} - \mathbf{H}\mathbf{f}\|_{2}^{2} + \lambda \mathcal{R}(|\mathbf{f}|)$$
(3.2)

Note that regularization is performed on the magnitude of the reflectivity field \mathbf{f} . For most SAR scenes, reflectivity phase at a certain location can be modeled as a uniformly distributed random variable uncorrelated with the phase at other locations [11]. Nevertheless, we need to estimate the complex-valued field \mathbf{f} in the process of ensuring good data fidelity. We can write $\mathbf{f} = \mathbf{\Theta}\mathbf{f}_m$ where $\mathbf{\Theta}$ is a diagonal matrix containing the phases of \mathbf{f} at each pixel in exponentiated form, i.e., $\mathbf{\Theta} = \text{diag}(e^{j\varphi(\mathbf{f})})$ where $\varphi(\cdot)$ denotes the phase, and \mathbf{f}_m denotes the magnitude of \mathbf{f} . Then, replacing \mathbf{f} with $\mathbf{\Theta}\mathbf{f}_m$ in (3.2), we get:

$$\{\hat{\mathbf{f}}_m, \hat{\mathbf{\Theta}}\} = \underset{\mathbf{f}_m, \mathbf{\Theta}}{\operatorname{argmin}} ||\mathbf{y} - \mathbf{H}\mathbf{\Theta}\mathbf{f}_m||_2^2 + \lambda \mathcal{R}(\mathbf{f}_m)$$
 (3.3)

It is pertinent to mention that, up to this point, the explicit definition of a regularizer has not been formulated. The regularization will be explained in Section 3.2 after we formulate ADMM iterations for (3.3) and decouple the data-fidelity term and the regularization term. By doing so, we will be able to plug-in any regularizer into the objective function without affecting the data-fidelity.

3.2 Variable Splitting and ADMM

Rewriting (3.3) in a suitable form by introducing an auxiliary variable with a constraint, we have

$$\{\hat{\mathbf{f}}_{m}, \hat{\mathbf{\Theta}}, \hat{\mathbf{h}}\} = \underset{\mathbf{f}_{m}, \mathbf{\Theta}, \mathbf{h}}{\operatorname{argmin}} ||\mathbf{y} - \mathbf{H}\mathbf{\Theta}\mathbf{f}_{m}||_{2}^{2} + \lambda \mathcal{R}(\mathbf{h})$$
s.t. $\mathbf{f}_{m} - \mathbf{h} = 0$ (3.4)

The augmented Lagrangian for (3.4) is given by

$$\mathcal{L}\left(\hat{\mathbf{f}}_{m}, \hat{\boldsymbol{\Theta}}, \hat{\mathbf{h}}, \hat{\mathbf{u}}\right) = ||\mathbf{y} - \mathbf{H}\boldsymbol{\Theta}\mathbf{f}_{m}||_{2}^{2} + \lambda \mathcal{R}(\mathbf{h}) + \frac{\rho}{2}||\mathbf{f}_{m} - \mathbf{h} + \mathbf{u}||_{2}^{2} + \frac{\rho}{2}||\mathbf{u}||_{2}^{2}$$
(3.5)

Let $\tilde{\mathbf{f}}^{(k)} = \mathbf{h}^{(k)} - \mathbf{u}^{(k)}$ and $\tilde{\mathbf{h}}^{(k)} = \mathbf{f}_m^{(k+1)} + \mathbf{u}^{(k)}$. Also, let us introduce a vector $\boldsymbol{\theta} \in \mathbb{C}^{N \times 1}$ that contains the diagonal elements of the phase matrix $\boldsymbol{\Theta}$, and the matrix \boldsymbol{B} whose diagonal elements are the reflectivity magnitudes \mathbf{f}_m . Finally, let us invoke

method. Hence, for notational simplicity, we use M and N to denote sizes of y and f, respectively.

the constraint that the magnitudes of the elements of $\boldsymbol{\theta}$ denoted as $|\theta_i|$ should be 1, simply because they contain phases of \mathbf{f} at each pixel in exponentiated vector form $e^{j\varphi(\mathbf{f})}$. Then, each iteration of the ADMM algorithm will perform the following steps:

$$\boldsymbol{\theta}^{(k+1)} = \underset{\boldsymbol{\theta}}{\operatorname{argmin}} ||\mathbf{y} - \mathbf{H}\mathbf{B}^{(k)}\boldsymbol{\theta}||_{2}^{2} + \lambda_{\boldsymbol{\theta}} \sum_{i=1}^{N} (|\theta_{i}| - 1)^{2}$$
(3.6)

$$\mathbf{f}_{m}^{(k+1)} = \underset{\mathbf{f}_{m}}{\operatorname{argmin}} ||\mathbf{y} - \mathbf{H}\boldsymbol{\Theta}^{(k)}\mathbf{f}_{m}||_{2}^{2} + \frac{\rho}{2}||\mathbf{f}_{m} - \tilde{\mathbf{f}}^{(k)}||_{2}^{2}$$

$$(3.7)$$

$$\mathbf{h}^{(k+1)} = \underset{\mathbf{h}}{\operatorname{argmin}} \lambda \mathcal{R}(\mathbf{h}) + \frac{\rho}{2} ||\tilde{\mathbf{h}}^{(k)} - \mathbf{h}||_{2}^{2}$$
(3.8)

$$\mathbf{u}^{(k+1)} = \mathbf{u}^{(k)} + \mathbf{f}_m^{(k+1)} - \mathbf{h}^{(k+1)}$$
(3.9)

where λ_{θ} and ρ are hyper-parameters.

The first two steps only depend on the choice of the forward model while the third step only depends on the choice of the regularizer. Therefore, using this plug-andplay framework, we can now write the minimization as two independent modules. Thus, this framework can be used to mix and match different prior models with forward models.

Sub-problem (3.6) is solved through a fixed point algorithm, which can also be shown to be equivalent to a particular quasi-Newton iterative minimization algorithm:

$$\mathbf{G}\boldsymbol{\theta}^{(n+1)} = (\mathbf{H}\mathbf{B})^H \mathbf{y} + \lambda_{\boldsymbol{\theta}} e^{j\varphi(\boldsymbol{\theta}^{(n)})}$$
(3.10)

where

$$\mathbf{G} = (\mathbf{H}\mathbf{B})^H (\mathbf{H}\mathbf{B}) + \lambda_{\theta} \mathbf{I} \tag{3.11}$$

Each iteration in (3.10) involves solving a set of linear equations. This is an indication that the approach here solves the non-quadratic optimization problem in (3.6) by turning it into a series of quadratic problems. This linear set of equations can efficiently be solved by the conjugate gradient algorithm with the convergence criterion $\frac{\|\boldsymbol{\theta}^{(n+1)} - \boldsymbol{\theta}^{(n)}\|}{\|\boldsymbol{\theta}^{(n)}\|} < 10^{-14}$. For this sub-problem, it is possible that one can come up with a more efficient algorithm than ours.

Sub-problem (3.7) has a closed form solution, i.e., taking the derivative with respect to \mathbf{f}_m and equating it to zero results in the following:

$$(\frac{\rho}{2}\mathbf{I} + \mathbf{\Theta}^H \mathbf{H}^H \mathbf{H} \mathbf{\Theta})\hat{\mathbf{f}}_m = \mathbf{\Theta}^H \mathbf{H}^H \mathbf{y} + \frac{\rho}{2}\tilde{\mathbf{f}}$$
(3.12)

We solve this complex-valued linear system using the conjugate gradient algorithm. This produces a complex solution. Therefore, we address it by considering its real part only and equating any negative value to zero, as in practice the imaginary part of this solution has values close to zero which can easily be ignored, and the real part usually has only a small number of negative values at most (which are also very close to zero). Note that within each iteration of ADMM, sub-problems (3.6) and (3.7) are solved iteratively until the convergence criterion is satisfied, before moving to the sub-problem (3.8). This process is called the model update.

Sub-problem (3.8) can be rewritten as

$$\mathbf{h}^{(k+1)} = \underset{\mathbf{h}}{\operatorname{argmin}} \frac{1}{2(\sqrt{\lambda/\rho})^2} ||\tilde{\mathbf{h}}^{(k)} - \mathbf{h}||_2^2 + \mathcal{R}(\mathbf{h})$$
(3.13)

which is equivalent to denoising the image $\tilde{\mathbf{h}}^{(k)}$ with a Gaussian denoiser with a noise level $\sqrt{\lambda/\rho}$ [190]. Hence, any Gaussian denoiser can act as a modular component of the overall algorithm to solve (3.8). In the work we present here, denoising is achieved through a trained CNN. Note that the parameter λ is not explicitly defined as it depends on ρ and the network settings. The details about the architecture and training will be discussed in Section 3.3.1.

Note that (3.9) is not an optimization problem, rather it is a direct update on the Lagrange multiplier \mathbf{u} . Algorithm 2 shows the summary of the PnP-CNN-SAR. Note that $\mathbf{f}_m^{[t]}$ stands for the outputs of the inner loop, i.e., the model update, while $\mathbf{f}_m^{(k)}$ stands for the outputs of the outer loop, i.e., the overall ADMM algorithm. Typical values for K and T are 20 and 100, respectively.

3.3 Experimental Results

In this section, we demonstrate the effectiveness of PnP-CNN-SAR on synthetic and real SAR scenes. We also tabulate results for various scenarios and compare them with FFT-based, NQR-based [1], and DL-based [3] SAR image reconstructions as well as a PnP framework with BM3D [186] regularizer. Finally, we give an analysis on the effect of training different networks for different scenarios, in case of synthetic scene experiments.

3.3.1 Setup

The experiments are conducted with variation in two major aspects, first is the type of observation SAR scene, second is the parameters of SAR observation phenomenon. The observation SAR scenes are further bifurcated to synthetic and real scenes. The

Algorithm 2: PnP-CNN-SAR for image reconstruction.

```
Require: \mathbf{y}, \mathbf{H}, \epsilon, \mathcal{D}_{\sigma}(x), K, T.
       \mathbf{f}^{(0)} \leftarrow \mathbf{H}^H \mathbf{y} {Conventional reconstruction}
       \mathbf{B}^{(0)} \leftarrow \operatorname{diag}\{\mathbf{f}_m^{(0)}\}\
       \boldsymbol{\theta}^{(0)} \leftarrow \angle \mathbf{f}^{(0)}
      \mathbf{h}^{(0)} \leftarrow \mathbf{f}_m^{(0)} \\ \mathbf{u}^{(0)} \leftarrow 0
       k \leftarrow 0
      while \frac{\|\mathbf{u}^{(k+1)} - \mathbf{u}^{(k)}\|}{\|\mathbf{u}^{(k)}\|} \ge \epsilon and k \le K do
              \tilde{\textbf{f}}^{(k)} \leftarrow \textbf{h}^{(k)} - \textbf{u}^{(k)}
            \mathbf{f}_{m}^{[0]} \leftarrow \mathbf{f}_{m}^{(k)}
\mathbf{while} \frac{\left\|\mathbf{f}_{m}^{[t+1]} - \mathbf{f}_{m}^{[t]}\right\|}{\left\|\mathbf{f}_{m}^{[t]}\right\|} \ge 10^{-4} \text{ and } t \le T \text{ do}
                      Calculate \boldsymbol{\theta}^{(t)} using (3.6) {Phase alignment}
                      Solve for \mathbf{f}_{m}^{(t)} using (3.7) {Magnitude update}
               end while
             \mathbf{f}_{m}^{(k+1)} \leftarrow \mathbf{f}_{m}^{[t]}
\tilde{\mathbf{h}}^{(k)} \leftarrow \mathbf{f}_{m}^{(k+1)} + \mathbf{u}^{(k)}
\mathbf{h}^{(k+1)} \leftarrow \mathcal{D}_{\sigma}(\tilde{\mathbf{h}}^{(k)}) \text{ {Prior update}}
\mathbf{u}^{(k+1)} \leftarrow \mathbf{u}^{(k)} + \mathbf{f}_{m}^{(k+1)} - \mathbf{h}^{(k+1)}
               k \leftarrow k + 1
       end while
      return \mathbf{f}_m^{(k+1)}
```

details of the real scenes are given in Section 3.3.3. It is pertinent to mention that, SAR scenes are complex valued but the synthetic scenes are real-valued and we only have access to the reflectivity magnitudes of the real scenes. Therefore, we add uniform random phase distributed over $[-\pi,\pi]$ to these real-valued (synthetic and real) SAR scenes to generate/simulate complex valued SAR reflectivities of a SAR scene.

The secondary segregation of the experiments brings SAR observation phenomenon into account. It is based on parameters of forward model \mathbf{H} and noise \mathbf{n} . The forward model \mathbf{H} consists of band-limited 2D Fourier transform where the band-limitation is denoted in terms of data availability percentage L. In particular $L = N_a/N_d$ where N_a is the number of available phase history samples and N_d is the number of phase history samples in full band-width data. In case of real SAR scene, full band-width data is not available, therefore, the existing bandwidth of the reference SAR scene is assumed to be full or 100%. The phase histories are acquired from the complex

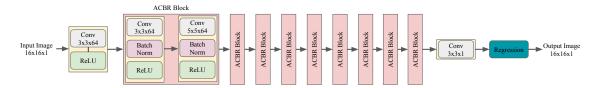


Figure 3.1 Network architecture for synthetic scene experiments. ACBR stands for Alternating Convolutional-Batch normalization-ReLU layers.

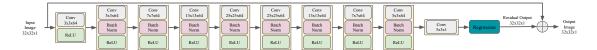


Figure 3.2 Network architecture for real scene experiments.

valued SAR reflectivities using the approach described in [12, 191]. For the implementation of different data availability levels, we crop the appropriate amount of phase history data with equal band-limitation in the range and azimuth directions. In the Fourier perspective, its a low-pass square box filter where the box is located in the center of the 2D Fourier spectrum, essentially eliminating high frequency complex-valued frequency components. The area of the box is calculated from the data availability level L. The box filtering effect is essentially limiting the transmitted signal bandwidth and the range of look angles, leading to reductions in nominal range and cross-range resolution, respectively. Note that in spotlight-mode SAR, phase history data (after pre-processing steps) consist of samples of the band-limited spatial Fourier transform of the reflectivity field on a polar grid (see Eq. (2.16) and Refs. [12, 191]). Conventional polar format imaging interpolates the data to a rectangular grid and then performs a 2D FFT. We use a slightly simplified forward model and generate band-limited data on a rectangular, rather than polar, grid. In practice, this would correspond to adding one more data preprocessing step, namely polar to rectangular interpolation, and then posing the image formation problem based on such interpolated data on the rectangular grid. The other parameter involved in SAR observation process is noise n, which is assumed to be i.i.d. in each pixel and circularly-symmetric complex Gaussian, whose real and imaginary parts are normally distributed with zero mean and $\sigma_{\mathbf{n}}$ standard deviation.

In both synthetic and real scene experiments, we have considered various scenarios, i.e., two different noise levels ($\sigma_{\mathbf{n}} \in \{0.1, 1\} \times \sigma_{\mathbf{y}}$, where $\sigma_{\mathbf{y}}$ is the standard deviation of the magnitude of the phase history data) and 6 different data availability levels in the phase history domain (100%, 90%, 80%, 70%, 50%, and 30%), hence, in total 12 types of experiments are conducted.

Our experimental procedure is composed of two phases, namely, network training and image reconstruction using the PnP-ADMM algorithm. The network training

is performed on two different architectures, one each for synthetic and real scene experiments. The network architecture for the synthetic scene experiments is shown in Figure 3.1. The architecture used for real scene experiments is inspired from U-Net [192] and is shown in Figure 3.2². It is worth noting that the network architectures we used for the synthetic and real scene experiments are different. These selections were made based on empirical exploration of several architectures for each scenario. These choices were driven by the types of features and the sizes of the scenes. While the architectures used here provide good results, we do not claim to have found the best ones. It would also be of interest to find architectures that perform well over all scenarios. The architecture we have proposed for our real scene experiments should be appropriate for generic real SAR imaging tasks. For the synthetic scene experiments, we have trained 12 different networks of the same architecture, i.e., one for each type of experiments considered, while for the real scene experiments, a single network is trained, which corresponds to the case of 100% data availability and $\sigma_{\rm n} = \sigma_{\rm y}$.

In training, we perform conventional reconstruction from the noisy phase histories to get noisy images. We then extract overlapping patches using the magnitudes of the conventionally reconstructed images, $(16 \times 16 \text{ for the synthetic scene experiments})$ and 32×32 for the real scene experiments), and finally augment the patches by rotation, with angles [90°, 180°, 270°]. We train CNNs to learn a mapping from these noisy patches to their clean counterparts in the case of synthetic scene experiments, and we perform residual learning for the real scene experiments. Note that as we do not have ground truth for the real scene experiments, we start our experiments with high-resolution TerraSAR-X reconstructions, which we call reference images. In the context of our experiments, these reference images can be viewed as conventionally reconstructed images from 'full-bandwidth' noiseless data, i.e., without any bandwidth reduction we impose, and without any noise we add on the phase history data. Based on these reference images, we then generate simulated data for various scenarios by appropriate Fourier transformation, band limitation, and noise addition operations. The training of the networks for the synthetic scene experiments took around 3 hours each, while the network for the real scene experiments is trained roughly in 3 days, on MATLAB R2019b, using NVIDIA TITAN Xp GPU. Networks used in the synthetic scene experiments are trained with RMSProp [193] solver, 10^{-3} learning rate, 0.9 learning rate drop factor with the period of one epoch, mini-batch size of 1500 for 100 epochs, while the network used in the real scene experiments is trained with RMSProp [193], 10^{-4} learning rate, 0.9 learning rate drop factor with

-

²We acknowledge that the architecture we propose is inspired from the U-Net architecture, however it is not precisely U-Net.

the period of one epoch, mini-batch size of 500 for 20 epochs.

In image reconstruction, we follow the steps mentioned in Algorithm 2. We initialize the solution with the conventional FFT-based reconstruction. Then we solve equations (3.6) and (3.7) iteratively (within each iteration of ADMM), as mentioned in Section 3.2. Then we solve (3.8) by a denoising procedure using our trained CNN. The procedure include, first extracting overlapping patches from the input image $(16 \times 16 \text{ patches for the synthetic, and } 32 \times 32 \text{ patches for the real scene cases}),$ and feeding the patches to the CNN to get denoised patches. Then we combine the overlapping patches to obtain the overall denoised image. These steps are repeated iteratively until the stopping criteria is satisfied. The stopping criterion is chosen as $\frac{\|\mathbf{u}^{(k+1)} - \mathbf{u}^{(k)}\|}{N} < \epsilon$. Since the data availability and noise levels have a great impact on the performance, various ϵ values are employed for different experiments. We observed that as the data availability reduces or the noise increases, i.e., as the problem at hand gets more challenging, a higher value of ϵ yields better results, while for the less challenging cases, a smaller ϵ is preferred. For synthetic experiments, we used ϵ values ranging from 5×10^{-5} to 5×10^{-3} , while for the real case, the range was from 1×10^{-3} to 1×10^{-1} . The value of ρ is chosen as 12 throughout the experiments and the value of λ_{θ} is calculated from the data availability level, i.e., $\lambda_{\theta} = \frac{2}{L^2}$ where L is the data availability level.

We compare our proposed framework with FFT-based, NQR-based [1], DL-based [3] reconstructions, and a PnP framework with BM3D [186] regularizer (PnP-BM3D) in terms of signal-to-noise ratio (SNR) and structural similarity (SSIM) index [194] values of the reconstructed images. The formula of SSIM is given as:

$$SSIM(\mathbf{x}, \mathbf{y}) = \frac{(2\mu_{\mathbf{x}}\mu_{\mathbf{y}} + c_1)(2\sigma_{\mathbf{x}\mathbf{y}} + c_2)}{(\mu_{\mathbf{x}}^2 + \mu_{\mathbf{y}}^2 + c_1)(\sigma_{\mathbf{x}}^2 + \sigma_{\mathbf{y}}^2 + c_2)}$$
(3.14)

where $\mu_{\mathbf{x}}$, $\mu_{\mathbf{y}}$, $\sigma_{\mathbf{x}}$, $\sigma_{\mathbf{y}}$, and $\sigma_{\mathbf{x}\mathbf{y}}$ are the local means, standard deviations, and cross-covariance for images \mathbf{x} and \mathbf{y} . $c_1 = (k_1 l)^2$ and $c_2 = (k_2 l)^2$ are constants to maintain formula validity, avoiding the denominator being zero. l represents the dynamic range of the pixel value. $k_1 = 0.01$ and $k_2 = 0.03$ are default values. A larger SSIM value indicates a better similarity of the two images.

The problem formulation for the NQR-based [1] reconstruction is given as:

$$\hat{\mathbf{f}} = \underset{\mathbf{f}}{\operatorname{argmin}} \|\mathbf{y} - \mathbf{H}\mathbf{f}\|_{2}^{2} + \lambda_{1} \|\mathbf{f}\|_{p}^{p} + \lambda_{2} \|\mathbf{D}\mathbf{f}_{m}\|_{p}^{p}$$
(3.15)

where **D** is a discrete approximation to the 2D derivative operator (gradient). Here, the first term is the data-fidelity term while second and third terms are aimed at enhancing point-based and region-based features respectively. The parameters λ_1

and λ_2 are optimized to obtain the highest SNR for each image in each scenario while p = 0.9 is used for all the considered scenarios.

The problem formulation for the DL-based [3] reconstruction is given as:

$$\{|\hat{\mathbf{f}}|, \hat{\boldsymbol{\Theta}}, \hat{\mathbf{D}}, \hat{\alpha}_{i}\} = \underset{\mathbf{f}_{m}, \boldsymbol{\Theta}, \mathbf{D}, \alpha_{i}}{\operatorname{argmin}} \lambda \|\mathbf{y} - \mathbf{H}\boldsymbol{\Theta}\mathbf{f}_{m}\|_{2}^{2}$$

$$+ \sum_{i} \|\mathbf{E}_{i}|\mathbf{f}| - \mathbf{D}\alpha_{i}\|_{2}^{2} + \sum_{i} \mu_{i} \|\alpha_{i}\|_{0}$$
s.t. $|\boldsymbol{\Theta}_{ij}| = 1 \ \forall j$ (3.16)

where **D** and α_i are the dictionary and the sparse coefficients to be learned, respectively, and \mathbf{E}_i are the patch extraction operators. See [3] for details.

For PnP-BM3D experiments, the steps in the Algorithm 2 are followed, and in the prior update step, the BM3D prior is used instead of the CNN-based prior. The MATLAB implementation of BM3D is downloaded from the authors' website³.

3.3.2 Synthetic Scene Experiments

The training and the test images for the synthetic scene experiments are shown in Figures 3.3a and 3.3b, respectively. Note that training is not performed using the training images themselves, rather 16×16 overlapping patches that are extracted from the training images. Consequently, in every ADMM iteration, during the prior update step, overlapping patches are extracted, each patch is passed through the network, and then patches are combined to get the reconstructed image.

Although we have trained a network for each scenario, we have compared the performances of all networks in each scenario. Table 3.1 shows the average SNR and SSIM values for different experiments using different networks, respectively. Note that, the network that gives the best result for a particular scenario is not always the one that is trained for that scenario.

The best results from Table 3.1 are also presented in Table 3.2 for comparison with other methods. Table 3.2 shows that our framework outperforms other methods in terms of both SNR and SSIM in 11 out of 12 scenarios. Figure 3.4 shows the effects of the noise level and data availability on the SNR and SSIM values.

Figures 3.5 and 3.6 show reconstruction results for *Image* 7 with noise $\sigma_{\mathbf{n}} = 0.1 \sigma_{\mathbf{y}}$, and *Image* 2 with noise $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$, respectively, for qualitative analysis. It appears that our proposed framework is better at preserving the details and more robust to

 $^{^3}$ http://www.cs.tut.fi/~foi/GCF-BM3D/index.html#ref_software

Table 3.1 Average [SNR/SSIM] values for synthetic scene experiments using different networks. Each row represents a network trained with a particular scenario considered, and each column represents a test scenario.

	20% 30%	-0.551/0.184 -1.632/0.139 1.176/0.217 -0.452/0.166 1.543/0.265 -0.266/0.178		0.752/0.276 -2.433/0.150 2.062/ 0.293 -1.443/0.164 2.285/0.291 -1.307/0.167 2.660 /0.287 -1.244/0.177 1.868/0.180 1.027 /0.142 -0.209/0.072 0.351/0.039
	%02	-0.051/0.230 -0 1.611/0.297 1 1.107/0.300 1		3.182/ 0.420 4.356/0.388 2. 4.614/0.401 2.340/0.174 2.0340/0.177
$\sigma_{\mathbf{y}}$	%08	0.273/0.258 1.518/0.341 0.571/0.286	-0.836/0.191 -2.928/0.080 -4.070/0.040	5.015/0.484 5.034/0.437 5.279/0.458 3.689/0.368 -0.481/0.174 -3.163/0.061
	%06	0.629/0.282 1.625/0.360 0.648/0.305	-1.008/0.197 -3.000/0.087 -4.032/0.054	6.425/0.539 5.444/0.468 5.312/0.470 3.588/0.378 -0.786/0.182 -3.390/0.063
	100%	1.623/0.347 2.585/0.439 2.055/0.413	1.160/0.385 -0.891/0.270 -2.386/0.190	8.242/0.610 6.749/0.541 6.710/0.549 4.383/0.461 1.335/0.346 -1 651/0.212
	30%	-2.008/0.198 -1.074/0.240 -0.956/0.237	1.043/ 0.278 2.041 /0.276 -1.349/0.164	-3.082/0.171 -2.542/0.183 -2.187/0.205 -2.147/0.217 0.184/0.241 1.713/0.173
	20%	0.217/0.317 2.975/0.406 3.754/0.465	5.480/0.490 3.748/0.392 -1.896/0.133	0.381/0.383 1.584/0.407 1.652/0.394 1.981/0.406 4.152/0.363 2.303/0.212
	%02	4.146/0.504 9.931/0.689 11.417/0.726	9.103/0.660 3.781/0.458 -2.366/0.104	4.791/0.590 6.662/0.605 5.468/0.570 5.308/0.550 5.482/0.454 1.176/0.205
$0.1\sigma_{\mathbf{y}}$	%08	11.447/0.699 16.552/0.816 15.254/0.802	11.820/0.747 4.404/0.521 -2.378/0.111	8.767/0.701 10.229/0.700 8.353/0.668 7.090/0.636 6.167/0.502 0.560/0.216
	%06		12.741/0.766 5.414/0.567 -2.459/0.119	10.773/0.743 11.295/0.731 10.197/0.712 7.952/0.665 6.647/0.537 0.900/0.253
	100%	32.170/0.905 19.074/ 0.869 21.282/0.861 19.57 4/0.848 17.571/0.840 16.839/0.828	14.088/0.804 8.948/0.716 2.352/0.517	13.572/0.800 12.913/0.780 11.849/0.759 8.647/0.716 8.294/0.649 4.764/0.536
		100% 90% 80%	70% 50% 30%	100% 90% 80% 70% 50%
		,	$0.1\sigma_{\mathbf{y}}$	$\sigma_{\mathbf{y}}$

Table 3.2 Average SNR and SSIM values for different noise and data availability levels for the synthetic scene experiments. Best results are shown in bold, second best results are shown in red.

		SNR	(dB)	SSIM		
Available Data	Method	$0.1\sigma_{\mathbf{y}}$	$\sigma_{\mathbf{y}}$	$0.1\sigma_{\mathbf{y}}$	$\sigma_{\mathbf{y}}$	
	FFT-based	25.55	5.84	0.906	0.449	
	NQR-based [1]	27.31	12.36	0.985	0.631	
100%	DL-based [3]	26.31	6.13	0.877	0.343	
	PnP-BM3D	25.95	5.98	0.869	0.351	
	PnP-CNN-SAR	37.92	13.99	0.996	0.687	
	FFT-based	11.29	5.62	0.539	0.366	
	NQR-based [1]	17.84	11.03	0.845	0.549	
90%	DL-based [3]	24.76	4.81	0.869	0.269	
	PnP-BM3D	17.77	5.88	0.737	0.289	
	PnP-CNN-SAR	25.33	12.18	0.954	0.591	
	FFT-based	9.65	5.64	0.474	0.336	
	NQR-based [1]	15.55	10.02	0.809	0.486	
80%	DL-based [3]	22.20	5.42	0.852	0.272	
	PnP-BM3D	12.14	5.89	0.544	0.254	
	PnP-CNN-SAR	22.30	11.03	0.921	0.488	
	FFT-based	7.94	5.47	0.390	0.294	
	NQR-based [1]	10.83	8.97	0.596	0.431	
70%	DL-based [3]	12.22	6.02	0.566	0.254	
	PnP-BM3D	8.68	5.84	0.409	0.229	
	PnP-CNN-SAR	17.17	10.37	0.831	0.433	
	FFT-based	6.00	5.26	0.289	0.227	
	NQR-based [1]	8.97	8.02	0.432	0.316	
50%	DL-based [3]	7.44	6.55	0.344	0.209	
	PnP-BM3D	7.09	5.44	0.293	0.186	
	PnP-CNN-SAR	11.23	8.41	0.596	0.335	
	FFT-based	4.41	4.52	0.191	0.158	
	NQR-based [1]	7.70	7.32	0.296	0.263	
30%	DL-based [3]	5.71	6.31	0.216	0.156	
	PnP-BM3D	5.36	5.15	0.188	0.131	
	PnP-CNN-SAR	7.79	6.78	0.390	0.216	

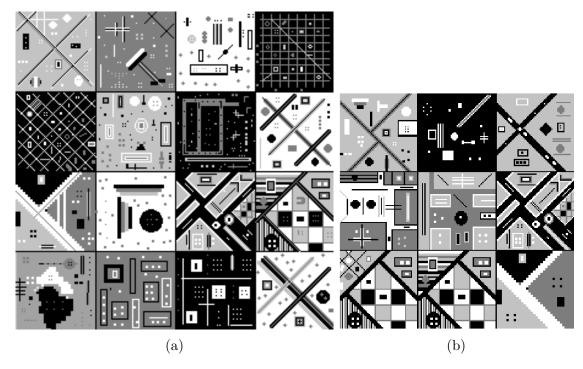


Figure 3.3 Synthetic SAR scenes used for (a) training the CNN and learning the dictionary, and (b) performance evaluation.

data limitations, and gives the best visual quality in the majority of scenarios.

3.3.3 Real Scene Experiments

For the real scene experiments, TerraSAR-X [195] images are used and we designed two types of experimental settings in which we considered the following scenarios: 1) When the training and testing are performed on non-overlapping windows extracted from the same scene that are randomly split into training and test data sets, and 2) when the training is done using non-overlapping windows extracted from one scene, and testing is performed on non-overlapping windows extracted from a different scene. These two experimental settings will be called same scene experiments, and different scene experiments, respectively. For these experiments, two TerraSAR-X images are used to form the data set. The first image, shown in Figure 3.7, contains the view of the city of Wonsan in Democratic People's Republic of Korea and, the second image, shown in Figure 3.8, contains the view of the Kapıkule Border Crossing near the Turkey-Bulgaria-Greece border tripoint. These images will be referred as Wonsan and Kapikule, respectively. The data collected for the Wonsan image has the incidence angle of 42.88° while the incidence angle for Kapıkule is 44.77°. After post-processing, both images have the pixel resolution of 1.60 m \times 1.60 m.

158 non-overlapping 170×170 windows are extracted from the Wonsan image, and

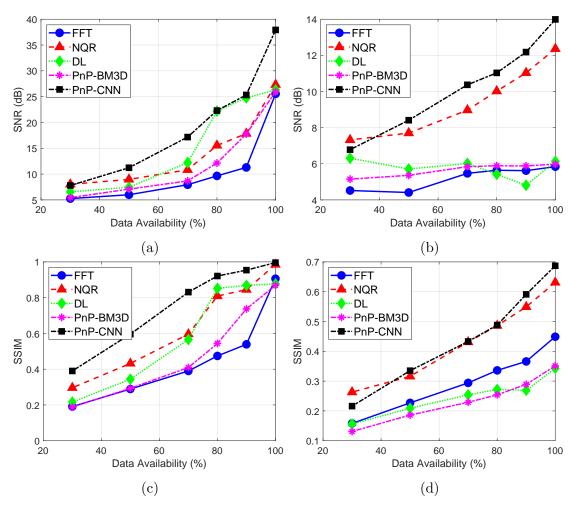


Figure 3.4 Average SNR values for synthetic scene experiments for (a) $\sigma_{\bf n}=0.1\sigma_{\bf y}$, and (b) $\sigma_{\bf n}=\sigma_{\bf y}$. Average SSIM values for synthetic scene experiments for (c) $\sigma_{\bf n}=0.1\sigma_{\bf y}$, and (d) $\sigma_{\bf n}=\sigma_{\bf y}$.

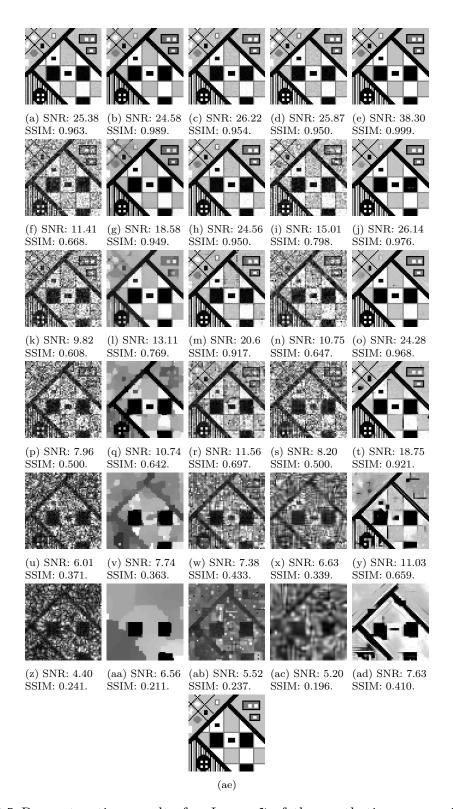


Figure 3.5 Reconstruction results for Image~7 of the synthetic scenes with noise $\sigma_{\bf n}=0.1\sigma_{\bf y}$. First row, full data, second row, 90% data, third row, 80% data, fourth row, 70% data, fifth row, 50% data, sixth row, 30% data. First column, FFT-based reconstruction, second column, NQR-based reconstruction, third column, DL-based reconstruction, fourth column, PnP-BM3D based reconstruction, fifth column, PnP-CNN-SAR based reconstruction. (ae) Ground truth.

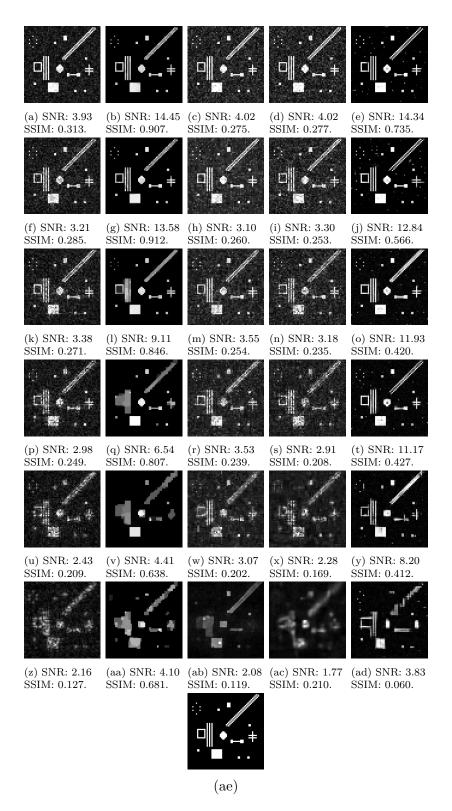


Figure 3.6 Reconstruction results for Image~2 of the synthetic scenes with noise $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$. First row, full data, second row, 90% data, third row, 80% data, fourth row, 70% data, fifth row, 50% data, sixth row, 30% data. First column, FFT-based reconstruction, second column, NQR-based reconstruction, third column, DL-based reconstruction, fourth column, PnP-BM3D based reconstruction, fifth column, PnP-CNN-SAR based reconstruction. (ae) Ground truth.



Figure 3.7 Wonsan image.

110 non-overlapping windows with the same size are extracted from the Kapıkule image. Randomly selected 127 windows from the Wonsan image are used for training the network that is to be used for both scenarios. Remaining 31 windows from the Wonsan image constitute the test set for the same scene experiments, while all windows from the Kapıkule image are used as the test set for the different scene experiments. Figures 3.9, 3.10, and 3.11 show the training windows, test windows for the same scene experiments, and the test windows for the different scene experiments, respectively. Note that the number of training windows are not enough to train the network and therefore the training is performed using 32×32 overlapping patches that are extracted from the training windows. Hence, the prior update step of the ADMM iteration is performed the same way as described in Section 3.3.2.

3.3.3.1 Same Scene Experiments

Columns under "Wonsan" in Table 3.3 show the average SNR and SSIM values for the same scene experiments. Figure 3.12 shows the effects of the noise level and data availability on the SNR and SSIM values. Results show that our framework outperforms other methods in 10 scenarios in terms of SNR, and in all scenarios in terms of SSIM. More detailed results can be found in the supplementary material.

Figures 3.13 and 3.14 show the reconstruction results for *Image 26* with noise $\sigma_{\mathbf{n}} = 0.1 \sigma_{\mathbf{y}}$ and *Image 12* with noise $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$, respectively. It appears that the perceptual quality of the images produced by PnP-CNN-SAR is better, and our framework shows its value especially in the case of higher noise, where other methods tend to produce blurry and/or noisy images. It is worth noting that, as the amount



Figure 3.8 Kapıkule image.

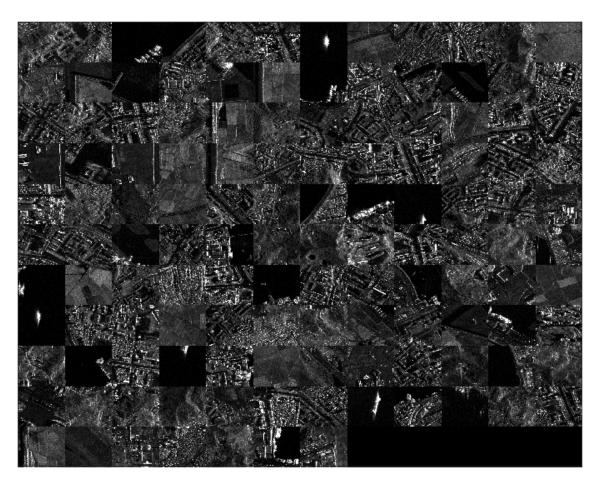


Figure 3.9 Windows extracted from the Wonsan image that are used for training the CNN for PnP-CNN-SAR and learning the dictionary for DL-based reconstruction.

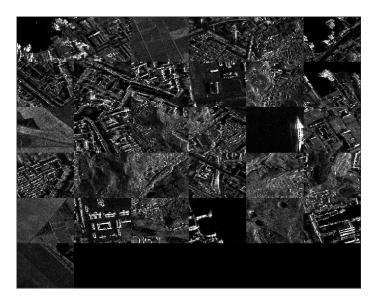


Figure 3.10 Windows extracted from the Wonsan image that are used for performance evaluation in the same scene experiments.



Figure 3.11 Windows extracted from the Kapıkule image that are used for performance evaluation in the different scene experiments.

Table 3.3 Average SNR and SSIM values for different noise and data availability levels for the real scene experiments. Best results are shown in bold, second best results are shown in red.

		Wonsan				Kapıkule			
	Method	SNR (dB)		SSIM		SNR (dB)		SSIM	
Available Data		$0.1\sigma_{\mathbf{y}}$	$\sigma_{\mathbf{y}}$	$0.1\sigma_{\mathbf{y}}$	$\sigma_{\mathbf{y}}$	$0.1\sigma_{\mathbf{y}}$	$\sigma_{\mathbf{y}}$	$0.1\sigma_{\mathbf{y}}$	$\sigma_{\mathbf{y}}$
	FFT-based	26.65	6.33	0.973	0.457	26.63	6.62	0.989	0.597
	NQR-based [1]	23.56	9.70	0.963	0.535	23.11	10.28	0.964	0.537
100%	DL-based [3]	22.95	16.13	0.976	0.924	24.66	11.08	0.981	0.647
	PnP-BM3D	26.65	6.42	0.973	0.462	26.63	6.75	0.979	0.459
	PnP-CNN-SAR	40.03	36.82	0.998	0.997	38.52	36.27	0.998	0.997
	FFT-based	13.23	6.16	0.802	0.423	13.21	6.49	0.758	0.516
	NQR-based [1]	11.64	9.09	0.683	0.486	11.70	9.65	0.705	0.492
90%	DL-based [3]	12.71	8.90	0.783	0.538	14.22	10.21	0.822	0.584
	PnP-BM3D	14.89	6.10	0.838	0.418	15.27	6.71	0.844	0.422
	PnP-CNN-SAR	16.78	15.85	0.900	0.857	16.87	15.43	0.892	0.851
	FFT-based	10.02	6.01	0.704	0.390	10.07	6.39	0.662	0.471
	NQR-based [1]	8.83	8.07	0.506	0.409	9.47	8.79	0.518	0.434
80%	DL-based [3]	9.97	8.07	0.671	0.470	11.22	9.50	0.699	0.520
	PnP-BM3D	11.65	6.18	0.731	0.400	12.36	6.82	0.729	0.401
	PnP-CNN-SAR	13.29	12.55	0.830	0.781	13.60	12.98	0.796	0.749
	FFT-based	8.41	5.84	0.632	0.362	8.41	6.28	0.576	0.426
	NQR-based [1]	7.49	7.19	0.427	0.355	8.22	8.29	0.439	0.391
70%	DL-based [3]	8.65	7.54	0.587	0.419	9.65	8.94	0.601	0.467
	PnP-BM3D	9.53	6.19	0.632	0.379	10.49	7.03	0.608	0.384
	PnP-CNN-SAR	10.75	10.34	0.736	0.693	11.20	10.99	0.688	0.649
	FFT-based	6.22	5.42	0.510	0.315	6.33	5.86	0.358	0.275
	NQR-based [1]	5.86	6.60	0.295	0.301	7.78	7.59	0.393	0.355
50%	DL-based [3]	6.85	6.51	0.450	0.333	7.63	7.70	0.455	0.380
	PnP-BM3D	7.68	6.15	0.452	0.334	8.81	7.39	0.419	0.348
	PnP-CNN-SAR	8.11	8.09	0.580	0.562	8.62	8.81	0.527	0.516
	FFT-based	4.53	4.57	0.394	0.267	4.67	4.94	0.220	0.182
	NQR-based [1]	5.71	6.14	0.270	0.284	7.01	6.84	0.346	0.323
30%	DL-based [3]	5.14	5.06	0.334	0.258	5.67	5.89	0.339	0.298
	PnP-BM3D	5.98	5.87	0.336	0.277	6.77	7.14	0.330	0.312
	PnP-CNN-SAR	5.56	5.89	0.406	0.363	5.98	6.48	0.384	0.339

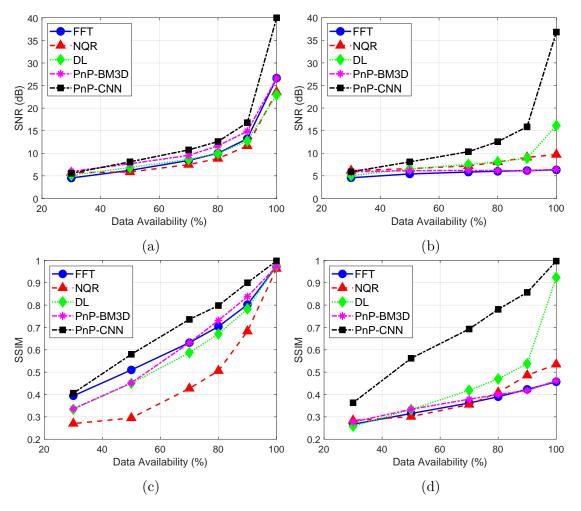


Figure 3.12 Average SNR values for same scene experiments for (a) $\sigma_{\mathbf{n}} = 0.1 \, \sigma_{\mathbf{y}}$, and (b) $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$. Average SSIM values for same scene experiments for (c) $\sigma_{\mathbf{n}} = 0.1 \, \sigma_{\mathbf{y}}$, and (d) $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$.

of available data reduces, the degradation of the reconstruction quality is more prominent in real case, compared to the synthetic case. This is most likely due to that in real case, all experiments are performed using the same network, i.e., trained with 100% data and $\sigma_{\bf n} = \sigma_{\bf y}$ noise, while in synthetic case, networks trained for all scenarios considered and results of the network that performed the best in each scenario are presented. A similar degradation scheme is also visible in case of different scene experiments, which will be discussed next.

3.3.3.2 Different Scene Experiments

Columns under "Kapıkule" in Table 3.3 show the average SNR and SSIM values for different scene experiments. Figure 3.15 shows the effects of the noise level and data availability on the SNR and SSIM values. Note that since FFT-based, NQR-based [1], and PnP-BM3D methods are not learning-based, the issue of training and testing on different scenes is only relevant to DL-based [3] and PnP-CNN-SAR.

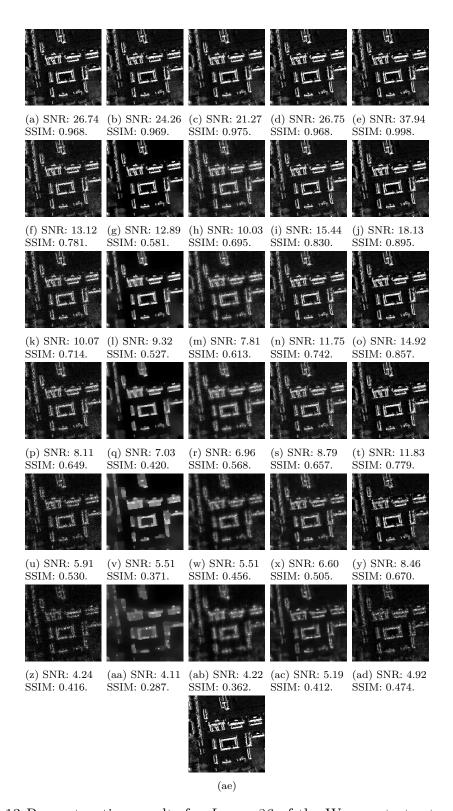


Figure 3.13 Reconstruction results for Image~26 of the Wonsan test set with noise $\sigma_{\mathbf{n}}=0.1\sigma_{\mathbf{y}}$. First row, full data, second row, 90% data, third row, 80% data, fourth row, 70% data, fifth row, 50% data, sixth row, 30% data. First column, FFT-based reconstruction, second column, NQR-based reconstruction, third column, DL-based reconstruction, fourth column, PnP-BM3D based reconstruction, fifth column, PnP-CNN-SAR based reconstruction. (ae) Reference image.

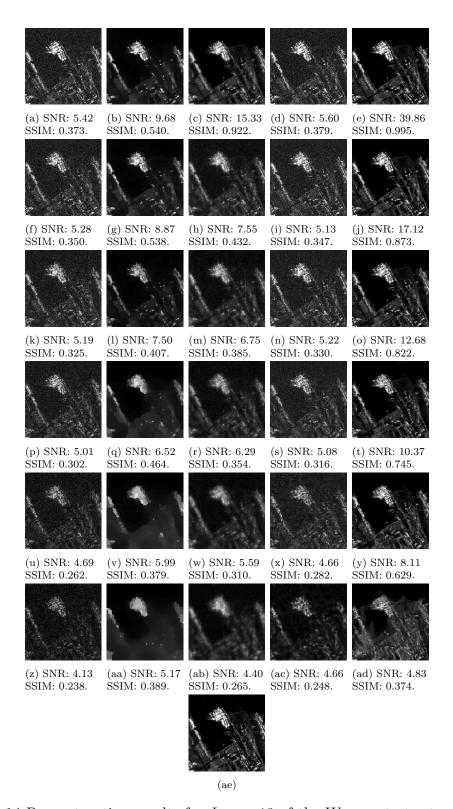


Figure 3.14 Reconstruction results for *Image 12* of the Wonsan test set with noise $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$. First row, full data, second row, 90% data, third row, 80% data, fourth row, 70% data, fifth row, 50% data, sixth row, 30% data. First column, FFT-based reconstruction, second column, NQR-based reconstruction, third column, DL-based reconstruction, fourth column, PnP-BM3D based reconstruction, fifth column, PnP-CNN-SAR based reconstruction. (ae) Reference image.

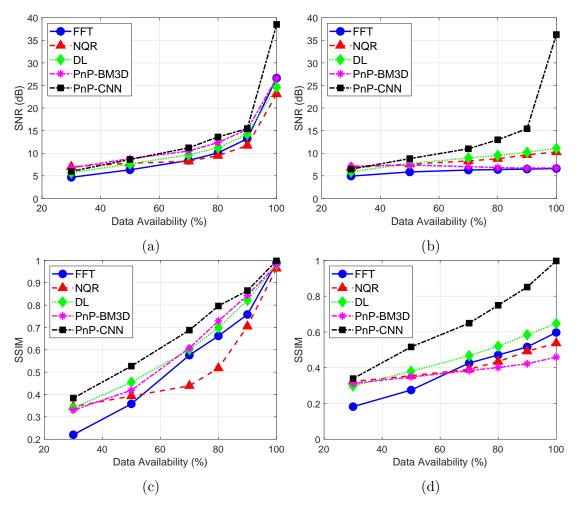


Figure 3.15 Average SNR values for different scene experiments for (a) $\sigma_{\mathbf{n}} = 0.1 \sigma_{\mathbf{y}}$, and (b) $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$. Average SSIM values for different scene experiments for (c) $\sigma_{\mathbf{n}} = 0.1 \sigma_{\mathbf{y}}$, and (d) $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$.

Nevertheless, we show results from all methods. Results show that our framework outperforms other methods in 9 scenarios in terms of SNR and, in all scenarios in terms of SSIM. More detailed results can be found in the supplementary material.

Figures 3.16 and 3.17 show the reconstruction results for *Image 16* with noise $\sigma_{\mathbf{n}} = 0.1 \sigma_{\mathbf{y}}$ and *Image 78* with noise $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$, respectively. Once again, PnP-CNN-SAR produces visually better reconstructions, and is able to capture/preserve more features and details compared to other methods. The performance of our approach on the different scene experiments suggests that it offers good generalization capability.

3.3.4 Training Different Networks for Different Scenarios

In Table 3.1 we have shown the average SNR and SSIM values that are achieved using different networks in each scenario. Our first observation is that, the network that gives the best result for a particular scenario in terms of SNR, is not necessarily



Figure 3.16 Reconstruction results for Image~16 of the Kapıkule test set with noise $\sigma_{\mathbf{n}}=0.1\sigma_{\mathbf{y}}$. First row, full data, second row, 90% data, third row, 80% data, fourth row, 70% data, fifth row, 50% data, sixth row, 30% data. First column, FFT-based reconstruction, second column, NQR-based reconstruction, third column, DL-based reconstruction, fourth column, PnP-BM3D based reconstruction, fifth column, PnP-CNN-SAR based reconstruction. (ae) Reference image.

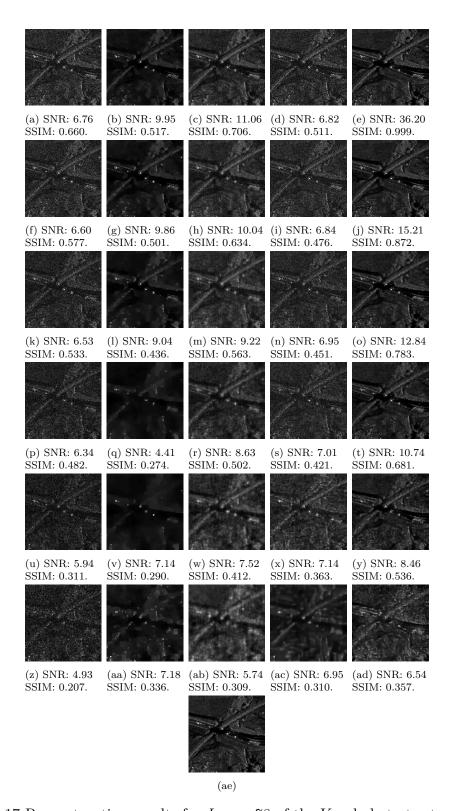


Figure 3.17 Reconstruction results for *Image 78* of the Kapıkule test set with noise $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$. First row, full data, second row, 90% data, third row, 80% data, fourth row, 70% data, fifth row, 50% data, sixth row, 30% data. First column, FFT-based reconstruction, second column, NQR-based reconstruction, third column, DL-based reconstruction, fourth column, PnP-BM3D based reconstruction, fifth column, PnP-CNN-SAR based reconstruction. (ae) Reference image.

Table 3.4 Decision process of which network to use in which case.

	Data availability is known	Data availability is unknown
Noise level is	Use a network trained with same	Use a network trained with the
known	amount of noise and higher data	same noise level and 70-90% data
	availability	availability
Noise level is	Use a network trained with high	Use a network trained with high
unknown	noise and higher data availability	noise and 70-90% data availabil-
		ity

the one trained for that scenario, rather, in most cases, the one that is trained with slightly higher data availability and same noise level. Another important observation is that, in each 6×6 block of the Table 3.1, the SNR values almost always decrease from left to right, which shows that regardless of for which case the network is trained, it is always easier for the network to reconstruct images from higher data availability levels. One last thing to note is that the performance of the networks trained with higher noise on the scenarios with lower noise is much better than the performance of the networks trained with lower noise on the scenarios with higher noise. These observations tell us that, when deciding for what kind of network one should train and use, a good rule of thumb would be; if both the data availability and the noise levels are known for the test scenario at hand, use a network trained with the same amount of noise and a slightly higher data; if only noise level is known, using a network trained with 70-90% data and same noise level would be a reasonable approach; if only data availability is known, safest approach would be to train with slightly higher data and high noise; and if neither is known, training with 70-90% data availability and high noise would be preferred approach. Table 3.4 shows the summary of the suggested decision process.

Reconstruction results in Figures 3.5 and 3.6 show that SNR and SSIM do not measure the performance of the methods well, especially in the case of lower data availability levels, i.e., 50% and 30%. To further elaborate that, in Figure 3.18, we show the reconstruction results of all 12 networks for *Image* 7 of the synthetic test set for the case of 30% data availability and noise $\sigma_{\bf n}=0.1\sigma_{\bf y}$. While the network trained with 50% data availability and noise $\sigma_{\bf n}=0.1\sigma_{\bf y}$ performs the best in terms of both SNR and SSIM, it is not entirely clear that this is the best reconstruction in terms of visual quality. It seem that each network was able capture some features from the original scene and missing many others. There also seems to be a trend that networks trained with lower data availability tend to generate smoother reconstructions, compared to networks trained with higher data availability levels.

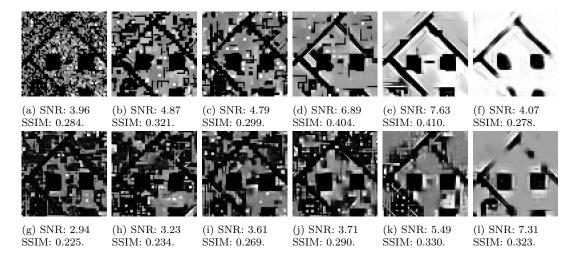


Figure 3.18 Reconstruction results of Image 7 of the synthetic test set in the case of 30% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$ using different networks. First row, results of the networks trained with noise $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$. Second row, results of the networks trained with noise $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$. Each column represents a data availability level, in a decreasing order from left to right, i.e., (a) is the result of the network trained with full data and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$, (b) is the result of the network trained with 90% data and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$ etc. Note that (e) has the highest SNR among all reconstructions and hence, is the one presented in Figure 3.5ad.

3.3.5 Runtime Analysis

Table 3.5 shows the comparison of runtimes across different methods for the same scene experiments. PnP-CNN-SAR works faster than the NQR-based method in all scenarios, and PnP-BM3D in two scenarios. It is also worth noting that PnP-CNN-SAR works considerably slower in lower data availability scenarios, compared to the scenarios with higher data availability.

Figures 3.19 and 3.20 show the percentage shares of various steps within the overall algorithm time for the same scene experiments, for PnP-CNN-SAR. These results show that the prior update step dominates the overall procedure and takes around 60-90% of the computation time, while the phase alignment step never exceeds 30%. Magnitude update and other steps do not contribute as much to the computational expense, compared to these two steps. Reconstructing the entire image at once, instead of using a patch-based approach, would likely reduce the computation time of the prior update step. However this would be at the expense of increasing the computation time of the phase alignment step, which is observed in the PnP-BM3D experiments, for which the results are shown in Figures 3.21 and 3.22. Note however that abandoning the patch-based approach would likely result in performance degradation. For comparison, we also show the percentage shares of various steps within the overall algorithm time for the same scene experiments, for DL-based reconstruction [3], in Figures 3.23 and 3.24.

Table 3.5 Average reconstruction time for one image for the same scene experiments. FFT-based reconstruction is not shown as it practically works in real time.

Scenario	NQR-based [1]	DL-based [3]	PnP-BM3D	PnP-CNN-SAR
Full Data, $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$	206.8 s	12.0 s	13.6 s	16.7 s
90% Data, $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$	223.1 s	$15.6 \; { m s}$	19.3 s	23.0 s
80% Data, $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$	229.7 s	17.9 s	31.8 s	22.1 s
70% Data, $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$	232.8 s	19.7 s	27.6 s	24.5 s
50% Data, $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$	236.5 s	22.4 s	34.8 s	102.8 s
30% Data, $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$	232.1 s	27.0 s	$25.2 \mathrm{\ s}$	91.4 s
Full Data, $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$	209.9 s	15.1 s	19.1 s	25.7 s
90% Data, $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$	229.4 s	16.9 s	17.6 s	27.4 s
80% Data, $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$	235.7 s	17.8 s	10.9 s	27.1 s
70% Data, $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$	230.2 s	21.6 s	11.3 s	26.5 s
50% Data, $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$	225.3 s	$27.3 \mathrm{\ s}$	16.0 s	113.8 s
30% Data, $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$	218.5 s	$32.0 \mathrm{\ s}$	$18.6 \mathrm{\ s}$	84.8 s

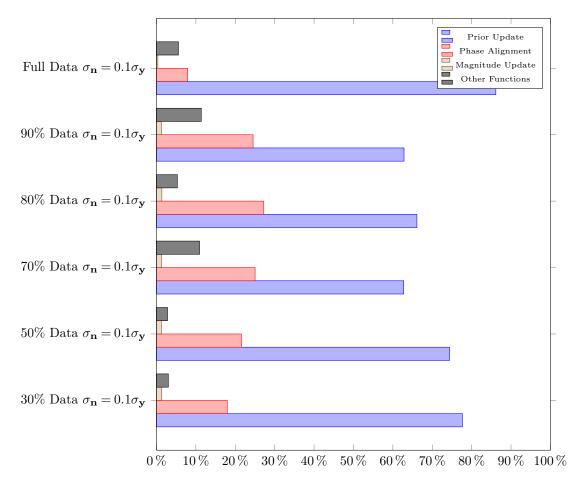


Figure 3.19 Percentage shares of various steps within the overall algorithm time for the same scene experiment scenarios with $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$, for PnP-CNN-SAR.

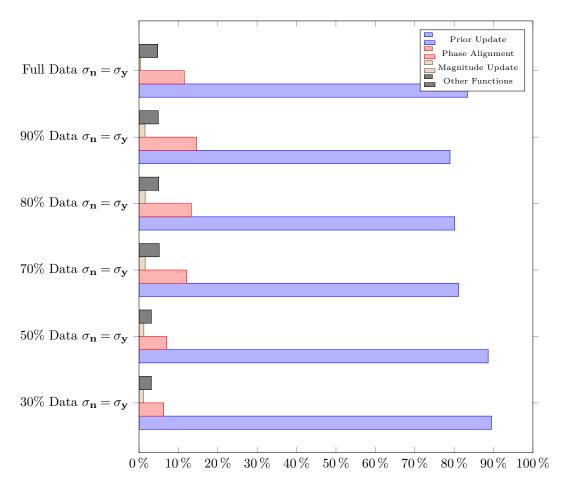


Figure 3.20 Percentage shares of various steps within the overall algorithm time for the same scene experiment scenarios with $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$, for PnP-CNN-SAR.

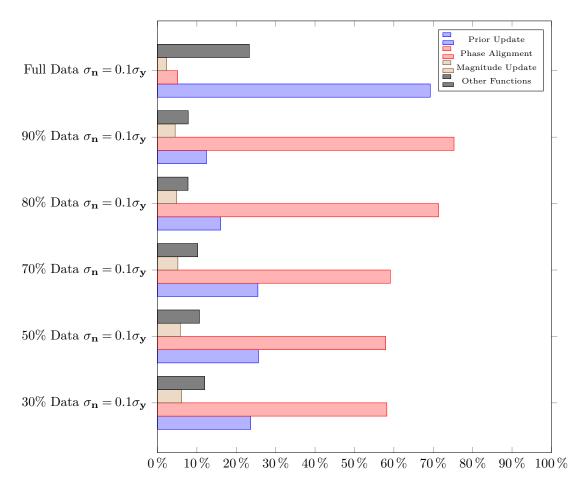


Figure 3.21 Percentage shares of various steps within the overall algorithm time for the same scene experiment scenarios with $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$, for PnP-BM3D.

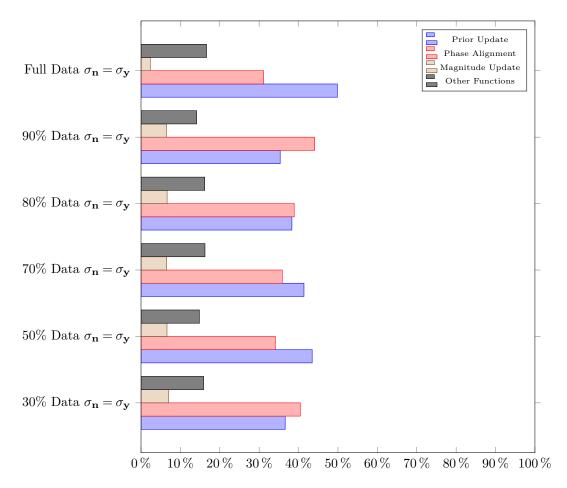


Figure 3.22 Percentage shares of various steps within the overall algorithm time for the same scene experiment scenarios with $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$, for PnP-BM3D.

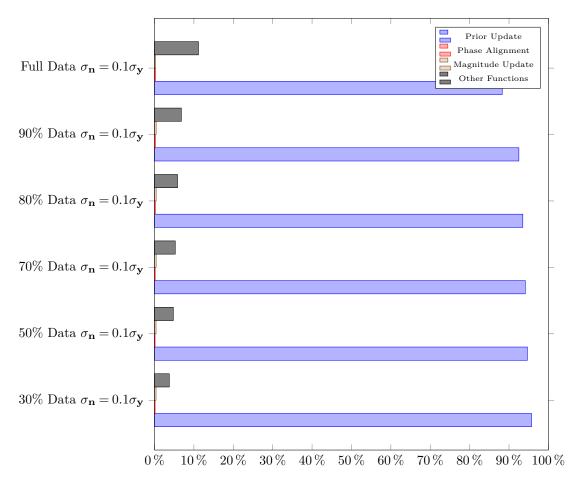


Figure 3.23 Percentage shares of various steps within the overall algorithm time for the same scene experiment scenarios with $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$, for DL-based reconstruction [3].

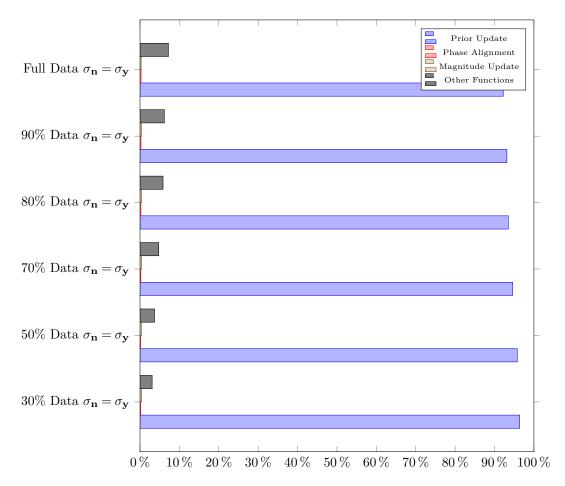


Figure 3.24 Percentage shares of various steps within the overall algorithm time for the same scene experiment scenarios with $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$, for DL-based reconstruction [3].

3.4 Summary

The results in this chapter show that performing regularization using deep priors with the help of PnP priors framework can produce better reconstructions in terms of visual quality compared to conventional as well as state-of-the-art SAR image reconstruction methods, especially in the more challenging, lower data availability and higher noise scenarios, and even when the performance metrics fail to fully reflect the performance of our proposed framework. Overall, these results suggest that deep learning methods may have the potential to learn complicated spatial patterns and enable their incorporation as priors into computational radar imaging.

In this chapter, we have only focused on image reconstruction, and did not account for possible phase errors. In Chapter 4, we extend the PnP-CNN-SAR framework to jointly address image reconstruction and phase error correction.

Chapter 4

PnP-CNN-SAR-AF for Joint Image Reconstruction and Phase Error Correction

In Chapter 3, we have introduced the PnP-CNN-SAR framework for SAR image reconstruction and demonstrated its effectiveness in various scenarios. However, in this framework, we did not account for phase errors. Although PnP-CNN-SAR can be combined with a postprocessing autofocus algorithm to eliminate phase errors, previous studies have shown that joint image reconstruction and phase error correction can produce better results than performing these two tasks sequentially [52,75,78,79,84,92,93]. Hence, in this chapter, we extend the PnP-CNN-SAR framework to address phase errors along with image reconstruction and introduce the PnP-CNN-SAR-AF framework (where AF stands for autofocus). This chapter includes the formulation and preliminary results of PnP-CNN-SAR-AF for 1D phase errors, and the formulation and implementation for 2D separable and non-separable phase errors are left for future work.

4.1 Objective Function

Consider the modified discrete observation model:

$$\mathbf{y} = \mathbf{H}(\boldsymbol{\phi})\mathbf{f} + \mathbf{n} \tag{4.1}$$

where ϕ stands for the 1D phase error. Note that the observation model matrix **H** is no longer constant, rather it is a function of ϕ . Now let us introduce the regularized

objective function to solve (4.1):

$$\hat{\mathbf{f}}, \hat{\boldsymbol{\phi}} = \underset{\mathbf{f}, \boldsymbol{\phi}}{\operatorname{argmin}} \|\mathbf{y} - \mathbf{H}(\boldsymbol{\phi})\mathbf{f}\|_{2}^{2} + \lambda \mathcal{R}(\mathbf{f}_{m})$$
(4.2)

where $\mathbf{f}_m = |\mathbf{f}|$. As we have already discussed in Chapter 3, for most SAR scenes, reflectivity phase at a certain location can be modeled as a uniformly distributed random variable uncorrelated with the phase at other locations [11]. Nevertheless, we need to estimate the complex-valued field \mathbf{f} in the process of ensuring good data fidelity. Again, we can write $\mathbf{f} = \mathbf{\Theta} \mathbf{f}_m$ where $\mathbf{\Theta}$ is a diagonal matrix containing the phases of \mathbf{f} at each pixel in exponentiated form, i.e., $\mathbf{\Theta} = \operatorname{diag}(e^{j\varphi(\mathbf{f})})$ where $\varphi(\cdot)$ denotes the phase. Then we get:

$$\hat{\mathbf{f}}_{m}, \hat{\boldsymbol{\theta}}, \hat{\boldsymbol{\phi}} = \underset{\mathbf{f}_{m}, \boldsymbol{\theta}, \boldsymbol{\phi}}{\operatorname{argmin}} \|\mathbf{y} - \mathbf{H}(\boldsymbol{\phi})\boldsymbol{\Theta}\mathbf{f}_{m}\|_{2}^{2} + \lambda \mathcal{R}(\mathbf{f}_{m}). \tag{4.3}$$

Once again, we do not explicitly define the regularizer until we formulate ADMM iterations for (4.3), and decouple the data-fidelity term and the regularization term.

4.2 Variable Splitting and ADMM

Rewriting (4.3) in a suitable form by introducing an auxiliary variable with a constraint, we have

$$\hat{\mathbf{f}}_{m}, \hat{\boldsymbol{\theta}}, \hat{\boldsymbol{\phi}}, \hat{\mathbf{h}} = \underset{\mathbf{f}_{m}, \boldsymbol{\theta}, \boldsymbol{\phi}, \mathbf{h}}{\operatorname{argmin}} \|\mathbf{y} - \mathbf{H}(\boldsymbol{\phi})\boldsymbol{\Theta}\mathbf{f}_{m}\|_{2}^{2} + \lambda \mathcal{R}(\mathbf{h}) \text{ s.t. } \mathbf{f}_{m} - \mathbf{h} = 0.$$
(4.4)

The augmented Lagrangian for (4.4) is given by

$$\mathcal{L}(\hat{\mathbf{f}}_{m}, \hat{\boldsymbol{\theta}}, \hat{\boldsymbol{\phi}}, \hat{\mathbf{h}}, \hat{\mathbf{u}}) = \|\mathbf{y} - \mathbf{H}(\boldsymbol{\phi})\boldsymbol{\Theta}\mathbf{f}_{m}\|_{2}^{2} + \lambda \mathcal{R}(\mathbf{h}) + \frac{\rho}{2} \|\mathbf{f}_{m} - \mathbf{h} + \mathbf{u}\|_{2}^{2} + \frac{\rho}{2} \|\mathbf{u}\|_{2}^{2}.$$
(4.5)

Let $\tilde{\mathbf{f}}^{(k)} = \mathbf{h}^{(k)} - \mathbf{u}^{(k)}$, $\tilde{\mathbf{h}}^{(k)} = \mathbf{f}_m^{(k+1)} + \mathbf{u}^{(k)}$. Also, let us introduce a vector $\boldsymbol{\theta} \in \mathbb{C}^{N \times 1}$ that contains the diagonal elements of the phase matrix $\boldsymbol{\Theta}$, and the matrix \mathbf{B} whose diagonal elements are the reflectivity magnitudes \mathbf{f}_m . Finally, let us invoke the constraint that the magnitudes of the elements of $\boldsymbol{\theta}$ denoted as $|\theta_i|$ should be 1, simply because they contain phases of \mathbf{f} at each pixel in exponentiated vector form $e^{j\varphi(\mathbf{f})}$. Then, each iteration of the ADMM algorithm will perform model update, phase error estimation and correction, prior update, and Lagrange multiplier update steps.

Model update:

$$\boldsymbol{\theta}^{(k+1)} = \underset{\boldsymbol{\theta}}{\operatorname{argmin}} \left\| \mathbf{y} - \mathbf{H}(\boldsymbol{\phi}^{(k)})^{(k)} \mathbf{B}^{(k)} \boldsymbol{\theta} \right\|_{2}^{2} + \lambda_{\boldsymbol{\theta}} \sum_{i=1}^{N} (|\theta_{i}| - 1)^{2}$$
(4.6)

$$\mathbf{f}_{m}^{(k+1)} = \underset{\mathbf{f}_{m}}{\operatorname{argmin}} \left\| \mathbf{y} - \mathbf{H}(\boldsymbol{\phi}^{(k)})^{(k)} \boldsymbol{\Theta}^{(k)} \mathbf{f}_{m} \right\|_{2}^{2} + \frac{\rho}{2} \left\| \mathbf{f}_{m} - \tilde{\mathbf{f}}^{(k)} \right\|_{2}^{2}. \tag{4.7}$$

Phase error estimation and correction:

For every cross-range position $p = 1 \dots P$

$$\Delta \phi_p^{(k+1)} = -\arctan(-\frac{I}{R}) \tag{4.8}$$

where

$$R = \Re\{(\mathbf{f}^{(k+1)})^H \mathbf{H}_p(\phi_p^{(k)})^H \mathbf{y}_p\}$$

$$\tag{4.9}$$

and

$$I = \Im\{(\mathbf{f}^{(k+1)})^H \mathbf{H}_p(\phi_p^{(k)})^H \mathbf{y}_p\}$$

$$\tag{4.10}$$

$$\phi_p^{(k+1)} = \phi_p^{(k)} + \Delta \phi_p^{(k+1)} \tag{4.11}$$

$$\mathbf{H}_{p}(\phi_{p}^{(k+1)}) = \exp\{j\Delta\phi_{p}^{(k+1)}\}\mathbf{H}_{p}(\phi_{p}^{(k)}). \tag{4.12}$$

Prior update:

$$\mathbf{h}^{(k+1)} = \underset{\mathbf{h}}{\operatorname{argmin}} \lambda \mathcal{R}(\mathbf{h}) + \frac{\rho}{2} \left\| \tilde{\mathbf{h}}^{(k)} - \mathbf{h} \right\|_{2}^{2}. \tag{4.13}$$

Lagrange multiplier update:

$$\mathbf{u}^{(k+1)} = \mathbf{u}^{(k)} + \mathbf{f}_m^{(k+1)} - \mathbf{h}^{(k+1)}. \tag{4.14}$$

 $\lambda_{\boldsymbol{\theta}}$ and ρ are hyper-parameters.

The model update and phase error estimation and correction steps only depend on the forward model, while the prior update step only depends on the regularizer. ADMM first iteratively solves for θ and \mathbf{f}_m using (4.6) and (4.7) until the phase is aligned. Then for each cross-range position p, corresponding phase error is estimated using (4.8), and the model matrix is updated using (4.11) and (4.12). Note that (4.6), (4.7), and (4.13) are optimization problems, while (4.8)-(4.12) and (4.14) are direct updates on the corresponding variables.

As is done in Chapter 3, sub-problems (4.6) and (4.7) are solved with conjugate gradient algorithm using (3.10) and (3.12), respectively. Sub-problem (4.13) is once again replaced with a CNN. However, in contrast to PnP-CNN-SAR, in PnP-CNN-SAR-AF, we investigate training the CNN using images corrupted only with noise, as well as using noisy and phase error-corrupted images. Algorithm 3 summarizes the PnP-CNN-SAR-AF framework. Note that $\mathbf{f}_m^{[t]}$ stands for the outputs of the inner loop, i.e., the model update, while $\mathbf{f}_m^{(k)}$ stands for the outputs of the outer loop, i.e., the overall ADMM algorithm. Typical value for T is 100, while the value of K depends on the amount of phase error in the data, and can be as low as 20, and as high as 1000.

4.3 Experimental Results

In this section, we demonstrate the effectiveness of PnP-CNN-SAR-AF on real SAR scenes. We also tabulate results for various scenarios, and compare them with FFT-based image reconstruction and sparsity driven autofocus (SDA) [52].

4.3.1 Setup

In both synthetic and real scene experiments, we have considered various scenarios, i.e., a single data availability level (100%) and 7 different 1D phase error levels (0; $[-\pi/6,\pi/6]$; $[-\pi/4,\pi/4]$; $[-\pi/3,\pi/3]$; $[-3\pi/8,3\pi/8]$; $[-\pi/2,\pi/2]$; $[-2\pi/3,2\pi/3]$) in the phase history domain, and a single noise level ($\sigma_{\bf n}=0.1\sigma_{\bf y}$, where $\sigma_{\bf y}$ is the standard deviation of the magnitude of the phase history data), hence a total of 7 different scenarios. Note that the case of 0 phase error is equivalent to the scenario considered in Chapter 3.

The experimental procedure is the same as the procedure described in Chapter 3, i.e., same network architecture, same patch-based approach, same data augmentation procedure, etc., except that we trained a new network with the setting of 100% data availability, $\sigma_{\bf n}=0.1\sigma_{\bf y}$ noise, and $[-\pi/6,\pi/6]$ phase error, and also, the algorithm has additional steps for phase error estimation and correction, as well as model matrix update, as described in the previous section. We have only performed same scene experiments, using the training and test windows from the Wonsan image.

Algorithm 3: PnP-CNN-SAR-AF for joint image reconstruction and 1D phase error correction.

```
Require: \mathbf{y}, \mathbf{H}(\boldsymbol{\phi}^{(0)}), \epsilon, \mathcal{D}_{\sigma}(x), K, T.
     \mathbf{f}^{(0)} \leftarrow \mathbf{H}(\boldsymbol{\phi}^{(0)})^H \mathbf{y} {Conventional reconstruction}
    \mathbf{B}^{(0)} \leftarrow \operatorname{diag}\{\mathbf{f}_m^{(0)}\}\
     \boldsymbol{\theta}^{(0)} \leftarrow \angle \mathbf{f}^{(0)}
    \mathbf{h}^{(0)} \leftarrow \mathbf{f}_m^{(0)} \\ \mathbf{u}^{(0)} \leftarrow 0
     k \leftarrow 0
    n \leftarrow 0
while \frac{\left\|\mathbf{u}^{(k+1)} - \mathbf{u}^{(k)}\right\|}{\left\|\mathbf{u}^{(k)}\right\|} \ge \epsilon \text{ and } k \le K \text{ do}
          \tilde{\mathbf{f}}^{(k)} \leftarrow \mathbf{h}^{(k)} - \mathbf{u}^{(k)}
          \mathbf{f}_{m}^{[0]} \leftarrow \mathbf{f}_{m}^{(k)}
\mathbf{while} \frac{\|\mathbf{f}_{m}^{[t+1]} - \mathbf{f}_{m}^{[t]}\|}{\|\mathbf{f}_{m}^{[t]}\|} \ge 10^{-4} \text{ and } t \le T \text{ do}
                 Calculate \boldsymbol{\theta}^{(t)} using (4.6) {Phase alignment}
                 Solve for \mathbf{f}_{m}^{(t)} using (4.7) {Magnitude update}
           end while
           for Every cross-range position p = 1 \dots P do
                 Calculate \Delta \phi_p^{(k+1)} using (4.8) {Phase error estimation}
                 Calculate \phi_p^{(k+1)} using (4.11) {Phase error correction}
                 Calculate \mathbf{H}_{p}(\phi_{p}^{(k+1)}) using (4.13) {Model matrix update}
           end for
          \mathbf{f}_{m}^{(k+1)} \leftarrow \mathbf{f}_{m}^{[t]}
\tilde{\mathbf{h}}^{(k)} \leftarrow \mathbf{f}_{m}^{(k+1)} + \mathbf{u}^{(k)}
          \mathbf{h}^{(k+1)} \leftarrow \mathcal{D}_{\sigma}(\tilde{\mathbf{h}}^{(k)}) \text{ {Prior update}}
\mathbf{u}^{(k+1)} \leftarrow \mathbf{u}^{(k)} + \mathbf{f}_{m}^{(k+1)} - \mathbf{h}^{(k+1)} \text{ {Lagrange multiplier update}}
           k \leftarrow k + 1
     end while return \mathbf{f}_m^{(k+1)}
```

Table 4.1 Experimental settings.

	PnP-CNN-SAR Algorithm	PnP-CNN-SAR-AF Algorithm
Network trained without phase error	PnP	PnP-AF
Network trained with phase error	PnP-pe	PnP-pe-AF

Table 4.2 Average SNR and SSIM values for different phase error levels and for PnP, PnP-pe, PnP-AF, and PnP-pe-AF.

Phase Error	PnP		PnP-pe		PnP-AF		PnP-pe-AF	
T Hoose Error	SNR (dB)	SSIM	SNR (dB)	SSIM	SNR (dB)	SSIM	SNR (dB)	SSIM
0	40.03	0.998	26.83	0.977	26.66	0.978	26.85	0.977
$[-\pi/6, \pi/6]$	14.77	0.801	14.05	0.776	19.96	0.904	23.59	0.959
$[-\pi/4, \pi/4]$	11.39	0.671	10.59	0.642	19.88	0.901	24.09	0.963
$[-\pi/3, \pi/3]$	8.87	0.549	8.23	0.525	19.88	0.902	24.09	0.964
$[-3\pi/8, 3\pi/8]$	7.96	0.499	7.29	0.472	19.86	0.900	24.02	0.963
$[-\pi/2, \pi/2]$	5.66	0.346	5.07	0.325	19.86	0.901	24.22	0.964
$[-2\pi/3, 2\pi/3]$	3.73	0.190	3.26	0.176	19.74	0.840	24.04	0.977

4.3.2 Results

To show the effectiveness of the additional phase error estimation and correction steps as well as training the network with data containing phase errors, we first present results of four different experimental settings, namely, PnP, PnP-pe, PnP-AF, and PnP-pe-AF. In PnP and PnP-pe, the experimental procedure follows the steps in Algorithm 2, i.e., no phase error estimation and correction steps, and in PnP-AF and PnP-pe-AF, the steps given in Algorithm 3 are followed. Also, in PnP and PnP-AF, the network used is the same as the one that is used for the real scene experiments in Chapter 3, and hence trained without phase errors, while in PnPpe and PnP-pe-AF, the network trained with data containing phase errors is used. Table 4.1 summarizes these four settings. Figures 4.1–4.7 show the reconstruction results of these settings for various scenarios for Image 26 from the Wonsan test set, and Table 4.2 shows the average SNR and SSIM values obtained with different settings. These results suggest that the additional steps for phase error estimation and correction improve the performance significantly. Another observation is that the network trained with data containing phase errors only help if it is used within the PnP-CNN-SAR-AF framework.

Figures 4.8–4.14 and 4.15–4.21 show the reconstruction results of FFT-based reconstruction, SDA, and PnP-pe-AF, for various scenarios for *Image 26* and *Image 28* from the Wonsan test set, respectively. Table 4.3 show the quantitative comparison between these methods. These results show that PnP-pe-AF outperforms other methods in terms of both visual quality and performance metrics. It is also worth noting that SDA performs quite well in all scenarios for sparse scenes however its

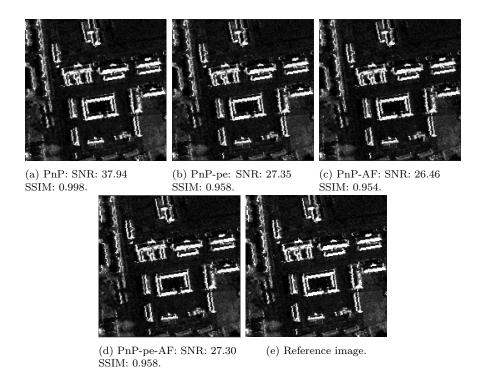


Figure 4.1 Reconstruction results for $Image\ 26$ of the Wonsan test set with no phase error.

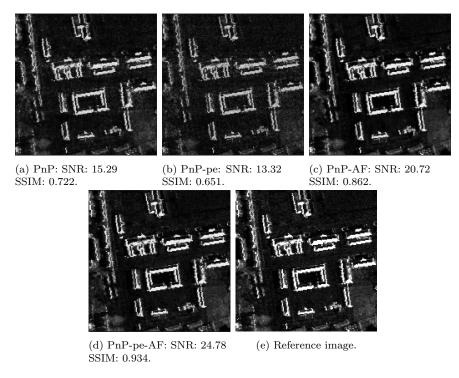


Figure 4.2 Reconstruction results for *Image 26* of the Wonsan test set with phase error in the range $[-\pi/6, \pi/6]$.

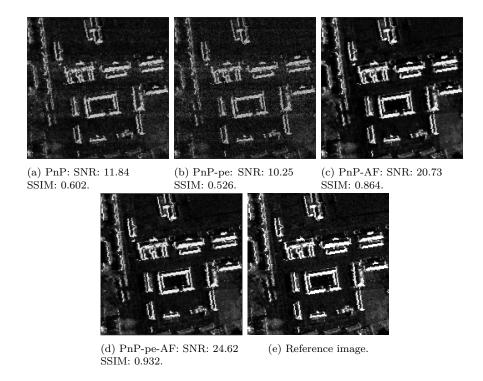


Figure 4.3 Reconstruction results for *Image 26* of the Wonsan test set with phase error in the range $[-\pi/4, \pi/4]$.

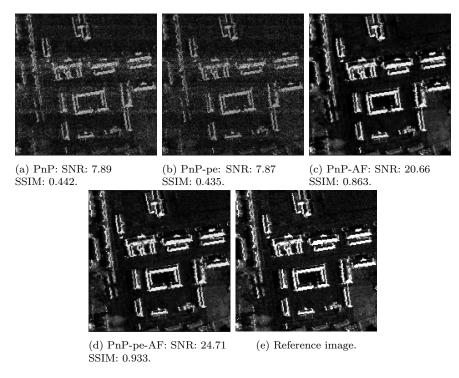


Figure 4.4 Reconstruction results for *Image 26* of the Wonsan test set with phase error in the range $[-\pi/3, \pi/3]$.

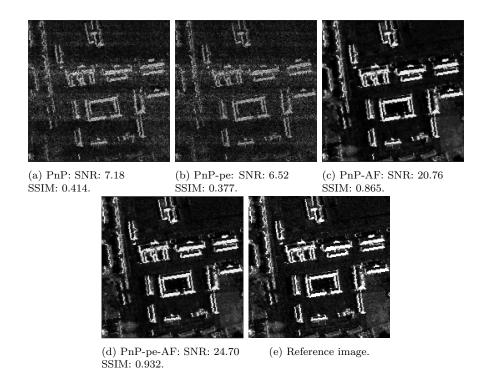


Figure 4.5 Reconstruction results for *Image 26* of the Wonsan test set with phase error in the range $[-3\pi/8, 3\pi/8]$.

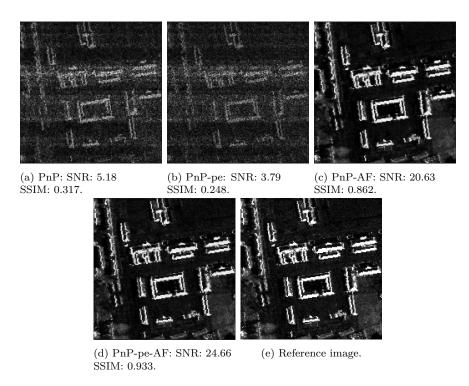


Figure 4.6 Reconstruction results for *Image 26* of the Wonsan test set with phase error in the range $[-\pi/2,\pi/2]$.

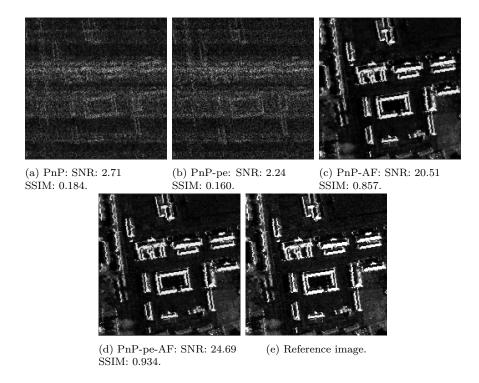


Figure 4.7 Reconstruction results for *Image 26* of the Wonsan test set with phase error in the range $[-2\pi/3, 2\pi/3]$.

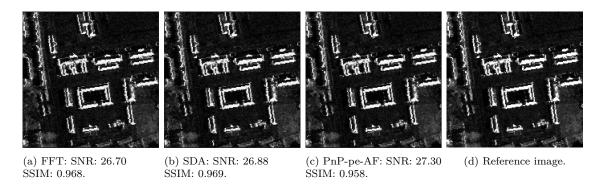


Figure 4.8 Visual comparison of reconstructions for *Image 26* of the Wonsan test set with no phase error.

performance degrades rapidly as the amount of phase error increases for non-sparse scenes. Performing these experiments for lower data availability scenarios would give more insight on the behaviors of the competing methods.

4.3.3 Summary

The preliminary results in this chapter show that joint image reconstruction and 1D phase error correction is indeed achievable with the PnP-CNN-SAR-AF algorithm. Our results suggest that both the phase error estimation and correction steps in the PnP-CNN-SAR-AF algorithm and training the network with phase errors improve the performance significantly. We have also observed that using a network trained

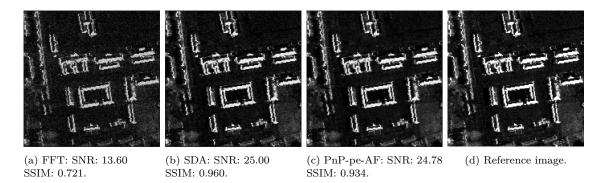


Figure 4.9 Visual comparison of reconstructions for *Image 26* of the Wonsan test set with phase error in the range $[-\pi/6, \pi/6]$.

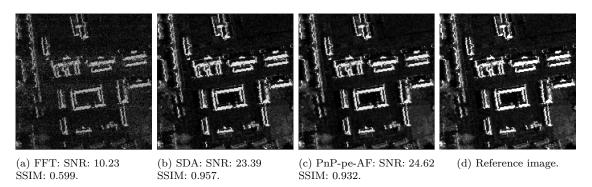


Figure 4.10 Visual comparison of reconstructions for *Image 26* of the Wonsan test set with phase error in the range $[-\pi/4, \pi/4]$.

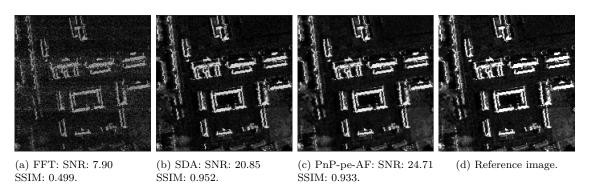


Figure 4.11 Visual comparison of reconstructions for *Image 26* of the Wonsan test set with phase error in the range $[-\pi/3, \pi/3]$.

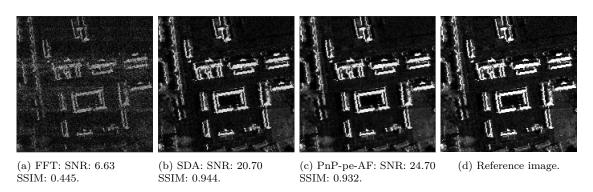


Figure 4.12 Visual comparison of reconstructions for *Image 26* of the Wonsan test set with phase error in the range $[-3\pi/8, 3\pi/8]$.

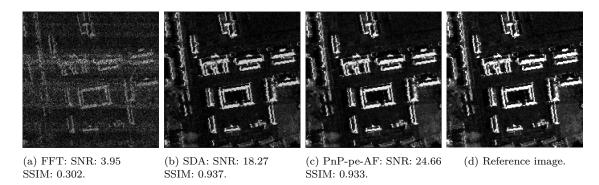


Figure 4.13 Visual comparison of reconstructions for *Image 26* of the Wonsan test set with phase error in the range $[-\pi/2, \pi/2]$.

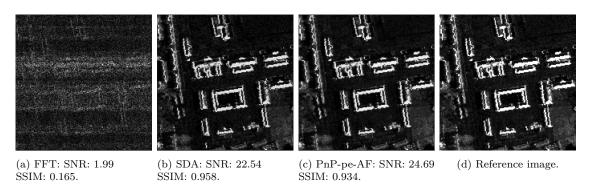


Figure 4.14 Visual comparison of reconstructions for *Image 26* of the Wonsan test set with phase error in the range $[-2\pi/3, 2\pi/3]$.

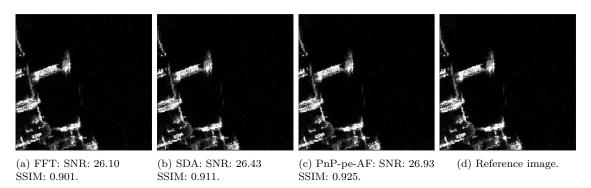


Figure 4.15 Visual comparison of reconstructions for $Image\ 28$ of the Wonsan test set with no phase error.

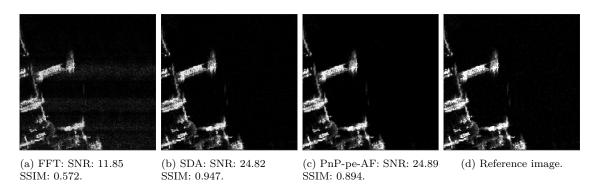


Figure 4.16 Visual comparison of reconstructions for *Image 28* of the Wonsan test set with phase error in the range $[-\pi/6, \pi/6]$.

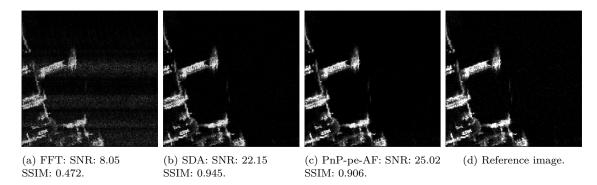


Figure 4.17 Visual comparison of reconstructions for *Image 28* of the Wonsan test set with phase error in the range $[-\pi/4, \pi/4]$.

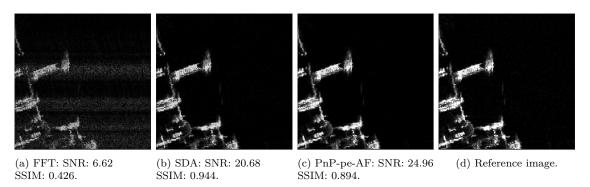


Figure 4.18 Visual comparison of reconstructions for *Image 28* of the Wonsan test set with phase error in the range $[-\pi/3, \pi/3]$.

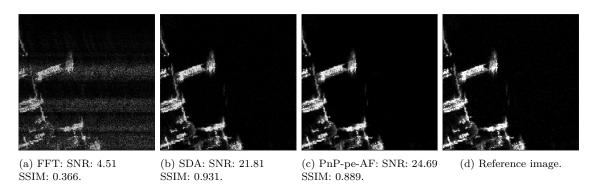


Figure 4.19 Visual comparison of reconstructions for *Image 28* of the Wonsan test set with phase error in the range $[-3\pi/8, 3\pi/8]$.

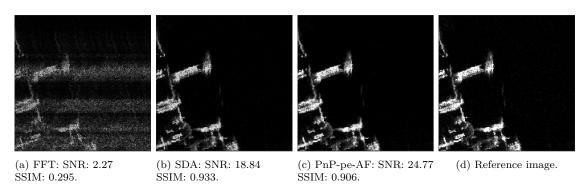
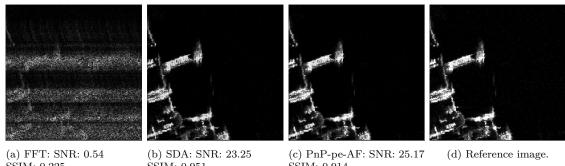


Figure 4.20 Visual comparison of reconstructions for *Image 28* of the Wonsan test set with phase error in the range $[-\pi/2,\pi/2]$.



SSIM: 0.225.

SSIM: 0.951.

SSIM: 0.914.

Figure 4.21 Visual comparison of reconstructions for *Image 28* of the Wonsan test set with phase error in the range $[-2\pi/3, 2\pi/3]$.

Table 4.3 Average SNR and SSIM values for different phase error levels and for FFT-based, SDA, and PnP-pe-AF methods.

Phase Error	FFT	ı	SDA	-	PnP-pe-AF		
Thase Error	SNR (dB)	SSIM	SNR (dB)	SSIM	SNR (dB)	SSIM	
0	26.64	0.973	26.71	0.974	26.85	0.977	
$[-\pi/6, \pi/6]$	13.26	0.750	19.89	0.875	23.59	0.959	
$[-\pi/4, \pi/4]$	9.99	0.618	18.62	0.851	24.09	0.963	
$[-\pi/3, \pi/3]$	7.76	0.502	16.90	0.803	24.09	0.964	
$[-3\pi/8, 3\pi/8]$	6.74	0.445	16.10	0.803	24.02	0.963	
$[-\pi/2, \pi/2]$	4.75	0.314	14.52	0.764	24.22	0.964	
$[-2\pi/3, 2\pi/3]$	3.04	0.171	12.05	0.676	24.04	0.977	

with phase error within the PnP-CNN-SAR framework does not generate good results, and the benefit of training the network with phase error can only be observed if it is used within the PnP-CNN-SAR-AF framework.

Chapter 5

SAR ATR in the Phase History Domain

In the previous two chapters, we have focused on the problems of image reconstruction and phase error correction in SAR imaging. In this chapter, we shift our focus towards another important SAR imaging problem, i.e., automatic target recognition (ATR), and we investigate the idea of performing the ATR task in the phase history domain. This is a joint work with Sara Atito, another Ph.D. student in our group.

5.1 Phase History Domain ATR Methodologies

In this section, we present two frameworks for SAR ATR that work in the phase history domain and utilize CNNs for learning and classification.

5.1.1 Phase History Domain Classification (PHDC)

Since the conventional reconstruction for SAR is performed using noisy measurements, and because SAR is a coherent imaging modality, formed images suffer from speckle which also affects the ATR performance. Note that only the magnitudes of the formed images are used in ATR since the phase does not contain much information about the spatial structure of the 2D projection of the scene produced by conventional SAR imaging. Moreover, we know from an information theoretic perspective that, image formation does not provide us any additional information about the scene, it is merely a visualization. Since ATR is performed by computers, in principle it does not really matter if we feed the ATR system with formed images or phase history data. In either case, the system would learn the dependency patterns in the data. Since the information contained in the formed images cannot

be more than the information in the phase histories, we predicted that, in an ideal scenario, using the phase history data instead of formed images may result in equal or better ATR performance. This is something we explore in this study. Hence, our first proposed framework uses phase history data as an input to the ATR system. To evaluate our framework, we used the MSTAR data set and two well-known CNN architectures, AlexNet and VGG16. AlexNet and VGG16 are both originally designed for image classification tasks and they require 3-channel real input data. However, the phase history data we have is complex. Therefore, we needed a way to feed the data to the networks, and using the magnitudes, real parts, and imaginary parts of the phase histories as the three channel inputs gave us the best result. Specifically, we assign the magnitudes, real parts, and imaginary parts of the phase histories of the images in the MSTAR data set as the red, green, and blue channels, respectively, and form an RGB image for each data point in the data set.

5.1.2 Image-Phase Removed Phase History Classification (IPRPHC)

Since the phase of the reconstructed image does not really contain any information about the observed scene, as it is highly dominated by speckle, it is usually discarded in image analysis. Motivated by this, our second framework performs the following steps: image formation, image-phase removal, and phase history data generation. We evaluate our framework in the same way described in Section 5.1.1.

5.2 Experimental Results

5.2.1 Setup

We now evaluate the performance of the ATR methodologies discussed in Section 5.1 using the MSTAR data set. The data set consists of 10 classes, i.e., tanks (T62, T72), armored vehicles (BRDM2, BMP2, BTR60, BTR70), a rocket launcher (2S1), an air defense unit (ZSU234), a military truck (ZIL131), and a bulldozer (D7). The MSTAR data set only provides formed images. Therefore, in order to obtain the phase histories, we undo the final steps of the MSTAR image formation using the method described in Refs. [1,191]. Also, we resize the generated images to 227×227 for AlexNet and to 224×224 for VGG16 to match their respective input sizes. Figure 5.1 shows sample data of each class for PHDC experiments and Figure 5.2 shows sample data from the same chips for IPRPHC experiments. In our experiments, we use the images at 17° depression angle for training and images at 15° depression angle for testing. Since CNNs require large number of images

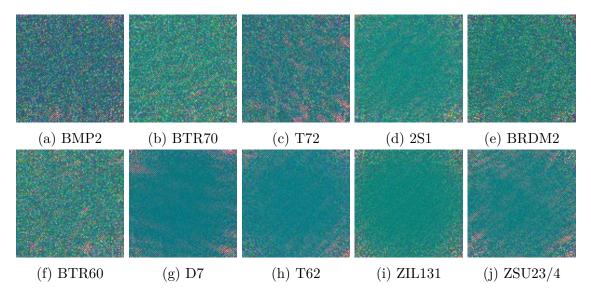


Figure 5.1 Sample data of each class for PHDC experiments.

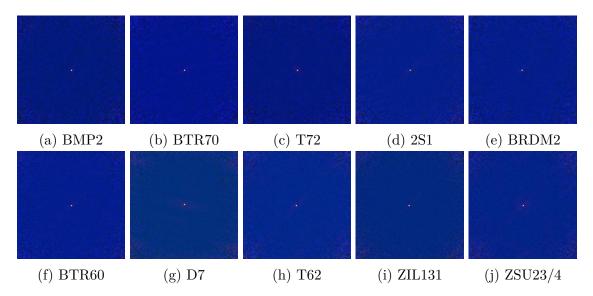


Figure 5.2 Sample data of each class for IPRPHC experiments.

for training, we augment our training set by rotating each image by $[2^{\circ}, 4^{\circ}, 6^{\circ}, 8^{\circ}, 10^{\circ}]$ clockwise and counter-clockwise. Table 5.1 shows the numbers of each type of target images in each of these sets before augmentation. All of the experiments are performed on MATLAB R2018a using NVIDIA TITAN Xp GPU.

In the following subsections, we present the results of phase history domain classification and image-phase removed phase history classification experiments and conclude this section with comparison of the results of other methods and a summary.

5.2.2 PHDC Results

We now present the recognition performance of the phase history domain classification using AlexNet and VGG16 architectures. Tables 5.2 and 5.3 show the results for

Table 5.1 Composition of the MSTAR data set.

	Trai	ning set	Test set				
Target	Depression	No. of images	Depression	No. of images			
BMP2	17°	233	15°	195			
BTR70	17°	233	15°	196			
T72	17°	232	15°	196			
2S1	17°	299	15°	274			
BRDM2	17°	298	15°	274			
BTR60	17°	256	15°	195			
D7	17°	299	15°	274			
T62	17°	299	15°	273			
ZIL131	17°	299	15°	274			
ZSU23/4	17°	299	15°	274			

Table 5.2 Confusion matrix for PHDC results with the AlexNet architecture.

			Predicted Class										
		BMP2	BTR70	T72	2S1	BRDM2	BTR60	D7	T62	ZIL131	ZSU23/4	Accuracy	
	BMP2	103	8	35	3	0	29	2	3	10	2	52,82%	
	BTR70	29	118	1	6	0	29	0	1	12	0	60,20%	
	T72	7	0	162	1	0	8	3	14	0	1	$82,\!65\%$	
SS	2S1	0	0	0	272	0	0	0	1	0	1	99,27%	
Class	BRDM2	3	0	7	0	257	3	0	0	0	4	$93,\!80\%$	
ne	BTR60	11	19	37	3	0	109	1	6	9	0	55,90%	
True	D7	0	0	0	0	0	0	257	4	1	12	$93,\!80\%$	
	T62	0	0	0	0	0	0	37	224	3	9	82,05%	
	ZIL131	0	0	0	0	0	0	3	0	271	0	98,91%	
	ZSU23/4	1	0	0	0	0	0	0	0	0	273	$99{,}64\%$	
						Overall	Accuracy					84,37%	

AlexNet and VGG16, respectively. The results show that tanks and armored vehicles are harder to distinguish. The overall performance is worse than the competing methods for both architectures.

5.2.3 IPRPHC Results

We now present the recognition performance of the image-phase removed phase history classification using AlexNet and VGG16 architectures. Tables 5.4 and 5.5 show the results for AlexNet and VGG16, respectively. The results show that this methodology works well and the performance is at the same level with the state-of-the-art techniques. Note that, with AlexNet, 7 of the classes, and with VGG16, 5 of the classes reached 100% accuracy.

Table 5.3 Confusion matrix for PHDC results with the VGG16 architecture.

			Predicted Class										
		BMP2	BTR70	T72	2S1	BRDM2	BTR60	D7	T62	ZIL131	ZSU23/4	Accuracy	
	BMP2	93	18	22	3	4	42	2	6	3	2	47,69%	
	BTR70	20	137	0	5	6	17	0	1	10	0	69,90%	
	T72	12	1	140	0	3	22	0	15	0	3	71,43%	
SS	2S1	0	0	0	272	0	0	0	0	0	2	99,27%	
Class	BRDM2	7	3	3	0	255	6	0	0	0	0	93,07%	
je (BTR60	23	29	30	2	3	92	0	8	8	0	47,18%	
True	D7	0	0	0	0	0	0	247	16	2	9	90,15%	
	T62	0	0	0	1	0	0	19	241	1	11	88,28%	
	ZIL131	0	0	0	0	0	0	4	2	268	0	97,81%	
	ZSU23/4	0	0	0	0	1	0	1	0	0	272	$99{,}27\%$	
						Overall	Accuracy					83,18%	

Table 5.4 Confusion matrix for IPRPHC results with the AlexNet architecture.

						Predict	ed Class					Class
		BMP2	BTR70	T72	2S1	BRDM2	BTR60	D7	T62	ZIL131	ZSU23/4	Accuracy
	BMP2	185	1	7	0	0	2	0	0	0	0	94,87%
	BTR70	4	189	0	0	0	3	0	0	0	0	96,43%
	T72	0	0	196	0	0	0	0	0	0	0	100,00%
$_{\rm SS}$	2S1	0	0	0	274	0	0	0	0	0	0	100,00%
Class	BRDM2	0	0	0	0	274	0	0	0	0	0	100,00%
_	BTR60	1	1	0	0	0	189	4	0	0	0	96,92%
True	D7	0	0	0	0	0	0	274	0	0	0	100,00%
	T62	0	0	0	0	0	0	0	273	0	0	100,00%
	ZIL131	0	0	0	0	0	0	0	0	274	0	100,00%
	ZSU23/4	0	0	0	0	0	0	0	0	0	274	$100,\!00\%$
						Overall	Accuracy					99,05%

Table 5.5 Confusion matrix for IPRPHC results with the VGG16 architecture.

			Predicted Class									
		BMP2	BTR70	T72	2S1	BRDM2	BTR60	D7	T62	ZIL131	ZSU23/4	Accuracy
	BMP2	182	2	7	0	0	4	0	0	0	0	93,33%
	BTR70	6	188	0	0	0	2	0	0	0	0	95,92%
	T72	0	0	195	0	0	1	0	0	0	0	99,49%
SS	2S1	3	0	0	271	0	0	0	0	0	0	98,91%
Class	BRDM2	0	0	0	0	274	0	0	0	0	0	100,00%
Je (BTR60	0	1	0	0	0	194	0	0	0	0	99,49%
True	D7	0	0	0	0	0	0	274	0	0	0	100,00%
	T62	0	0	0	0	0	0	0	273	0	0	100,00%
	ZIL131	0	0	0	0	0	0	0	0	274	0	100,00%
	ZSU23/4	0	0	0	0	0	0	0	0	0	274	$100,\!00\%$
						Overall	Accuracy					98,93%

Table 5.6 Confusion matrix for conventional reconstruction results with the AlexNet architecture.

						Predict	ed Class					Class
		BMP2	BTR70	T72	2S1	BRDM2	BTR60	D7	T62	ZIL131	ZSU23/4	Accuracy
	BMP2	191	0	2	0	1	1	0	0	0	0	97,95%
	BTR70	0	195	0	0	1	0	0	0	0	0	99,49%
	T72	0	0	196	0	0	0	0	0	0	0	100,00%
\mathbf{S}	2S1	0	0	0	273	1	0	0	0	0	0	99,64%
Class	BRDM2	4	1	0	0	268	1	0	0	0	0	97,81%
	BTR60	1	1	1	0	2	188	0	0	2	0	96,41%
True	D7	0	0	0	2	0	0	271	0	0	1	98,91%
	T62	1	0	0	0	0	0	0	265	0	7	97,07%
	ZIL131	0	0	0	0	0	0	2	0	272	0	$99,\!27\%$
	ZSU23/4	0	0	0	0	0	0	0	0	0	274	$100,\!00\%$
						Overall	Accuracy					98,68%

Table 5.7 Confusion matrix for conventional reconstruction results with the VGG16 architecture.

						Predict	ed Class					Class
		BMP2	BTR70	T72	2S1	BRDM2	BTR60	D7	T62	ZIL131	ZSU23/4	Accuracy
	BMP2	186	1	5	0	1	2	0	0	0	0	95,38%
	BTR70	0	196	0	0	0	0	0	0	0	0	100,00%
	T72	2	0	194	0	0	0	0	0	0	0	98,98%
SS	2S1	5	0	1	264	3	0	0	0	0	1	96,35%
Class	BRDM2	0	0	0	0	274	0	0	0	0	0	100,00%
	BTR60	0	1	1	0	4	189	0	0	0	0	96,92%
True	D7	0	0	0	0	0	0	271	1	2	0	98,91%
	T62	0	0	0	0	0	0	0	265	1	7	97,07%
	ZIL131	0	0	0	0	0	0	0	0	274	0	100,00%
	ZSU23/4	1	0	0	0	0	0	0	0	0	273	99,64%
						Overall	Accuracy					98,39%

5.2.4 Comparison & Summary

We now compare our results with existing ATR methodologies. First, we train AlexNet and VGG16 with the magnitude of the original MSTAR chips. For this, we feed the magnitude of the chips to all 3 channels of the input layers of both architectures. Tables 5.6 and 5.7 show the results for these networks, respectively.

We also compare our results with the methods in Refs. [117,119,120,123,125,126]. Note that, in the MSTAR dataset, there are additional images of different variations of two of the vehicles, namely, T72 and BMP2, in addition to the ones mentioned in Table 5.1, and [119,120,123] use these variants in the test set while [126] uses the variants in both training and test sets. Hence, it should be noted that the results of different methods compared here are not based on exactly the same training and test sets. Table 5.8 shows the comparison of the proposed methodologies with the existing methods.

Table 5.8 Comparison of the proposed frameworks with existing methods.

Method	Accuracy
Image Domain (AlexNet)	98,68%
Image Domain (VGG16)	$98,\!39\%$
PHDC (AlexNet)	$84,\!37\%$
PHDC (VGG16)	$83,\!18\%$
IPRPHC (AlexNet)	$99,\!05\%$
IPRPHC (VGG16)	98,93%
Song et. al. [119]	96,24%
Sun et. al. [120]	$91,\!48\%$
Chen and Wang [117]	84,7%
Chen et. al. [123]	$99,\!42\%$
El Housseini et. al. [125]	$92,\!63\%$
Wagner et. al. [126]	$98,\!47\%$

We observe that our first framework performed poorly compared to existing work while our second framework performed well, only surpassed by the method in Ref. 123. For our second framework, it seems like it was a reasonable judgment to remove the phase of the MSTAR chips before returning back to the phase history domain.

Another observation is that we reached slightly better results with AlexNet architecture compared to VGG16 for all types of experiments. The reason for this might be that VGG16 is a larger network which needs more data to train.

Chapter 6

Conclusions & Future Directions

6.1 Summary & Conclusions

In this dissertation, we have contributed to the SAR imaging literature. We have developed three frameworks that enhance the performance of various SAR imaging tasks. In particular, we have developed a SAR image formation framework that utilizes deep learning-based priors using the plug-and-play (PnP) priors [33] framework. We have also extended our framework to address phase errors along with image formation. Finally, we have proposed a SAR ATR framework that works in the data domain.

In Chapter 3, our image formation framework, namely PnP-CNN-SAR, was established. In this framework, the forward model and the prior model are decoupled using the PnP priors [33] framework. Within our framework, we have used a Fourier transform-based forward model and a CNN-based prior model. We conducted experiments with different settings to evaluate our proposed framework, i.e., we tested our framework in the presence of different levels of noise and data availability. We compared the results with three existing reconstruction methods, namely, FFT-based, NQR-based [1], and DL-based [3] reconstruction, and a PnP framework-based method with BM3D regularizer, using signal-to-noise ratio (SNR) and structural similarity (SSIM) index metrics. We have shown that PnP-CNN-SAR outperforms other methods in terms of visual quality. Overall, this study suggests that deep learning methods may have the potential to learn complicated spatial patterns and enable their incorporation as priors into computational radar imaging.

In Chapter 4, the extension of PnP-CNN-SAR, namely PnP-CNN-SAR-AF, to jointly address image formation and phase error correction, is presented. We have tested our framework in the presence of different levels of phase errors and data

availability. We have first shown the effectiveness of the additional phase error estimation and correction steps as well as training the network with phase errors, by designing four different experimental settings, namely, PnP, PnP-pe, PnP-AF, and PnP-pe-AF. In PnP and PnP-AF, the network trained without phase error is used while in PnP-pe and PnP-pe-AF, the network trained with phase error is used. Also, in PnP and PnP-pe, the PnP-CNN-SAR algorithm is used while in PnP-AF and PnP-pe-AF, the PnP-CNN-SAR-AF algorithm is used which has additional phase error estimation and correction steps. The results obtained in these settings have shown that the additional phase error estimation and correction steps and training the network with phase errors improve the performance significantly. These results have also shown that training the network with phase error is not enough by itself and this network should be used within the PnP-CNN-SAR-AF algorithm. Later, we compared the results of PnP-pe-AF with two existing methods, namely, FFTbased image reconstruction and SDA [52], using SNR and SSIM. We have shown that PnP-CNN-SAR-AF outperforms other methods in all scenarios in terms of both visual quality and the performance metrics.

In Chapter 5, we have introduced a SAR ATR framework that performs the recognition task in the phase history domain. We have designed two ATR modalities, namely, PHDC and IPRPHC. PHDC performs the ATR task using the collected phase history data directly while in IPRPHC, image formation, image phase removal, and phase history generation steps are performed before the ATR task. Within this framework, we have used two well-known CNNs, AlexNet [35] and VGG16 [36], which are designed for classification tasks. While the performance with the PHDC was subpar, we have reached near state-of-the-art performance levels with the IPRPHC modality. Nevertheless, presented results suggest that SAR ATR without image formation or in the phase histories domain can be an important research direction.

Overall, this dissertation brought a new point-of-view to SAR imaging problems, and has shown that deep neural networks can be effectively used in improving the performance of SAR imaging tasks. Although SAR ATR and other classification/detection/segmentation applications have already started benefiting from the deep learning tools, most of these works are essentially either postprocessing conventionally reconstructed images or feature extraction from those. The works presented in this dissertation show that deep neural networks can have a much broader role than just postprocessing within SAR imaging tasks, and can improve the overall performance significantly.

6.2 Topics for Future Research

All three frameworks presented in this dissertation can benefit from further analysis and extensions. In the following subsections, we will focus each of them individually, and present our suggestions.

6.2.1 Further Analysis and Extensions on PnP-CNN-SAR for Image Reconstruction

The PnP-CNN-SAR framework presented in Chapter 3 is open to further extensions and developments in many ways. First of all, in synthetic scene experiments, we have observed that a network trained for a specific scenario might not be the one that performs the best for that particular scenario, rather a network trained with a slightly higher data availability tends to perform better. A deeper analysis to understand the underlying reasons and an extension to the real scene case would be quite helpful. We have also observed that quantitative analyses based on SNR and SSIM do not necessarily harmonize well with qualitative assessments, especially in lower data availability scenarios, i.e., a reconstructed image might have a better perceptual quality compared to another one, while having a worse score in terms of the metrics used. (See, for example, last two rows of Figures 3.5, 3.6, 3.13, 3.14, 3.16, 3.17. In many of these cases, the NQR-based [1] reconstruction achieves the best performance in terms of SNR, while producing least useful reconstructions.) It is clear that new metrics that are more consistent with the visual quality of the reconstructed images are required. Lastly, we observed that the most time consuming process in our algorithm is the phase alignment step in the model update which can take up to 75% of the overall computation time. A more efficient algorithm for the phase alignment step would be useful for time-critical reconstruction tasks.

6.2.2 Further Analysis and Extensions on PnP-CNN-SAR-AF for Joint Image Reconstruction and 1D Phase Error Correction

In Chapter 4, we have used the PnP-CNN-SAR-AF framework to jointly address image reconstruction and 1D phase errors. This framework can easily be extended to address 2D separable and non-separable phase errors as well. Also, in this work, we have only investigated the effects of the amount of phase error in the case of full data availability and relatively low noise. The performance of the proposed framework can be evaluated in more challenging scenarios, i.e., in cases of lower data availability and higher noise levels. Finally, in this work, we have trained the

network with data generated with a particular scenario in mind. A network trained with data generated with various scenarios, i.e., different combinations of different phase error, data availability, and noise levels, would likely have more generalization power, and could give good results in a variety of scenarios.

6.2.3 Further Analysis and Extensions on SAR ATR in the Phase History Domain

The work presented in Chapter 5 can be improved in several ways. We have used an approximation to the phase histories in our experiments for the PHDC. Using the actual phase history data would potentially give better and more realistic results. Also, the CNN architectures we have used in the experiments are relatively old and smaller compared to the recent networks. We can benefit from using a deeper and/or task specific network. Note however that using a deeper network might require a larger training set and longer time to train. Also, the effect of the way the data are used in the CNN can be investigated. Since AlexNet and VGG16 are both designed for image classification, they require 3-channel real input data. However, the phase history data we have is complex. Therefore, we decided to use real and imaginary parts and the magnitude of the phase histories to feed to the 3 channels since this gave us the best result among other methods (using only one part for all channels or another combination like real, imaginary, and phase). But we are still not sure if this is the best way to use the data and that is open for further investigation. One can also use a CNN that is designed to work with complex data like the one proposed by Gao et. al. [196]. In addition, using different augmentation techniques and generating larger training sets might improve the results. Finally, in this work, we only considered the standard operation conditions where the depression angles for the training and test sets are fairly similar (17° vs. 15°). The performance of the proposed modalities can also be investigated for the extended operation conditions, i.e., where depression angles differ by a large margin (Training 17° - Test 30° and Training 17° - Test 45°).

Appendix A

Additional Results from PnP-CNN-SAR for Image Reconstruction

This appendix provides additional results and information in support of Chapter 3. The rest of this supplement is structured as follows. Sections A.1, A.2, and A.3 provide detailed quantitative and qualitative analyses for the synthetic scene experiments, same scene experiments, and different scene experiments, respectively. The results presented in this supplement show that PnP-CNN-SAR outperforms other methods in both quantitative metrics and reconstruction quality, in the majority of the scenarios considered.

A.1 Synthetic Scene Experiments

A.1.1 Experiment 1: 100% Data Availability and $\sigma_n = 0.1\sigma_y$

Table A.1 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR for all 9 images, and best SSIM for 7 images. Fig. A.1 shows the reconstruction results from all methods for this experiment. NQR-based and PnP-CNN-SAR results in near perfect reconstructions. Other methods also produce good reconstructions.

A.1.2 Experiment 2: 90% Data Availability and $\sigma_n = 0.1\sigma_y$

Table A.2 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR for 4 images, and best SSIM for 3 images. Fig. A.2 shows the

Table A.1 SNR and SSIM values for synthetic SAR scenes for the case of 100% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			SNR (dE	3)				SSIM		
	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR
mage 1	26.25	25.44	27.12	28.08	36.88	0.930	0.978	0.901	0.924	0.998
mage 2	23.72	31.15	24.09	24.69	36.02	0.691	0.999	0.596	0.599	0.989
mage 3	25.76	31.99	26.54	27.37	37.61	0.928	0.989	0.894	0.927	0.998
mage 4	26.01	24.66	27.06	28.37	34.29	0.956	0.976	0.940	0.959	0.996
mage 5 mage 6	26.56	25.51	27.30 25.68	28.68	39.64	0.954	0.959	0.936 0.945	0.952	0.999 0.997
mage 6 mage 7	25.04 25.38	24.94 24.58	26.22	25.88 26.29	38.79 38.30	0.957 0.963	0.989 0.989	0.945	0.945 0.957	0.999
mage 8	25.53	24.91	25.98	26.55	38.97	0.966	0.988	0.954	0.962	0.999
mage 9	25.72	32.61	26.85	28.74	40.79	0.808	0.996	0.770	0.827	0.992
verage	25.55	27.31	26.31	27.18	37.92	0.906	0.985	0.877	0.895	0.996
	(a) F	FT-base	ed		(b) NQR-	based		(c)	DL-bas	ed
		•				# 1			- +	

Figure A.1 Reconstruction results for the synthetic scenes for the case of 100% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$.

(e) PnP-CNN-SAR

(f) Ground Truth

(d) PnP-BM3D

reconstruction results from all methods for this experiment. DL-based and PnP-CNN-SAR generally produces better reconstructions compared to other methods.

Table A.2 SNR and SSIM values for synthetic SAR scenes for the case of 90% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			SNR (dE	3)		SSIM					
	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR	
Image 1	11.84	17.85	27.05	26.57	23.26	0.457	0.867	0.910	0.942	0.922	
Image 2	10.12	18.78	20.98	19.78	20.83	0.377	0.991	0.540	0.520	0.967	
Image 3	11.62	27.05	25.23	25.17	23.94	0.494	0.950	0.894	0.940	0.917	
Image 4	11.38	9.27	25.08	27.07	23.32	0.627	0.533	0.934	0.968	0.960	
Image 5	11.70	13.72	26.90	27.31	24.63	0.519	0.627	0.941	0.955	0.938	
Image 6	10.61	18.10	23.20	17.39	28.38	0.683	0.944	0.934	0.880	0.985	
Image 7	11.41	18.58	24.56	20.18	26.14	0.668	0.949	0.950	0.934	0.976	
Image 8	11.18	13.25	24.11	21.05	28.88	0.694	0.810	0.950	0.946	0.987	
Image 9	11.71	23.94	25.73	28.43	28.56	0.333	0.934	0.766	0.870	0.932	
Average	11.29	17.84	24.76	23.66	25.33	0.539	0.845	0.869	0.884	0.954	

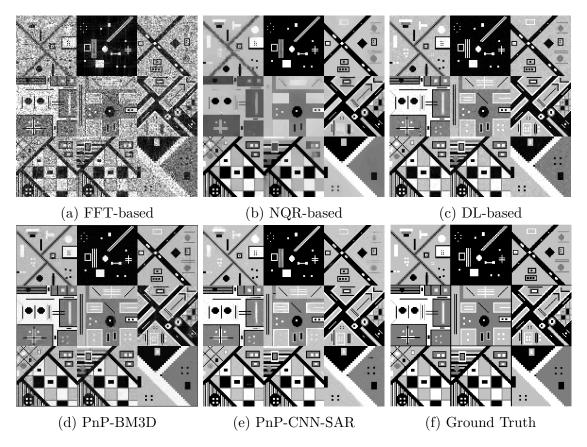


Figure A.2 Reconstruction results for the synthetic scenes for the case of 90% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$.

A.1.3 Experiment 3: 80% Data Availability and $\sigma_n = 0.1\sigma_y$

Table A.3 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR for 4 images, and best SSIM for 5 images. Fig. A.3 shows the reconstruction results from all methods for this experiment. DL-based and PnP-CNN-SAR generally produces better reconstructions compared to other methods.

Table A.3 SNR and SSIM values for synthetic SAR scenes for the case of 80% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			SNR (dE	3)		SSIM					
	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR	
Image 1	10.38	15.29	21.97	21.86	20.31	0.373	0.748	0.859	0.847	0.863	
Image 2	7.97	8.85	20.65	8.02	21.09	0.360	0.927	0.561	0.406	0.972	
Image 3	9.67	22.48	26.13	16.43	21.32	0.422	0.913	0.910	0.771	0.845	
Image 4	9.77	10.43	23.14	20.75	19.11	0.552	0.525	0.925	0.884	0.913	
Image 5	10.39	12.87	22.09	19.22	20.13	0.430	0.652	0.895	0.837	0.853	
Image 6	9.17	16.47	20.65	11.30	24.14	0.628	0.913	0.918	0.738	0.970	
Image 7	9.82	13.11	20.62	13.03	24.28	0.608	0.769	0.917	0.766	0.968	
Image 8	9.47	13.84	21.42	13.58	24.08	0.613	0.850	0.936	0.817	0.980	
Image 9	10.25	26.61	23.16	21.43	26.29	0.278	0.978	0.743	0.664	0.921	
Average	9.65	15.55	22.20	16.18	22.30	0.474	0.809	0.852	0.748	0.921	

A.1.4 Experiment 4: 70% Data Availability and $\sigma_n = 0.1\sigma_y$

Table A.4 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR for 8 images, and best SSIM for 8 images. Fig. A.4 shows

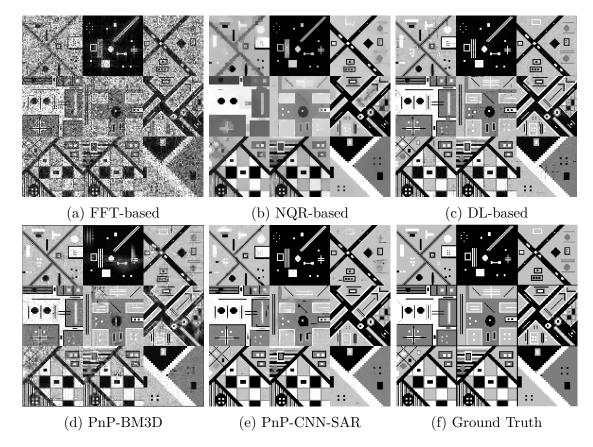


Figure A.3 Reconstruction results for the synthetic scenes for the case of 80% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$.

the reconstruction results from all methods for this experiment. PnP-CNN-SAR produces better reconstructions compared to other methods.

Table A.4 SNR and SSIM values for synthetic SAR scenes for the case of 70% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			SNR (dE	3)				SSIM		
	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR
Image 1	8.44	11.92	11.90	13.98	14.82	0.284	0.487	0.482	0.560	0.659
Image 2	6.52	7.67	17.84	2.90	15.27	0.322	0.896	0.536	0.208	0.963
Image 3	8.16	12.29	11.12	11.24	16.47	0.355	0.602	0.485	0.492	0.731
Image 4	8.32	9.21	11.76	13.32	17.44	0.434	0.377	0.613	0.688	0.860
Image 5	8.49	10.76	11.50	13.70	17.79	0.324	0.429	0.526	0.576	0.743
Image 6	7.58	8.72	11.22	6.73	18.21	0.529	0.570	0.714	0.421	0.926
Image 7	7.96	10.74	11.56	8.78	18.75	0.500	0.642	0.697	0.506	0.921
Image 8	7.84	8.87	10.81	8.95	18.42	0.529	0.571	0.675	0.572	0.935
${\rm Image}\ 9$	8.11	17.27	12.24	15.69	17.38	0.233	0.789	0.360	0.510	0.739
Average	7.94	10.83	12.22	10.59	17.17	0.390	0.596	0.566	0.504	0.831

A.1.5 Experiment 5: 50% Data Availability and $\sigma_n = 0.1\sigma_v$

Table A.5 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR for all 9 images, and best SSIM for 8 images. Fig. A.5 shows the reconstruction results from all methods for this experiment. PnP-CNN-SAR produces better reconstructions compared to other methods.

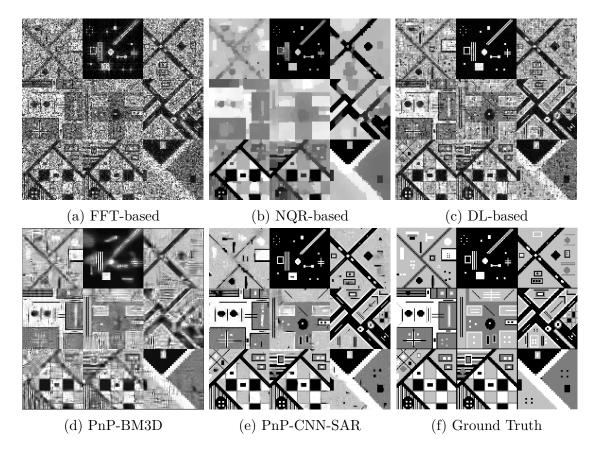


Figure A.4 Reconstruction results for the synthetic scenes for the case of 70% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$.

Table A.5 SNR and SSIM values for synthetic SAR scenes for the case of 50% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			SNR (dE	3)				SSIM		
	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR
Image 1	6.44	10.83	8.14	10.60	11.68	0.222	0.268	0.277	0.302	0.400
Image 2	4.39	6.03	6.08	2.49	8.49	0.280	0.839	0.346	0.141	0.887
Image 3	6.29	9.64	7.60	8.54	11.79	0.244	0.414	0.287	0.230	0.550
Image 4	6.26	8.22	7.40	8.76	9.83	0.297	0.265	0.344	0.398	0.472
Image 5	6.50	10.76	8.09	10.74	11.82	0.248	0.348	0.291	0.319	0.388
Image 6	5.53	6.97	6.69	5.97	10.86	0.364	0.414	0.429	0.267	0.714
Image 7	6.01	7.74	7.38	7.52	11.03	0.371	0.363	0.433	0.342	0.659
Image 8	6.11	7.05	7.19	7.26	10.64	0.394	0.318	0.460	0.369	0.689
Image 9	6.49	13.49	8.35	11.04	14.95	0.177	0.657	0.229	0.347	0.607
Average	6.00	8.97	7.44	8.10	11.23	0.289	0.432	0.344	0.302	0.596

A.1.6 Experiment 6: 30% Data Availability and $\sigma_n = 0.1\sigma_y$

Table A.6 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR for 5 images, and best SSIM for 6 images. Fig. A.6 shows the reconstruction results from all methods for this experiment. None of the methods are able to produce acceptable reconstructions for most of the images.

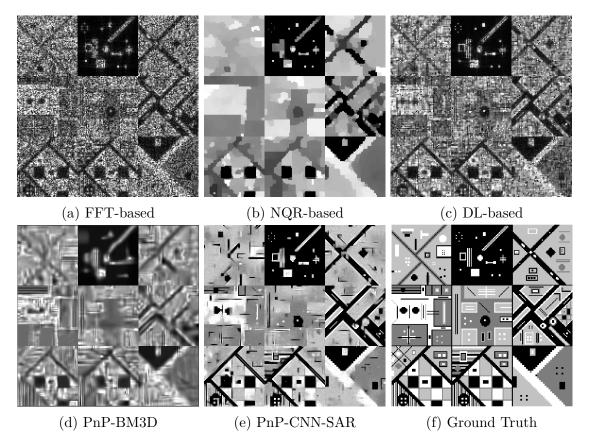


Figure A.5 Reconstruction results for the synthetic scenes for the case of 50% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$.

Table A.6 SNR and SSIM values for synthetic SAR scenes for the case of 30% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			SNR (dE	3)				SSIM		
	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR
Image 1	4.84	10.06	6.95	6.92	9.54	0.156	0.223	0.167	0.203	0.237
Image 2	3.14	4.08	3.17	2.60	5.21	0.215	0.704	0.200	0.122	0.766
Image 3	4.72	7.87	6.21	6.30	8.81	0.171	0.247	0.171	0.201	0.372
Image 4	4.61	7.74	5.94	6.02	5.97	0.182	0.219	0.224	0.215	0.160
Image 5	4.67	10.68	6.98	6.96	8.93	0.135	0.213	0.246	0.182	0.171
Image 6	4.09	5.38	4.88	5.01	7.22	0.237	0.189	0.229	0.187	0.526
Image 7	4.40	6.56	5.52	5.41	7.63	0.241	0.211	0.237	0.176	0.410
Image 8	4.38	5.97	5.38	5.30	7.12	0.258	0.172	0.270	0.168	0.455
Image 9	4.83	10.93	6.36	7.15	9.71	0.130	0.489	0.201	0.342	0.416
Average	4.41	7.70	5.71	5.74	7.79	0.191	0.296	0.216	0.200	0.390

A.1.7 Experiment 7: 100% Data Availability and $\sigma_n = \sigma_y$

Table A.7 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR for 8 images, and best SSIM for 6 images. Fig. A.7 shows the reconstruction results from all methods for this experiment. PnP-CNN-SAR produces better reconstructions compared to other methods.

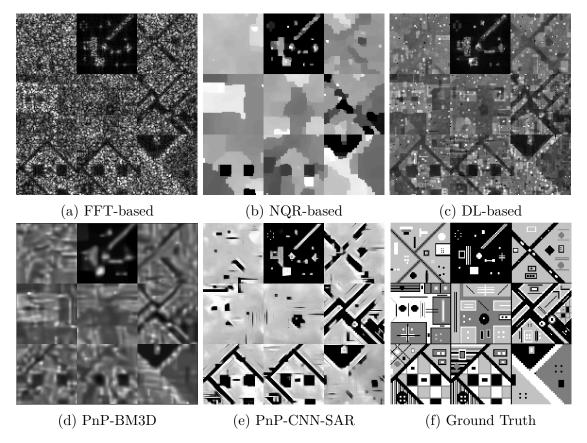


Figure A.6 Reconstruction results for the synthetic scenes for the case of 30% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$.

Table A.7 SNR and SSIM values for synthetic SAR scenes for the case of 100% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			SNR (dE	3)				SSIM		
	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR
Image 1	6.76	13.26	6.97	10.62	13.69	0.364	0.560	0.252	0.336	0.568
Image 2	3.93	14.45	4.02	3.75	14.34	0.313	0.907	0.275	0.244	0.735
Image 3	6.19	13.96	6.36	9.03	14.23	0.403	0.678	0.291	0.360	0.669
Image 4	6.31	10.04	6.77	9.60	12.79	0.507	0.447	0.387	0.501	0.668
Image 5	6.69	12.79	7.16	11.06	13.45	0.393	0.453	0.288	0.431	0.529
Image 6	5.21	9.94	5.60	6.63	14.23	0.622	0.651	0.486	0.520	0.832
Image 7	5.66	10.50	6.07	7.75	13.90	0.593	0.644	0.447	0.526	0.789
Image 8	5.54	10.19	5.91	7.66	13.15	0.591	0.655	0.462	0.558	0.802
Image 9	6.23	16.11	6.31	9.91	16.18	0.256	0.684	0.195	0.322	0.594
Average	5.84	12.36	6.13	8.45	13.99	0.449	0.631	0.343	0.422	0.687

A.1.8 Experiment 8: 90% Data Availability and $\sigma_n = \sigma_y$

Table A.8 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR for 5 images, and best SSIM for 6 images. Fig. A.8 shows the reconstruction results from all methods for this experiment. PnP-CNN-SAR produces better reconstructions compared to other methods.

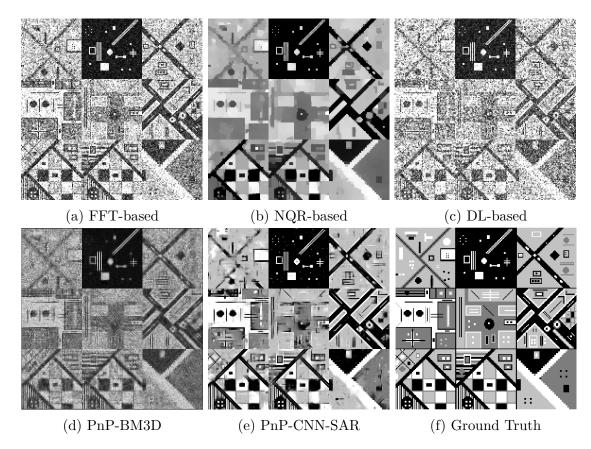


Figure A.7 Reconstruction results for the synthetic scenes for the case of 100% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$.

Table A.8 SNR and SSIM values for synthetic SAR scenes for the case of 90% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			SNR (dE	3)				SSIM		
	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR
Image 1	6.54	12.25	6.04	10.74	10.96	0.270	0.427	0.191	0.312	0.415
Image 2	3.21	13.58	3.10	2.98	12.84	0.285	0.912	0.260	0.225	0.566
Image 3	5.80	10.11	4.98	8.35	10.56	0.311	0.488	0.202	0.290	0.507
Image 4	6.41	9.82	5.39	9.03	11.03	0.426	0.432	0.312	0.431	0.621
Image 5	6.75	11.80	5.88	10.78	11.69	0.314	0.423	0.219	0.357	0.436
Image 6	4.99	7.81	3.78	6.08	12.80	0.518	0.468	0.393	0.447	0.793
Image 7	5.52	8.95	4.56	7.05	12.01	0.483	0.507	0.359	0.431	0.725
Image 8	5.47	9.06	4.01	6.84	12.59	0.483	0.551	0.329	0.444	0.767
Image 9	5.93	15.86	5.59	9.13	15.14	0.206	0.737	0.159	0.258	0.488
Average	5.62	11.03	4.81	7.89	12.18	0.366	0.549	0.269	0.355	0.591

A.1.9 Experiment 9: 80% Data Availability and $\sigma_n = \sigma_y$

Table A.9 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR for 6 images, and best SSIM for 5 images. Fig. A.9 shows the reconstruction results from all methods for this experiment. PnP-CNN-SAR produces better reconstructions compared to other methods.

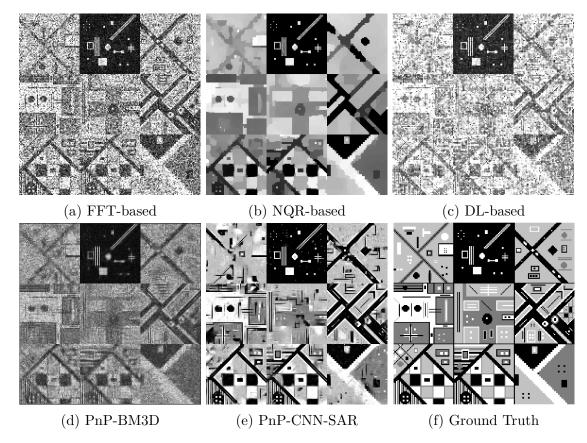


Figure A.8 Reconstruction results for the synthetic scenes for the case of 90% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$.

Table A.9 SNR and SSIM values for synthetic SAR scenes for the case of 80% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			SNR (dE	3)				SSIM		
	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR
Image 1	6.70	11.19	6.99	11.00	10.65	0.246	0.328	0.206	0.241	0.352
Image 2	3.38	9.11	3.55	2.07	11.93	0.271	0.846	0.254	0.176	0.420
Image 3	5.95	10.60	5.81	8.60	11.07	0.295	0.511	0.214	0.248	0.460
Image 4	6.36	9.16	5.71	8.60	9.94	0.375	0.357	0.295	0.343	0.498
Image 5	6.84	12.24	6.66	10.86	11.42	0.292	0.431	0.235	0.261	0.387
Image 6	4.89	8.69	4.26	5.24	11.11	0.468	0.578	0.362	0.279	0.707
Image 7	5.36	7.75	5.12	6.67	9.83	0.456	0.350	0.366	0.297	0.572
Image 8	5.29	6.50	4.72	6.29	10.36	0.437	0.273	0.360	0.322	0.635
Image 9	5.95	14.90	5.92	9.47	12.99	0.185	0.702	0.153	0.283	0.363
Average	5.64	10.02	5.42	7.64	11.03	0.336	0.486	0.272	0.272	0.488

A.1.10 Experiment 10: 70% Data Availability and $\sigma_n = \sigma_y$

Table A.10 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR for 8 images, and best SSIM for 6 images. Fig. A.10 shows the reconstruction results from all methods for this experiment. PnP-CNN-SAR produces better reconstructions compared to other methods.

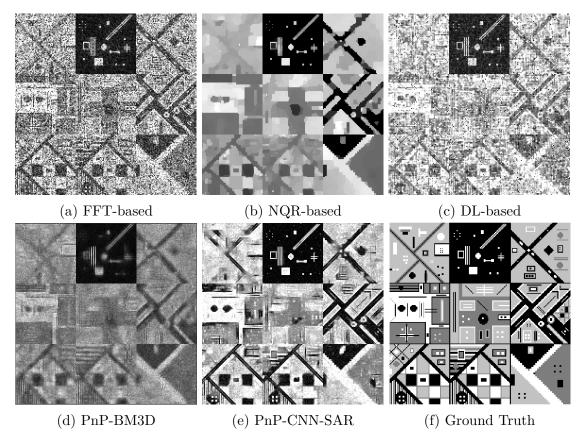


Figure A.9 Reconstruction results for the synthetic scenes for the case of 80% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$.

Table A.10 SNR and SSIM values for synthetic SAR scenes for the case of 70% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			SNR (dE	3)				SSIM		
	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR
Image 1	6.64	10.51	7.44	10.13	10.65	0.223	0.264	0.185	0.211	0.297
Image 2	2.98	6.54	3.53	2.54	11.17	0.249	0.807	0.239	0.175	0.427
Image 3	5.81	9.29	6.31	8.38	9.38	0.242	0.392	0.192	0.212	0.307
Image 4	6.21	8.93	6.51	8.38	10.02	0.314	0.327	0.291	0.290	0.475
Image 5	6.56	10.09	7.69	10.29	10.95	0.250	0.304	0.231	0.257	0.339
Image 6	4.61	7.33	4.73	5.59	9.36	0.412	0.480	0.330	0.301	0.606
Image 7	5.41	7.57	5.48	6.93	9.97	0.383	0.306	0.322	0.320	0.565
Image 8	5.15	7.01	5.48	6.64	9.10	0.402	0.348	0.340	0.331	0.548
Image 9	5.87	13.45	7.00	8.94	12.71	0.170	0.654	0.159	0.196	0.338
Average	5.47	8.97	6.02	7.54	10.37	0.294	0.431	0.254	0.255	0.433

A.1.11 Experiment 11: 50% Data Availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$

Table A.11 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR for 5 images, and best SSIM for 6 images. Fig. A.11 shows the reconstruction results from all methods for this experiment. PnP-CNN-SAR produces better reconstructions compared to other methods.

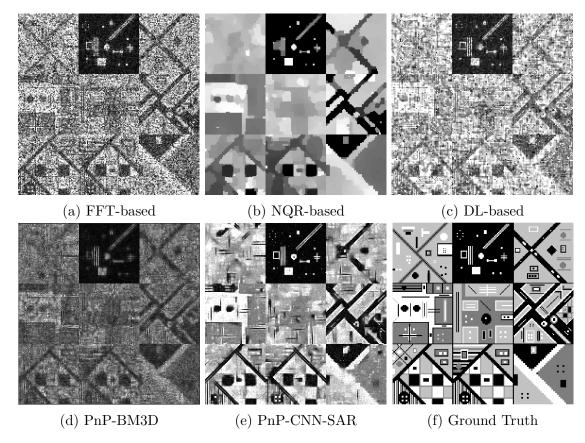


Figure A.10 Reconstruction results for the synthetic scenes for the case of 70% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$.

Table A.11 SNR and SSIM values for synthetic SAR scenes for the case of 50% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			SNR (dE	3)				SSIM		
	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR
Image 1	6.43	9.21	8.77	9.63	9.72	0.162	0.213	0.173	0.209	0.248
Image 2	2.43	4.41	3.07	1.58	8.20	0.209	0.638	0.202	0.098	0.412
Image 3	5.71	8.79	7.00	7.83	8.70	0.188	0.302	0.154	0.223	0.316
Image 4	5.78	8.35	7.23	7.67	7.72	0.233	0.258	0.227	0.191	0.260
Image 5	6.50	10.90	8.57	9.84	10.19	0.185	0.259	0.208	0.201	0.220
Image 6	4.49	5.83	4.95	5.02	7.30	0.317	0.245	0.248	0.134	0.466
Image 7	5.15	6.65	6.14	6.08	7.49	0.296	0.218	0.253	0.143	0.371
Image 8	5.00	6.47	5.71	5.87	7.18	0.314	0.227	0.269	0.170	0.433
Image 9	5.84	11.60	7.56	8.73	9.22	0.141	0.480	0.150	0.374	0.288
Average	5.26	8.02	6.55	6.92	8.41	0.227	0.316	0.209	0.194	0.335

A.1.12 Experiment 12: 30% Data Availability and $\sigma_n = \sigma_y$

Table A.12 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR for 3 images, and best SSIM for 3 images. Fig. A.12 shows the reconstruction results from all methods for this experiment. None of the methods are able to produce acceptable reconstructions for most of the images.

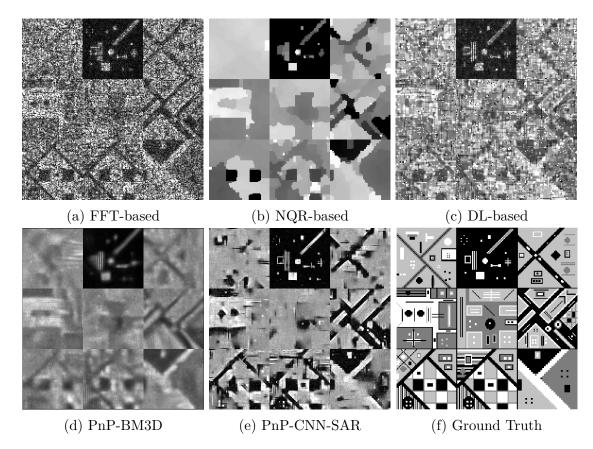


Figure A.11 Reconstruction results for the synthetic scenes for the case of 50% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$.

Table A.12 SNR and SSIM values for synthetic SAR scenes for the case of 30% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			SNR (dE	3)				SSIM		
	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR
Image 1	5.40	10.16	9.30	7.77	8.65	0.120	0.219	0.192	0.134	0.154
Image 2	2.16	4.10	2.08	1.68	3.83	0.149	0.681	0.119	0.085	0.285
Image 3	4.89	6.38	6.89	6.76	7.00	0.136	0.196	0.149	0.171	0.201
Image 4	4.71	7.62	7.06	6.53	6.90	0.145	0.186	0.192	0.163	0.167
Image 5	5.48	10.36	9.04	7.96	9.02	0.154	0.184	0.183	0.165	0.158
Image 6	3.99	5.12	4.17	4.62	5.07	0.209	0.174	0.067	0.116	0.204
Image 7	4.56	6.12	5.70	5.69	6.35	0.205	0.138	0.112	0.155	0.261
Image 8	4.40	5.68	5.14	5.39	5.90	0.201	0.129	0.091	0.152	0.208
Image 9	5.09	10.34	7.41	7.22	8.31	0.106	0.460	0.303	0.243	0.309
Average	4.52	7.32	6.31	5.96	6.78	0.158	0.263	0.156	0.154	0.216

A.2 Same Scene Experiments

A.2.1 Experiment 1: 100% Data Availability and $\sigma_n = 0.1\sigma_y$

Table A.13 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR and SSIM for all 31 images. Fig. A.13 shows the reconstruction results for a selection of images from the Wonsan test set from all methods for this experiment. All methods produce good reconstructions.

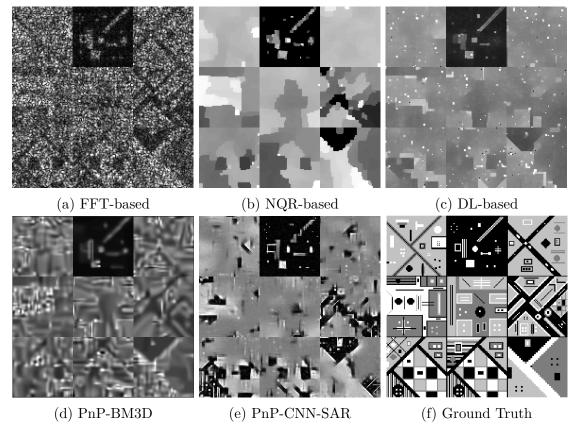


Figure A.12 Reconstruction results for the synthetic scenes for the case of 30% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$.

Table A.13 SNR and SSIM values for images from the Wonsan Test Set for the case of 100% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			SNR (dE	3)				SSIM		
	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR
Image 1	26.31	24.60	22.58	26.28	40.91	0.900	0.906	0.937	0.899	0.992
Image 2	26.74	24.59	22.61	26.79	43.99	0.986	0.985	0.982	0.986	1.000
Image 3	26.67	12.02	26.13	26.62	35.68	0.983	0.856	0.981	0.983	0.998
Image 4	26.75	25.65	21.41	26.82	39.38	0.989	0.989	0.984	0.989	0.999
Image 5	26.76	25.14	21.77	26.70	42.66	0.990	0.989	0.984	0.990	1.000
Image 6	26.36	25.10	21.73	26.35	40.02	0.938	0.941	0.957	0.939	0.99
Image 7	26.67	25.22	21.36	26.66	40.81	0.987	0.987	0.983	0.987	0.999
Image 8	26.75	25.45	21.88	26.73	41.24	0.984	0.984	0.981	0.984	0.999
Image 9	26.64	24.53	21.74	26.73	42.06	0.979	0.979	0.978	0.979	0.999
Image 10	26.63	25.84	21.99	26.75	38.59	0.978	0.979	0.980	0.979	0.999
Image 11	26.72	24.93	22.83	26.72	41.32	0.986	0.985	0.982	0.986	1.000
Image 12	26.29	24.93	22.24	26.36	41.16	0.927	0.931	0.953	0.928	0.994
Image 13	26.71	17.00	26.46	26.67	37.39	0.979	0.935	0.979	0.979	0.998
Image 14	26.65	24.58	21.25	26.68	41.48	0.980	0.980	0.978	0.980	0.999
Image 15	26.65	26.01	23.08	26.70	41.33	0.988	0.987	0.984	0.988	1.000
Image 16	26.69	24.71	21.94	26.73	41.57	0.987	0.986	0.982	0.987	0.999
Image 17	26.95	24.98	25.77	26.75	34.66	0.957	0.960	0.973	0.955	0.99
Image 18	26.73	25.46	22.94	26.69	43.06	0.985	0.985	0.982	0.985	1.000
Image 19	26.61	24.81	21.33	26.83	41.72	0.985	0.985	0.982	0.986	1.000
Image 20	26.74	25.58	23.12	26.69	39.18	0.987	0.987	0.984	0.987	0.999
Image 21	26.64	24.80	24.05	26.75	41.17	0.984	0.984	0.982	0.985	1.000
Image 22	26.73	24.76	21.51	26.78	40.01	0.985	0.984	0.982	0.985	0.999
Image 23	26.75	25.02	24.29	26.72	40.95	0.986	0.985	0.983	0.986	1.000
Image 24	26.78	24.63	23.35	26.73	41.59	0.981	0.981	0.981	0.981	0.999
Image 25	26.79	12.07	26.64	26.67	35.27	0.981	0.851	0.978	0.980	0.99
Image 26	26.74	24.26	21.27	26.75	37.94	0.968	0.969	0.975	0.968	0.998
Image 27	26.65	24.10	22.77	26.70	41.25	0.985	0.984	0.983	0.985	0.999
Image 28	25.97	25.25	22.20	25.97	37.01	0.898	0.904	0.942	0.898	0.98
Image 29	26.74	26.48	25.03	26.70	41.78	0.985	0.985	0.983	0.985	0.999
Image 30	26.52	24.40	21.48	26.52	40.51	0.947	0.950	0.959	0.949	0.99
Image 31	26.71	13.35	24.67	26.72	35.22	0.989	0.956	0.985	0.989	0.99
Average	26.65	23.56	22.95	26.65	40.03	0.973	0.963	0.976	0.973	0.998

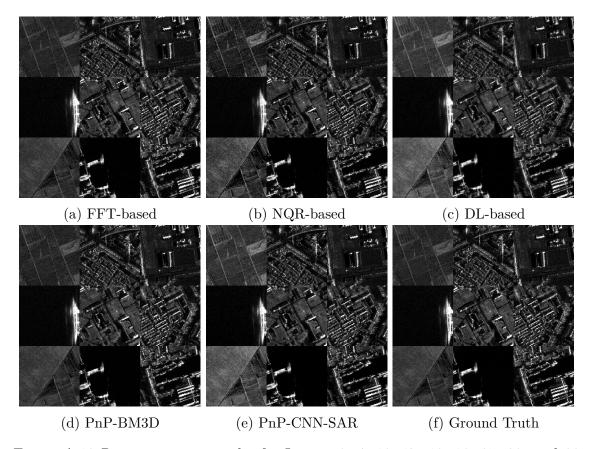


Figure A.13 Reconstruction results for Images 3, 4, 10, 17, 18, 19, 25, 28, and 30 from the Wonsan test set for the case of 100% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$.

A.2.2 Experiment 2: 90% Data Availability and $\sigma_n = 0.1\sigma_y$

Table A.14 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR for 27 images, and best SSIM for 29 images. Fig. A.14 shows the reconstruction results for a selection of images from the Wonsan test set from all methods for this experiment. All methods produce good reconstructions.

A.2.3 Experiment 3: 80% Data Availability and $\sigma_n = 0.1\sigma_y$

Table A.15 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR for 27 images, and best SSIM for 29 images. Fig. A.15 shows the reconstruction results for a selection of images from the Wonsan test set from all methods for this experiment. PnP-BM3D and PnP-CNN-SAR produce better reconstructions compared to other methods.

A.2.4 Experiment 4: 70% Data Availability and $\sigma_n = 0.1\sigma_y$

Table A.16 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR for 29 images, and best SSIM for 28 images. Fig. A.16 shows

Table A.14 SNR and SSIM values for images from the Wonsan Test Set for the case of 90% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			SNR (dE	5)				SSIM		
	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR
Image 1	13.19	13.25	12.32	15.55	18.06	0.706	0.608	0.698	0.783	0.925
Image 2	13.47	11.48	12.14	14.43	17.16	0.832	0.667	0.790	0.858	0.926
Image 3	13.55	9.82	16.45	17.73	17.00	0.762	0.601	0.829	0.872	0.843
Image 4	13.37	10.67	11.56	13.01	15.24	0.862	0.831	0.802	0.846	0.901
Image 5	13.31	11.57	11.73	13.25	15.71	0.859	0.839	0.809	0.851	0.913
Image 6	12.84	11.43	10.70	13.49	17.11	0.751	0.765	0.679	0.770	0.937
Image 7	13.21	11.90	11.15	13.01	16.40	0.850	0.842	0.781	0.841	0.934
Image 8	13.26	10.81	11.68	13.57	14.79	0.837	0.647	0.783	0.846	0.875
Image 9	13.16	11.92	11.14	13.96	15.49	0.822	0.676	0.748	0.846	0.888
Image 10	13.21	10.18	11.60	15.12	16.59	0.802	0.526	0.787	0.838	0.901
Image 11	13.36	11.64	13.08	14.22	16.07	0.830	0.816	0.822	0.853	0.892
Image 12	12.99	11.28	11.69	14.41	17.01	0.733	0.587	0.691	0.788	0.925
Image 13	13.28	12.60	16.45	17.48	17.66	0.729	0.710	0.814	0.848	0.849
Image 14	13.11	11.44	10.43	13.54	15.74	0.830	0.681	0.733	0.842	0.902
Image 15	13.36	12.47	13.34	14.39	16.34	0.835	0.829	0.831	0.850	0.902
Image 16	13.23	10.96	11.76	13.44	15.42	0.847	0.663	0.791	0.851	0.899
Image 17	13.10	12.23	15.62	18.22	17.89	0.818	0.689	0.854	0.784	0.907
Image 18	13.25	11.61	12.67	15.29	17.45	0.801	0.792	0.790	0.859	0.909
Image 19	13.32	11.35	10.77	13.31	16.94	0.846	0.829	0.780	0.854	0.925
Image 20	13.41	11.34	13.64	13.96	15.59	0.842	0.824	0.846	0.836	0.896
Image 21	13.40	12.66	14.31	15.31	16.48	0.813	0.811	0.831	0.855	0.890
Image 22	13.29	11.05	11.20	13.52	16.73	0.849	0.667	0.770	0.846	0.921
Image 23	13.29	11.62	14.52	15.80	16.59	0.797	0.505	0.827	0.862	0.878
Image 24	13.13	12.14	13.38	15.12	17.37	0.806	0.616	0.816	0.853	0.908
Image 25	13.40	10.13	16.74	18.21	17.95	0.711	0.559	0.806	0.860	0.835
Image 26	13.12	12.89	10.03	15.44	18.13	0.781	0.581	0.695	0.830	0.895
Image 27	13.29	9.57	13.08	14.52	16.29	0.827	0.450	0.827	0.856	0.897
Image 28	12.65	14.40	10.59	17.01	19.23	0.698	0.545	0.668	0.769	0.934
Image 29	13.32	12.99	15.14	16.81	16.54	0.786	0.787	0.843	0.871	0.881
Image 30	12.97	12.92	10.33	14.79	19.01	0.754	0.677	0.671	0.830	0.919
Image 31	13.40	10.54	14.89	13.75	16.18	0.835	0.557	0.858	0.816	0.884
Average	13.23	11.64	12.71	14.89	16.78	0.802	0.683	0.783	0.838	0.900

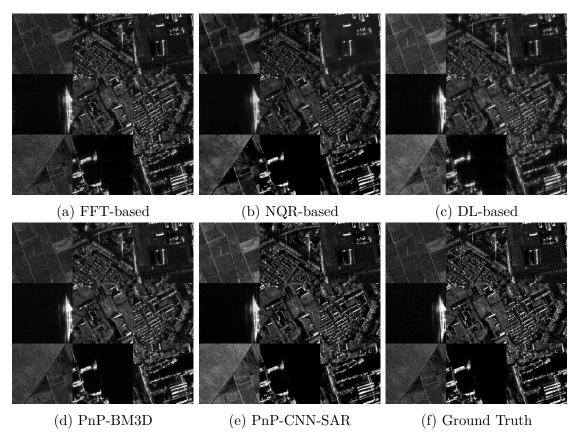


Figure A.14 Reconstruction results for Images 3, 4, 10, 17, 18, 19, 25, 28, and 30 from the Wonsan test set for the case of 100% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$.

Table A.15 SNR and SSIM values for images from the Wonsan Test Set for the case of 80% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			SNR (dE	3)				SSIM		
	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR
Image 1	9.77	8.98	9.69	11.73	13.41	0.653	0.485	0.632	0.697	0.867
Image 2	9.96	7.23	9.65	11.33	13.30	0.722	0.334	0.671	0.757	0.854
Image 3	10.15	8.55	13.04	14.50	13.69	0.597	0.566	0.681	0.747	0.715
Image 4	10.12	5.36	8.76	10.00	12.84	0.775	0.338	0.681	0.735	0.872
Image 5	10.20	8.90	9.16	10.55	12.57	0.763	0.732	0.681	0.747	0.848
Image 6	9.83	8.80	8.61	10.21	11.97	0.702	0.500	0.613	0.678	0.873
Image 7	9.99	9.35	8.56	10.06	12.84	0.765	0.754	0.659	0.737	0.873
Image 8	9.96	8.19	8.96	10.52	12.93	0.750	0.714	0.666	0.746	0.867
Image 9	9.99	7.64	8.87	10.90	13.60	0.742	0.483	0.650	0.752	0.868
Image 10	10.09	7.23	9.10	11.32	13.18	0.720	0.443	0.692	0.732	0.825
Image 11	10.07	7.48	10.12	11.28	13.07	0.712	0.306	0.686	0.734	0.813
Image 12	9.78	6.76	9.23	10.86	12.47	0.667	0.363	0.631	0.684	0.867
Image 13	10.10	11.59	13.12	14.07	14.40	0.571	0.511	0.668	0.712	0.724
Image 14	9.82	7.99	8.16	10.25	13.29	0.748	0.530	0.631	0.749	0.888
Image 15	10.10	9.27	10.37	11.45	12.85	0.721	0.487	0.697	0.741	0.817
Image 16	10.23	8.60	9.25	10.71	13.10	0.753	0.715	0.673	0.755	0.845
Image 17	9.86	12.48	12.43	14.19	13.84	0.780	0.655	0.803	0.719	0.850
Image 18	10.10	9.73	9.84	11.94	13.70	0.685	0.521	0.658	0.744	0.828
Image 19	9.85	8.77	8.32	10.24	13.38	0.752	0.574	0.663	0.755	0.871
Image 20	10.07	9.79	10.45	11.24	12.73	0.728	0.725	0.720	0.717	0.815
Image 21	10.19	9.16	11.28	12.14	13.08	0.685	0.335	0.698	0.732	0.785
Image 22	9.95	8.12	8.66	10.39	13.32	0.761	0.522	0.659	0.740	0.876
Image 23	10.28	9.75	11.52	12.42	13.13	0.663	0.345	0.691	0.736	0.772
Image 24	10.07	8.27	10.56	11.74	13.57	0.696	0.358	0.695	0.742	0.820
Image 25	10.19	10.45	13.21	15.28	14.44	0.543	0.502	0.631	0.731	0.684
Image 26	10.07	9.32	7.81	11.75	14.92	0.714	0.527	0.613	0.742	0.857
Image 27	10.17	7.71	10.09	11.47	13.00	0.710	0.353	0.695	0.739	0.814
Image 28	9.67	10.31	8.45	13.10	12.61	0.685	0.529	0.646	0.703	0.884
Image 29	10.12	8.71	11.99	13.41	13.60	0.645	0.386	0.707	0.754	0.788
Image 30	9.65	9.55	8.13	10.98	13.86	0.689	0.573	0.585	0.732	0.883
Image 31	10.11	9.66	11.67	11.05	13.43	0.713	0.522	0.735	0.680	0.788
Average	10.02	8.83	9.97	11.65	13.29	0.704	0.506	0.671	0.731	0.830

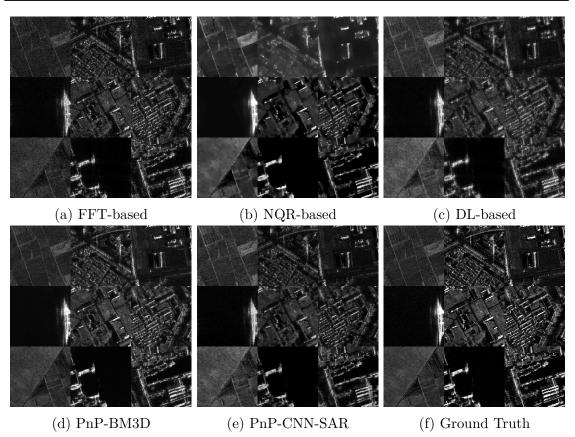


Figure A.15 Reconstruction results for Images 3, 4, 10, 17, 18, 19, 25, 28, and 30 from the Wonsan test set for the case of 100% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$.

the reconstruction results for a selection of images from the Wonsan test set from all methods for this experiment. PnP-CNN-SAR produce better reconstructions compared to other methods.

Table A.16 SNR and SSIM values for images from the Wonsan Test Set for the case of 70% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			3)		SSIM					
	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR
Image 1	8.32	7.53	8.47	9.30	10.56	0.605	0.432	0.579	0.626	0.793
Image 2	8.49	6.97	8.47	9.20	10.94	0.650	0.279	0.587	0.655	0.768
Image 3	8.52	8.24	11.01	12.12	11.59	0.503	0.427	0.550	0.597	0.572
Image 4	8.45	6.64	7.65	8.47	10.18	0.705	0.634	0.594	0.648	0.793
Image 5	8.43	7.18	7.86	8.71	10.09	0.678	0.631	0.583	0.646	0.752
Image 6	8.18	6.47	7.52	8.58	10.25	0.652	0.430	0.562	0.612	0.820
Image 7	8.35	7.30	7.46	8.43	10.21	0.696	0.664	0.585	0.652	0.800
Image 8	8.29	6.62	7.89	8.66	10.31	0.681	0.368	0.592	0.653	0.792
Image 9	8.30	6.53	7.62	8.92	10.85	0.669	0.372	0.563	0.665	0.799
Image 10	8.36	6.24	7.89	9.01	10.30	0.648	0.398	0.620	0.646	0.741
Image 11	8.42	6.16	8.74	9.62	10.49	0.619	0.296	0.584	0.633	0.697
Image 12	8.32	7.48	8.14	9.11	10.35	0.635	0.441	0.582	0.618	0.807
Image 13	8.53	10.62	11.17	11.39	11.81	0.481	0.466	0.538	0.560	0.554
Image 14	8.13	6.74	7.17	8.48	10.47	0.687	0.438	0.558	0.668	0.819
Image 15	8.64	6.78	9.09	9.56	10.62	0.641	0.270	0.599	0.620	0.708
Image 16	8.36	7.00	7.94	8.99	10.38	0.669	0.621	0.563	0.656	0.753
Image 17	8.57	10.52	10.38	10.26	10.90	0.748	0.620	0.757	0.685	0.796
Image 18	8.48	7.83	8.68	9.70	10.83	0.606	0.340	0.564	0.629	0.724
Image 19	8.29	6.86	7.29	8.62	10.51	0.687	0.635	0.580	0.664	0.795
Image 20	8.44	7.11	9.20	9.57	10.19	0.635	0.299	0.634	0.605	0.710
Image 21	8.56	8.73	9.70	10.12	10.87	0.598	0.330	0.586	0.618	0.668
Image 22	8.37	4.66	7.53	8.64	10.73	0.690	0.253	0.577	0.659	0.801
Image 23	8.63	8.26	9.89	10.21	10.92	0.569	0.284	0.572	0.605	0.645
Image 24	8.67	7.49	9.13	9.67	11.12	0.626	0.316	0.596	0.636	0.720
Image 25	8.56	9.20	11.21	12.22	11.97	0.445	0.425	0.500	0.555	0.500
Image 26	8.11	7.03	6.96	8.79	11.83	0.649	0.420	0.568	0.657	0.779
Image 27	8.31	6.83	8.82	9.49	10.54	0.616	0.252	0.588	0.625	0.710
Image 28	8.32	8.96	7.75	9.72	10.03	0.684	0.512	0.651	0.651	0.842
Image 29	8.55	8.23	10.35	11.10	11.33	0.556	0.317	0.606	0.616	0.681
Image 30	8.13	8.38	7.11	8.90	10.80	0.640	0.528	0.525	0.652	0.812
Image 31	8.56	7.71	10.13	9.97	11.23	0.633	0.531	0.639	0.570	0.677
Average	8.41	7.49	8.65	9.53	10.75	0.632	0.427	0.587	0.632	0.736

A.2.5 Experiment 5: 50% Data Availability and $\sigma_n = 0.1\sigma_y$

Table A.17 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR for 23 images, and best SSIM for 29 images. Fig. A.17 shows the reconstruction results for a selection of images from the Wonsan test set from all methods for this experiment. PnP-CNN-SAR produce better reconstructions compared to other methods.

A.2.6 Experiment 6: 30% Data Availability and $\sigma_n = 0.1\sigma_y$

Table A.18 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR for 5 images, and best SSIM for 23 images. Fig. A.18 shows the reconstruction results for a selection of images from the Wonsan test set from all methods for this experiment. PnP-CNN-SAR produce better reconstructions compared to other methods.

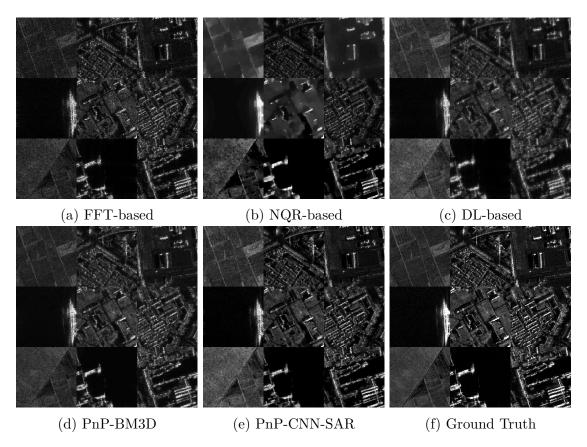


Figure A.16 Reconstruction results for Images 3, 4, 10, 17, 18, 19, 25, 28, and 30 from the Wonsan test set for the case of 100% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{v}}$.

Table A.17 SNR and SSIM values for images from the Wonsan Test Set for the case of 50% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			3)		SSIM					
	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR
Image 1	6.10	7.06	6.68	7.26	8.20	0.541	0.410	0.505	0.487	0.656
Image 2	6.26	4.80	6.78	7.44	7.97	0.513	0.190	0.437	0.461	0.595
Image 3	6.61	7.93	8.68	10.63	8.81	0.384	0.418	0.374	0.404	0.393
Image 4	6.10	5.76	5.99	6.41	7.40	0.555	0.343	0.441	0.444	0.620
Image 5	6.17	6.03	6.33	6.77	7.52	0.527	0.266	0.419	0.421	0.580
Image 6	6.09	3.54	6.19	6.54	7.56	0.579	0.200	0.488	0.468	0.651
Image 7	6.13	6.09	6.02	6.25	7.46	0.563	0.561	0.444	0.435	0.631
Image 8	6.03	3.42	6.28	6.84	7.64	0.546	0.175	0.451	0.469	0.625
Image 9	6.10	4.06	6.11	6.82	7.70	0.548	0.187	0.438	0.484	0.657
Image 10	6.21	5.12	6.31	7.22	7.74	0.533	0.357	0.505	0.492	0.618
Image 11	6.29	5.51	7.03	7.80	7.93	0.478	0.207	0.419	0.424	0.539
Image 12	6.15	6.76	6.49	7.00	7.95	0.558	0.407	0.502	0.474	0.642
Image 13	6.49	11.04	8.67	9.97	9.22	0.348	0.354	0.364	0.362	0.386
Image 14	6.04	3.13	5.75	6.38	7.62	0.563	0.150	0.429	0.478	0.674
Image 15	6.32	7.33	7.05	7.90	8.08	0.486	0.282	0.429	0.416	0.539
Image 16	6.19	4.40	6.36	6.94	7.70	0.519	0.151	0.408	0.426	0.589
Image 17	6.35	8.81	7.98	8.33	8.26	0.680	0.588	0.694	0.648	0.719
Image 18	6.31	5.13	6.84	7.78	8.31	0.472	0.184	0.413	0.414	0.557
Image 19	6.06	2.99	5.87	6.57	7.73	0.550	0.177	0.437	0.476	0.642
Image 20	6.31	5.50	7.24	7.83	8.00	0.500	0.263	0.464	0.401	0.543
Image 21	6.40	6.44	7.73	8.57	8.34	0.447	0.200	0.420	0.414	0.492
Image 22	6.16	3.40	6.08	6.61	7.71	0.553	0.175	0.439	0.465	0.639
Image 23	6.30	6.29	7.80	8.70	8.45	0.410	0.196	0.393	0.392	0.469
Image 24	6.28	5.72	7.25	7.93	8.24	0.475	0.230	0.446	0.455	0.562
Image 25	6.51	9,99	8.75	10.75	9.25	0.326	0.336	0.330	0.346	0.328
Image 26	5.91	5.51	5.51	6.60	8.46	0.530	0.371	0.456	0.505	0.670
Image 27	6.25	5.63	6.94	7.69	7.86	0.480	0.198	0.428	0.419	0.531
Image 28	5.88	6.72	6.02	7.11	8.23	0.668	0.482	0.629	0.591	0.710
Image 29	6.45	5.49	8.11	9.66	8.79	0.430	0.251	0.455	0.401	0.505
Image 30	6.05	5.70	5.65	6.55	8.25	0.540	0.373	0.411	0.473	0.687
Image 31	6.41	6.26	8.05	9.22	8.95	0.511	0.472	0.495	0.475	0.537
Average	6.22	5.86	6.85	7.68	8.11	0.510	0.295	0.450	0.452	0.580

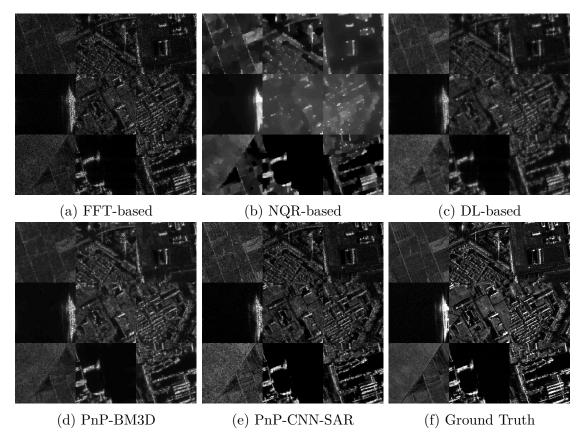


Figure A.17 Reconstruction results for Images 3, 4, 10, 17, 18, 19, 25, 28, and 30 from the Wonsan test set for the case of 100% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{v}}$.

Table A.18 SNR and SSIM values for images from the Wonsan Test Set for the case of 30% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			3)		SSIM					
	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR
Image 1	4.58	5.75	5.05	5.78	5.75	0.489	0.372	0.440	0.431	0.490
Image 2	4.60	5.39	5.09	5.76	5.40	0.371	0.174	0.296	0.302	0.40
Image 3	4.86	6.55	6.37	7.94	5.95	0.291	0.327	0.253	0.335	0.268
Image 4	4.46	3.37	4.57	5.10	5.08	0.416	0.180	0.313	0.289	0.410
Image 5	4.41	4.13	4.81	5.33	5.14	0.366	0.166	0.272	0.247	0.38
Image 6	4.36	5.60	4.60	5.16	5.71	0.501	0.379	0.414	0.384	0.502
Image 7	4.39	5.50	4.52	4.92	5.25	0.416	0.349	0.313	0.279	0.428
Image 8	4.39	5.61	4.75	5.33	5.24	0.410	0.288	0.320	0.311	0.40
Image 9	4.27	4.16	4.76	5.19	5.15	0.400	0.173	0.314	0.328	0.43
Image 10	4.43	4.98	4.73	5.81	5.34	0.415	0.354	0.401	0.416	0.46
Image 11	4.61	6.07	5.25	6.03	5.49	0.338	0.199	0.278	0.270	0.356
Image 12	4.38	4.73	4.90	5.59	6.18	0.481	0.238	0.419	0.413	0.512
Image 13	4.87	10.37	6.43	7.64	6.15	0.257	0.312	0.242	0.264	0.25
Image 14	4.30	3.52	4.41	4.92	4.97	0.419	0.149	0.306	0.325	0.443
Image 15	4.56	6.82	5.34	6.18	5.66	0.345	0.273	0.285	0.265	0.369
Image 16	4.50	4.90	4.78	5.43	5.37	0.372	0.137	0.263	0.254	0.372
Image 17	4.91	8.67	5.75	6.89	5.88	0.607	0.605	0.637	0.649	0.58
Image 18	4.55	5.37	5.18	5.92	5.48	0.339	0.183	0.285	0.279	0.37
Image 19	4.44	3.51	4.44	5.14	5.11	0.404	0.183	0.304	0.310	0.419
Image 20	4.67	6.17	5.45	6.29	5.62	0.373	0.265	0.341	0.305	0.388
Image 21	4.68	6.90	5.77	6.52	5.67	0.324	0.166	0.281	0.241	0.32
Image 22	4.45	4.02	4.59	5.22	5.13	0.406	0.172	0.304	0.313	0.420
Image 23	4.67	7.72	5.87	6.64	5.80	0.284	0.181	0.252	0.234	0.32
Image 24	4.59	5.95	5.34	6.26	5.53	0.359	0.232	0.304	0.309	0.38
Image 25	4.90	8.26	6.44	8.02	6.13	0.246	0.294	0.211	0.285	0.225
Image 26	4.24	4.11	4.22	5.19	4.92	0.416	0.287	0.362	0.412	0.474
Image 27	4.61	5.82	5.19	6.05	5.37	0.349	0.208	0.290	0.274	0.356
Image 28	4.09	5.08	4.57	5.50	6.50	0.666	0.466	0.631	0.605	0.576
Image 29	4.76	7.77	5.99	7.10	6.17	0.322	0.265	0.321	0.297	0.359
Image 30	4.30	4.68	4.31	5.16	5.34	0.427	0.314	0.318	0.362	0.498
Image 31	4.69	5.60	5.94	7.44	5.91	0.408	0.462	0.385	0.443	0.403
Average	4.53	5.71	5.14	5.98	5.56	0.394	0.270	0.334	0.336	0.406

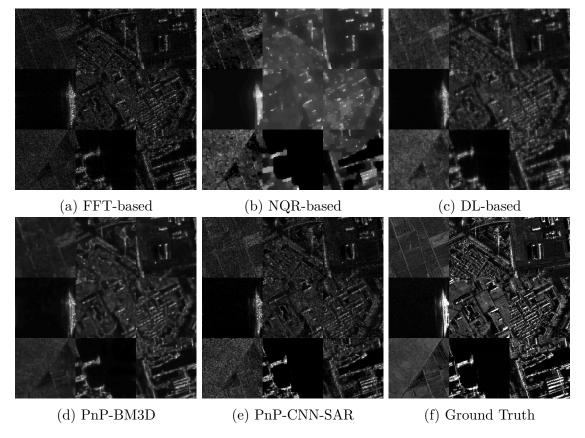


Figure A.18 Reconstruction results for Images 3, 4, 10, 17, 18, 19, 25, 28, and 30 from the Wonsan test set for the case of 100% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$.

A.2.7 Experiment 7: 100% Data Availability and $\sigma_n = \sigma_y$

Table A.19 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR and best SSIM for all 31 images. Fig. A.19 shows the reconstruction results for a selection of images from the Wonsan test set from all methods for this experiment. DL-based and PnP-CNN-SAR produce better reconstructions compared to other methods.

A.2.8 Experiment 8: 90% Data Availability and $\sigma_n = \sigma_y$

Table A.20 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR and best SSIM for all 31 images. Fig. A.20 shows the reconstruction results for a selection of images from the Wonsan test set from all methods for this experiment. PnP-CNN-SAR produce better reconstructions compared to other methods.

Table A.19 SNR and SSIM values for images from the Wonsan Test Set for the case of 100% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			SNR (dE	3)		SSIM					
	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR	
Image 1	5.41	10.06	15.71	5.46	39.11	0.345	0.479	0.924	0.343	0.99	
Image 2	6.48	9.31	15.65	6.59	40.53	0.500	0.535	0.925	0.504	1.000	
Image 3	6.97	10.14	20.18	7.21	33.66	0.408	0.481	0.921	0.420	0.997	
Image 4	6.49	8.61	14.28	6.61	36.84	0.578	0.582	0.924	0.582	0.999	
Image 5	6.52	9.06	14.69	6.63	37.93	0.557	0.544	0.922	0.563	0.999	
Image 6	5.37	9.12	14.76	5.49	39.21	0.395	0.552	0.917	0.401	0.99	
Image 7	6.21	9.00	14.24	6.28	38.49	0.548	0.582	0.923	0.555	0.999	
Image 8	6.26	7.45	14.82	6.33	37.92	0.522	0.459	0.925	0.526	0.999	
Image 9	6.17	8.73	14.67	6.18	35.93	0.511	0.561	0.922	0.508	0.99	
Image 10	6.54	10.86	14.85	6.68	37.82	0.415	0.556	0.931	0.420	0.998	
Image 11	6.71	9.20	15.93	6.80	36.95	0.504	0.470	0.923	0.511	0.999	
Image 12	5.42	9.68	15.33	5.60	39.86	0.373	0.540	0.922	0.379	0.995	
Image 13	6.93	12.16	20.66	6.97	32.68	0.372	0.486	0.921	0.370	0.998	
Image 14	5.98	8.57	14.14	6.03	35.48	0.526	0.577	0.921	0.530	0.997	
Image 15	6.64	10.21	16.19	6.78	38.05	0.505	0.535	0.925	0.511	0.999	
Image 16	6.46	9.00	14.89	6.54	34.12	0.558	0.563	0.920	0.565	0.998	
Image 17	5.51	13.30	19.52	5.57	35.60	0.193	0.601	0.949	0.199	0.994	
Image 18	6.61	9.11	16.00	6.63	39.79	0.486	0.456	0.924	0.485	0.999	
Image 19	6.28	8.77	14.17	6.45	37.46	0.548	0.580	0.924	0.555	0.999	
Image 20	6.70	10.09	16.26	6.89	37.06	0.495	0.524	0.928	0.508	0.999	
Image 21	6.76	10.11	17.34	6.81	36.62	0.466	0.497	0.924	0.475	0.999	
Image 22	6.42	8.68	14.40	6.43	36.85	0.537	0.556	0.923	0.540	0.999	
Image 23	6.88	10.35	17.66	7.03	33.88	0.472	0.454	0.922	0.476	0.998	
Image 24	6.54	10.14	16.53	6.63	37.66	0.458	0.539	0.928	0.464	0.999	
Image 25	7.01	8.08	21.07	7.10	31.47	0.367	0.402	0.917	0.367	0.993	
Image 26	6.04	9.95	14.06	6.15	37.39	0.418	0.569	0.926	0.417	0.998	
Image 27	6.67	9.66	15.82	6.70	37.46	0.491	0.511	0.924	0.494	0.999	
Image 28	4.74	11.68	15.28	4.67	37.08	0,206	0.721	0.922	0.208	0.98	
Image 29	6.90	11.57	18.58	7.00	38.68	0.444	0.535	0.929	0.454	0.999	
Image 30	5.68	8.23	14.40	5.74	36.19	0.464	0.526	0.916	0.464	0.992	
Image 31	6.86	9.81	18.07	7.16	33.57	0.513	0.599	0.930	0.522	0.99	
Average	6.33	9.70	16.13	6.42	36.82	0.457	0.535	0.924	0.462	0.997	

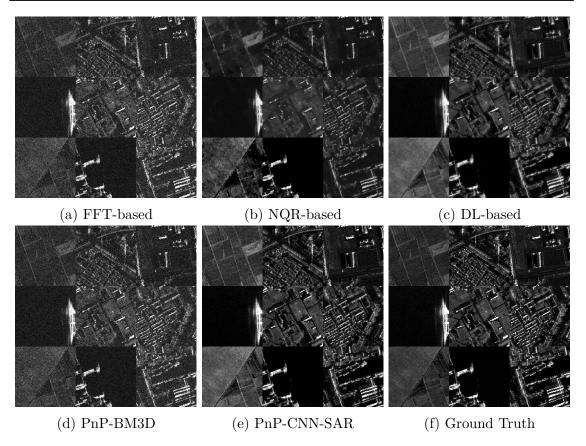


Figure A.19 Reconstruction results for Images 3, 4, 10, 17, 18, 19, 25, 28, and 30 from the Wonsan test set for the case of 100% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$.

Table A.20 SNR and SSIM values for images from the Wonsan Test Set for the case of 90% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			SNR (dE	3)				SSIM		
	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR
Image 1	5.23	9.27	7.59	4.97	17.27	0.317	0.486	0.387	0.309	0.854
Image 2	6.19	6.88	8.78	6.12	15.93	0.457	0.319	0.572	0.447	0.891
Image 3	7.03	7.55	12.23	7.65	15.43	0.377	0.452	0.556	0.391	0.798
Image 4	6.21	8.20	7.97	6.25	14.06	0.535	0.530	0.600	0.537	0.872
Image 5	6.40	8.00	8.42	6.31	15.24	0.524	0.454	0.600	0.512	0.901
Image 6	5.18	8.47	7.25	5.02	17.84	0.372	0.543	0.445	0.363	0.882
Image 7	6.15	7.77	7.71	5.99	13.95	0.519	0.496	0.578	0.511	0.869
Image 8	6.01	8.26	8.03	5.96	14.08	0.483	0.512	0.563	0.478	0.859
Image 9	5.85	7.93	7.76	5.49	14.29	0.472	0.499	0.542	0.449	0.861
Image 10	6.30	9.66	8.28	6.27	16.88	0.385	0.518	0.546	0.391	0.874
Image 11	6.48	8.93	9.34	6.55	15.39	0.459	0.458	0.587	0.461	0.870
Image 12	5.28	8.87	7.55	5.13	17.12	0.350	0.538	0.432	0.347	0.873
Image 13	6.82	11.91	11.90	7.16	16.32	0.327	0.383	0.524	0.337	0.806
Image 14	5.78	7.74	7.34	5.40	15.13	0.500	0.515	0.545	0.472	0.887
Image 15	6.61	8.75	9.51	6.67	15.66	0.468	0.432	0.595	0.469	0.879
Image 16	6.30	7.90	8.46	6.18	14.74	0.520	0.467	0.591	0.509	0.881
Image 17	5.40	12.87	9.17	5.07	14.39	0.187	0.610	0.405	0.193	0.705
Image 18	6.41	7.50	8.91	6.25	16.31	0.443	0.333	0.563	0.429	0.873
Image 19	6.07	7.94	7.57	5.97	15.34	0.515	0.515	0.574	0.501	0.892
Image 20	6.58	9.62	9.73	6.73	15.13	0.460	0.484	0.615	0.463	0.875
Image 21	6.60	10.11	10.22	6.77	15.60	0.424	0.461	0.577	0.423	0.860
Image 22	6.05	7.42	7.80	5.83	15.01	0.497	0.469	0.565	0.483	0.883
Image 23	6.81	10.21	10.71	6.94	15.60	0.425	0.443	0.587	0.427	0.854
Image 24	6.42	9.56	9.46	6.28	16.15	0.424	0.485	0.557	0.412	0.872
Image 25	6.98	11.86	12.35	7.51	16.31	0.329	0.368	0.505	0.336	0.787
Image 26	5.82	9.50	7.13	4.75	17.03	0.393	0.552	0.481	0.360	0.850
Image 27	6.48	8.87	9.30	6.49	15.67	0.450	0.442	0.579	0.444	0.873
Image 28	4.38	11.17	6.34	4.13	21.56	0.191	0.721	0.275	0.196	0.836
Image 29	6.80	10.91	11.16	7.16	16.27	0.405	0.481	0.607	0.414	0.860
Image 30	5.42	8.69	7.01	4.93	16.03	0.434	0.605	0.482	0.412	0.851
Image 31	6.83	9.40	10.87	7.04	15.50	0.476	0.507	0.636	0.483	0.848
Average	6.16	9.09	8.90	6.10	15.85	0.423	0.486	0.538	0.418	0.857

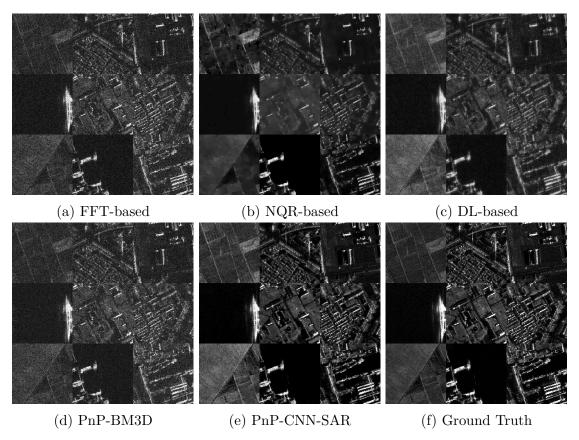


Figure A.20 Reconstruction results for Images 3, 4, 10, 17, 18, 19, 25, 28, and 30 from the Wonsan test set for the case of 90% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$.

A.2.9 Experiment 9: 80% Data Availability and $\sigma_n = \sigma_y$

Table A.21 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR and best SSIM for all 31 images. Fig. A.21 shows the reconstruction results for a selection of images from the Wonsan test set from all methods for this experiment. PnP-CNN-SAR produce better reconstructions compared to other methods.

Table A.21 SNR and SSIM values for images from the Wonsan Test Set for the case of 80% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			SNR (dE	3)		SSIM						
	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR		
Image 1	5.07	7.78	6.94	4.99	12.73	0.290	0.522	0.351	0.293	0.789		
Image 2	6.08	8.01	7.86	6.31	12.85	0.425	0.406	0.493	0.433	0.826		
Image 3	7.04	9.29	11.76	7.83	12.45	0.343	0.430	0.486	0.375	0.653		
Image 4	5.95	6.84	6.92	6.07	12.10	0.491	0.426	0.510	0.501	0.842		
Image 5	6.13	6.79	7.42	6.29	12.41	0.476	0.356	0.510	0.491	0.834		
Image 6	5.04	6.23	6.32	5.04	12.49	0.343	0.460	0.389	0.348	0.825		
Image 7	5.80	7.17	6.80	5.89	12.44	0.472	0.441	0.496	0.481	0.855		
Image 8	5.93	7.22	7.14	5.92	12.19	0.450	0.434	0.491	0.455	0.830		
Image 9	5.71	6.57	6.82	5.68	12.77	0.436	0.404	0.465	0.437	0.828		
Image 10	6.10	7.76	7.35	6.37	12.29	0.369	0.430	0.510	0.389	0.782		
Image 11	6.48	7.98	8.44	6.58	12.41	0.425	0.380	0.505	0.433	0.784		
Image 12	5.19	7.50	6.75	5.22	12.68	0.325	0.407	0.385	0.330	0.822		
Image 13	6.93	11.67	11.48	7.14	13.01	0.300	0.358	0.461	0.304	0.655		
Image 14	5.48	6.75	6.21	5.54	12.40	0.456	0.435	0.454	0.457	0.829		
Image 15	6.37	8.37	8.60	6.58	12.22	0.427	0.381	0.514	0.437	0.787		
Image 16	6.18	7.44	7.45	6.18	12.35	0.481	0.406	0.494	0.478	0.822		
Image 17	5.19	10.91	8.61	5.28	12.78	0.192	0.337	0.414	0.206	0.716		
Image 18	6.20	7.25	8.21	6.47	12.50	0.402	0.285	0.493	0.415	0.781		
Image 19	5.81	7.21	6.59	5.90	12.76	0.468	0.451	0.491	0.476	0.836		
Image 20	6.42	8.56	8.78	6.70	11.93	0.424	0.418	0.535	0.437	0.782		
Image 21	6.47	8.16	9.51	6.72	12.61	0.381	0.245	0.498	0.392	0.760		
Image 22	5.91	6.26	6.74	5.90	12.79	0.466	0.390	0.480	0.464	0.844		
Image 23	6.75	8.16	9.90	6.97	12.44	0.382	0.252	0.500	0.397	0.735		
Image 24	6.26	8.43	8.67	6.47	12.75	0.387	0.406	0.491	0.396	0.776		
Image 25	6.98	10.92	11.97	7.53	12.48	0.296	0.382	0.433	0.310	0.593		
Image 26	5.56	7.12	6.03	5.29	11.99	0.371	0.451	0.429	0.362	0.766		
Image 27	6.36	7.92	8.43	6.57	12.48	0.415	0.372	0.500	0.423	0.779		
Image 28	4.21	9.47	5.78	4.24	13.29	0.185	0.667	0.270	0.197	0.790		
Image 29	6.79	9.80	10.41	7.18	12.60	0.376	0.393	0.529	0.392	0.733		
Image 30	5.16	7.57	6.13	5.21	13.27	0.397	0.435	0.419	0.403	0.813		
Image 31	6.71	8.89	10.21	7.39	12.56	0.443	0.511	0.574	0.476	0.750		
Average	6.01	8.07	8.07	6.18	12.55	0.390	0.409	0.470	0.400	0.781		

A.2.10 Experiment 10: 70% Data Availability and $\sigma_n = \sigma_y$

Table A.22 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR for 28 images, and best SSIM for all 31 images. Fig. A.22 shows the reconstruction results for a selection of images from the Wonsan test set from all methods for this experiment. PnP-CNN-SAR produce better reconstructions compared to other methods.

A.2.11 Experiment 11: 50% Data Availability and $\sigma_n = \sigma_y$

Table A.23 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR for 24 images, and best SSIM for 27 images. Fig. A.23 shows the reconstruction results for a selection of images from the Wonsan test set from all methods for this experiment. PnP-CNN-SAR produce better reconstructions compared to other methods.

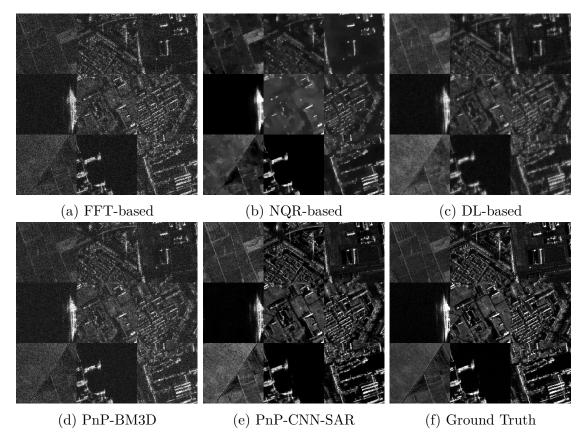


Figure A.21 Reconstruction results for Images 3, 4, 10, 17, 18, 19, 25, 28, and 30 from the Wonsan test set for the case of 80% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$.

Table A.22 SNR and SSIM values for images from the Wonsan Test Set for the case of 70% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			SNR (dE	3)				SSIM		
	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR
Image 1	4.93	7.48	6.44	4.97	9.93	0.272	0.521	0.320	0.280	0.70
Image 2	5.96	6.73	7.30	6.23	10.47	0.397	0.251	0.435	0.411	0.740
Image 3	6.94	8.58	11.12	8.35	10.72	0.317	0.414	0.419	0.358	0.530
Image 4	5.77	6.29	6.41	5.93	9.96	0.463	0.380	0.449	0.465	0.764
Image 5	5.96	6.44	6.83	6.19	10.21	0.439	0.317	0.433	0.454	0.74
Image 6	4.79	6.24	5.91	4.84	9.83	0.317	0.389	0.357	0.328	0.756
Image 7	5.63	6.10	6.35	5.75	9.79	0.445	0.370	0.444	0.453	0.764
Image 8	5.66	6.02	6.66	5.81	9.96	0.416	0.358	0.433	0.424	0.746
Image 9	5.50	6.22	6.31	5.67	10.20	0.405	0.322	0.413	0.412	0.74
Image 10	5.86	6.74	6.66	5.99	10.38	0.345	0.407	0.473	0.388	0.716
Image 11	6.31	5.08	7.85	6.72	10.39	0.399	0.233	0.437	0.408	0.684
Image 12	5.01	6.52	6.29	5.08	10.37	0.302	0.464	0.354	0.316	0.745
Image 13	6.68	11.45	11.00	7.49	11.16	0.262	0.281	0.395	0.285	0.522
Image 14	5.31	5.68	5.80	5.45	10.16	0.425	0.353	0.403	0.426	0.762
Image 15	6.20	7.65	7.95	6.74	10.13	0.393	0.330	0.442	0.416	0.690
Image 16	5.89	6.38	6.94	6.19	10.11	0.433	0.290	0.424	0.451	0.72
Image 17	4.90	10.18	8.29	5.06	10.58	0.188	0.345	0.447	0.232	0.64
Image 18	6.12	4.89	7.61	6.48	10.46	0.378	0.224	0.437	0.392	0.68
Image 19	5.64	6.29	6.11	5.73	10.25	0.437	0.387	0.429	0.441	0.76
Image 20	6.27	7.19	8.20	6.80	10.20	0.392	0.298	0.478	0.418	0.693
Image 21	6.37	8.08	8.90	6.84	10.61	0.348	0.229	0.426	0.366	0.65
Image 22	5.62	5.51	6.25	5.77	10.03	0.429	0.332	0.424	0.432	0.767
Image 23	6.65	7.04	9.31	7.15	10.45	0.350	0.209	0.426	0.366	0.617
Image 24	6.16	7.32	8.04	6.45	10.80	0.357	0.260	0.432	0.368	0.684
Image 25	6.89	10.50	11.50	8.19	10.99	0.267	0.370	0.376	0.300	0.469
Image 26	5.29	6.12	5.59	5.40	9.84	0.346	0.414	0.398	0.359	0.703
Image 27	6.23	7.80	7.74	6.59	10.36	0.385	0.341	0.434	0.395	0.675
Image 28	4.15	7.92	5.30	3.95	10.36	0.180	0.651	0.276	0.201	0.73
Image 29	6.70	9.97	9.85	7.57	10.81	0.343	0.374	0.478	0.374	0.640
Image 30	5.04	6.19	5.69	5.13	10.14	0.371	0.395	0.375	0.380	0.754
Image 31	6.59	8.21	9.70	7.53	10.76	0.423	0.508	0.526	0.457	0.65
Average	5.84	7.19	7.54	6.19	10.34	0.362	0.355	0.419	0.379	0.693

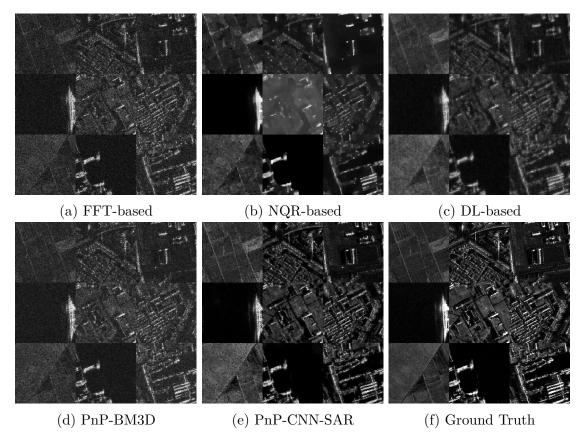


Figure A.22 Reconstruction results for Images 3, 4, 10, 17, 18, 19, 25, 28, and 30 from the Wonsan test set for the case of 70% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$.

Table A.23 SNR and SSIM values for images from the Wonsan Test Set for the case of 50% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			SNR (dE	3)				SSIM		
	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR
Image 1	4.59	6.24	5.57	4.71	8.08	0.232	0.287	0.276	0.255	0.606
Image 2	5.53	5.14	6.32	6.00	8.26	0.339	0.201	0.333	0.346	0.606
Image 3	6.50	8.83	9.39	9.51	9.28	0.274	0.380	0.309	0.338	0.375
Image 4	5.29	5.68	5.57	5.49	7.74	0.395	0.333	0.344	0.383	0.642
Image 5	5.48	4.09	5.86	5.89	7.85	0.379	0.188	0.315	0.359	0.598
Image 6	4.59	5.29	5.10	4.37	7.67	0.288	0.268	0.300	0.283	0.648
Image 7	5.14	5.82	5.41	5.17	7.75	0.378	0.346	0.333	0.356	0.647
Image 8	5.19	5.98	5.74	5.39	7.88	0.358	0.333	0.335	0.345	0.637
Image 9	5.06	5.20	5.42	5.35	8.02	0.349	0.245	0.317	0.354	0.659
Image 10	5.36	5.83	5.76	5.73	8.13	0.334	0.386	0.425	0.380	0.611
Image 11	5.69	6.13	6.78	6.67	8.26	0.329	0.236	0.321	0.348	0.543
Image 12	4.69	5.99	5.59	4.66	8.11	0.262	0.379	0.310	0.282	0.629
Image 13	6.37	11.19	9.27	8.44	9.70	0.224	0.275	0.293	0.264	0.378
Image 14	4.89	5.34	5.00	4.99	7.85	0.356	0.318	0.302	0.354	0.680
Image 15	5.85	6.66	6.88	6.73	8.41	0.343	0.243	0.332	0.340	0.553
Image 16	5.45	5.65	5.97	5.92	7.92	0.366	0.192	0.299	0.360	0.599
Image 17	4.72	9.07	7.27	4.59	8.66	0.228	0.335	0.490	0.297	0.704
Image 18	5.66	6.50	6.62	6.52	8.47	0.318	0.226	0.328	0.329	0.551
Image 19	5.17	4.82	5.25	5.42	7.91	0.368	0.239	0.323	0.369	0.650
Image 20	5.77	7.42	7.07	6.83	8.10	0.344	0.351	0.390	0.364	0.556
Image 21	5.97	7.00	7.74	7.19	8.71	0.301	0.249	0.319	0.321	0.504
Image 22	5.19	5.87	5.38	5.40	7.90	0.374	0.328	0.317	0.363	0.656
Image 23	6.17	8.53	7.97	7.72	8.60	0.288	0.243	0.305	0.328	0.463
Image 24	5.67	6.20	7.01	6.53	8.42	0.302	0.214	0.339	0.340	0.561
Image 25	6.54	9.95	9.72	9.56	9.54	0.233	0.320	0.266	0.280	0.317
Image 26	4.81	5.16	4.90	4.93	2.99	0.309	0.350	0.355	0.343	0.254
Image 27	5.73	6.53	6.75	6.62	8.28	0.318	0.224	0.332	0.339	0.529
Image 28	3.98	6.94	4.77	3.48	6.68	0.178	0.647	0.294	0.215	0.604
Image 29	6.12	7.93	8.38	8.27	8.87	0.288	0.261	0.381	0.346	0.504
Image 30	4.66	4.71	4.92	4.77	7.85	0.318	0.235	0.288	0.324	0.650
Image 31	6.21	9.07	8.30	7.85	9.03	0.390	0.493	0.440	0.448	0.518
Average	5.42	6.60	6.51	6.15	8.09	0.315	0.301	0.333	0.334	0.562

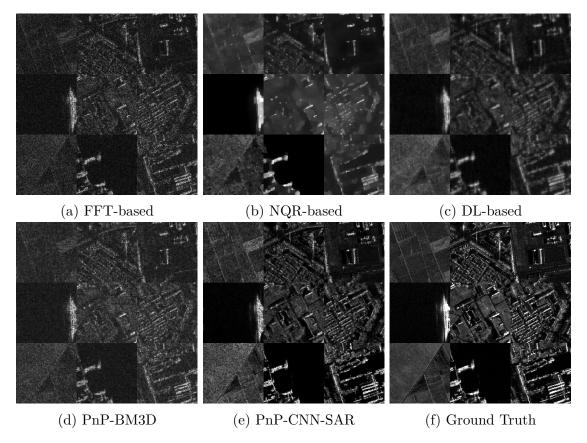


Figure A.23 Reconstruction results for Images 3, 4, 10, 17, 18, 19, 25, 28, and 30 from the Wonsan test set for the case of 50% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$.

A.2.12 Experiment 12: 30% Data Availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$

Table A.24 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR for 14 images, and best SSIM for 25 images. Fig. A.24 shows the reconstruction results for a selection of images from the Wonsan test set from all methods for this experiment. PnP-CNN-SAR produce better reconstructions compared to other methods.

A.3 Different Scene Experiments

A.3.1 Experiment 1: 100% Data Availability and $\sigma_n = 0.1\sigma_y$

Table A.25 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR and best SSIM for all 110 images. Fig. A.25 shows the reconstruction results for a selection of images from the Kapıkule test set from all methods for this experiment. All methods produce good reconstructions.

Table A.24 SNR and SSIM values for images from the Wonsan Test Set for the case of 30% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			SNR (dE	3)				SSIM		
	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR
Image 1	4.04	5.32	4.55	4.70	5.61	0.203	0.307	0.245	0.232	0.410
Image 2	4.60	5.73	4.94	5.66	5.94	0.269	0.199	0.231	0.264	0.368
Image 3	5.51	7.28	6.99	9.30	7.12	0.249	0.313	0.230	0.338	0.26
Image 4	4.32	5.32	4.36	4.98	5.34	0.311	0.295	0.249	0.261	0.373
Image 5	4.49	4.31	4.67	5.35	5.60	0.291	0.144	0.208	0.239	0.359
Image 6	3.86	5.27	4.21	4.28	4.32	0.248	0.334	0.261	0.241	0.374
Image 7	4.30	5.08	4.31	4.81	5.10	0.307	0.291	0.246	0.252	0.352
Image 8	4.32	5.33	4.57	5.09	5.34	0.287	0.291	0.249	0.267	0.359
Image 9	4.30	4.61	4.38	4.87	5.43	0.281	0.273	0.233	0.255	0.382
Image 10	4.47	5.69	4.54	5.56	5.57	0.310	0.379	0.368	0.379	0.451
Image 11	4.85	6.82	5.29	6.23	6.19	0.269	0.275	0.226	0.259	0.334
Image 12	4.13	5.17	4.40	4.66	4.83	0.238	0.389	0.265	0.248	0.374
Image 13	5.33	9.98	7.02	8.30	7.22	0.187	0.237	0.210	0.252	0.246
Image 14	4.08	4.93	4.04	4.56	5.17	0.284	0.281	0.216	0.259	0.391
Image 15	4.83	6.47	5.29	6.33	6.10	0.274	0.207	0.233	0.257	0.328
Image 16	4.48	5.31	4.75	5.40	5.67	0.277	0.175	0.200	0.232	0.336
Image 17	4.24	8.94	5.62	5.21	6.86	0.297	0.360	0.538	0.419	0.580
Image 18	4.74	5.42	5.13	6.02	5.93	0.259	0.162	0.230	0.256	0.322
Image 19	4.33	5.00	4.23	5.00	5.29	0.295	0.280	0.228	0.265	0.374
Image 20	4.83	6.70	5.48	6.53	6.38	0.292	0.292	0.298	0.296	0.368
Image 21	4.99	7.33	5.85	6.71	6.32	0.232	0.187	0.216	0.235	0.304
Image 22	4.31	5.09	4.26	4.98	5.22	0.295	0.278	0.227	0.269	0.377
Image 23	5.21	6.84	6.07	7.15	6.72	0.236	0.162	0.203	0.229	0.292
Image 24	4.87	6.58	5.31	6.15	6.01	0.252	0.225	0.249	0.267	0.355
Image 25	5.56	7.85	7.17	9.15	7.25	0.201	0.263	0.188	0.281	0.194
Image 26	4.09	4.61	3.91	4.63	5.17	0.276	0.323	0.304	0.327	0.431
Image 27	4.84	6.32	5.31	6.25	6.03	0.267	0.211	0.236	0.256	0.341
Image 28	3.47	5.64	3.76	3.72	5.76	0.211	0.608	0.326	0.246	0.519
Image 29	5.27	8.93	6.38	7.84	7.29	0.248	0.321	0.292	0.298	0.33
Image 30	4.00	4.53	3.97	4.54	5.07	0.264	0.286	0.225	0.264	0.370
Image 31	5.09	7.87	6.28	8.12	6.65	0.352	0.462	0.369	0.440	0.389
Average	4.57	6.14	5.06	5.87	5.89	0.267	0.284	0.258	0.277	0.363

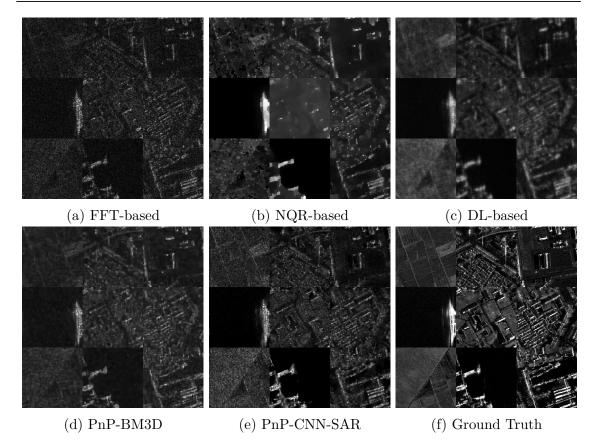


Figure A.24 Reconstruction results for Images 3, 4, 10, 17, 18, 19, 25, 28, and 30 from the Wonsan test set for the case of 30% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$.

Table A.25 SNR and SSIM values for images from the Kapıkule Test Set for the case of 100% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

SNR (dB)	SSIM
FFT-based NQR-based DL-based PnP-BM3D PnP-CNN-SAR	FFT-based NQR-based DL-based PnP-BM3D PnP-CNN-SAR

Table A.25 continued from previous page

			Ta	ble A.25 co	ntinued from	previous p	age			
			SNR (dB)					SSIM		
	06.70	10.46	05.45	00.50	90.07	0.001	0.000	0.000	0.004	0.999
Image 1 Image 2	$\frac{26.72}{26.46}$	10.46 26.22	$25.45 \\ 24.76$	26.56 26.61	$38.87 \\ 36.84$	$0.991 \\ 0.987$	$0.823 \\ 0.973$	$0.982 \\ 0.978$	0.984 0.974	0.996
Image 3	26.69	14.24	25.84	26.69	34.78	0.993	0.917	0.985	0.987	0.998
Image 4	26.69	26.34	24.16	26.72	37.22	0.992	0.985	0.984	0.985	0.999
Image 5	26.62	23.93	22.61	26.58	39.61	0.987	0.974	0.977	0.973	0.998
Image 6	26.68	18.46	26.26	26.67	39.05	0.991	0.941	0.983	0.983	0.999
Image 7 Image 8	$\frac{26.81}{26.63}$	25.18 25.45	24.43 23.84	26.71 26.73	$37.30 \\ 39.35$	$0.989 \\ 0.991$	$0.979 \\ 0.982$	0.982 0.984	0.979 0.983	0.999 0.999
Image 9	26.68	25.43 25.23	22.91	26.65	39.87	0.987	0.982	0.977	0.983	0.997
Image 10	26.69	26.61	25.08	26.65	41.54	0.986	0.974	0.979	0.973	0.997
Image 11	26.67	26.44	25.91	26.66	39.57	0.990	0.981	0.982	0.981	0.999
Image 12	26.68	24.97	21.11	26.75	38.48	0.989	0.979	0.978	0.979	0.998
Image 13 Image 14	$\frac{26.47}{26.77}$	$\frac{26.49}{25.98}$	$24.30 \\ 24.93$	$26.45 \\ 26.65$	$39.95 \\ 37.79$	$0.983 \\ 0.991$	$0.967 \\ 0.982$	$0.975 \\ 0.983$	$0.967 \\ 0.982$	$0.996 \\ 0.999$
Image 14 Image 15	26.68	25.99	25.27	26.69	36.78	0.993	0.982	0.986	0.982	0.999
Image 16	26.16	25.31	21.18	26.22	39.01	0.963	0.937	0.955	0.934	0.993
Image 17	26.78	12.41	26.30	26.64	36.99	0.991	0.902	0.983	0.985	0.998
Image 18	26.71	26.53	25.33	26.70	36.79	0.993	0.986	0.985	0.986	0.999
Image 19	26.66	25.79	23.87	26.72	39.39	0.994	0.988	0.986	0.989	0.999
Image 20 Image 21	26.61 26.47	25.67 18.22	22.84 24.73	$26.68 \\ 26.47$	$37.68 \\ 38.78$	$0.991 \\ 0.983$	$0.982 \\ 0.956$	0.983 0.974	0.982 0.971	$0.998 \\ 0.997$
Image 21	26.43	26.16	24.45	26.46	39.77	0.933	0.960	0.974	0.959	0.996
Image 23	26.58	23.45	25.47	26.65	41.90	0.989	0.977	0.981	0.981	0.999
Image 24	26.69	26.70	24.78	26.75	36.22	0.993	0.987	0.986	0.987	0.999
Image 25	26.75	12.56	27.02	26.66	34.80	0.990	0.900	0.981	0.982	0.997
Image 26	26.70	26.68	24.69	26.71	40.17	0.992	0.986	0.985	0.985	0.999
Image 27 Image 28	26.64 26.63	20.15 23.47	25.80 26.15	$26.68 \\ 26.70$	$35.81 \\ 36.99$	$0.993 \\ 0.991$	$0.980 \\ 0.981$	0.985 0.983	0.987 0.983	0.998 0.998
Image 28 Image 29	26.69	26.58	24.69	26.59	44.13	0.991 0.992	0.985	0.983	0.984	1.000
Image 30	26.72	17.05	25.39	26.71	39.42	0.992	0.969	0.983	0.985	0.999
Image 31	26.65	26.64	25.12	26.73	40.05	0.995	0.989	0.986	0.989	0.999
Image 32	26.70	12.57	25.34	26.68	36.69	0.993	0.901	0.984	0.988	0.999
Image 33	26.51	16.85 26.28	25.40 24.28	$26.55 \\ 26.34$	36.82	0.985	0.954	0.976	0.974	0.996 0.996
Image 34 Image 35	26.31 26.69	20.28 20.67	24.28 25.32	26.62	$40.94 \\ 43.21$	$0.972 \\ 0.993$	0.949 0.980	$0.964 \\ 0.985$	0.948 0.987	1.000
Image 36	26.53	26.52	25.02	26.53	39.51	0.985	0.971	0.976	0.972	0.998
Image 37	26.77	13.17	25.57	26.70	37.05	0.993	0.912	0.984	0.987	0.999
Image 38	26.62	13.18	26.04	26.66	36.24	0.991	0.911	0.982	0.984	0.998
Image 39	26.70	10.07	25.61	26.76	36.33	0.992	0.877	0.983	0.987	0.998
Image 40 Image 41	26.64 26.61	25.20 26.39	23.78 25.53	$26.67 \\ 26.68$	$37.94 \\ 36.59$	$0.989 \\ 0.992$	0.978 0.984	0.981 0.985	0.979 0.984	0.999 0.998
Image 41 Image 42	26.82	16.63	25.59	26.73	38.63	0.992	0.938	0.984	0.986	0.999
Image 43	26.40	23.49	24.88	26.39	40.50	0.976	0.951	0.969	0.956	0.996
Image 44	26.73	26.45	24.96	26.68	36.94	0.994	0.988	0.986	0.988	0.999
Image 45	26.61	25.75	22.87	26.72	39.01	0.992	0.983	0.983	0.984	0.999
Image 46 Image 47	$26.54 \\ 26.75$	24.37 26.47	20.98 24.68	26.53 26.68	$39.67 \\ 37.42$	$0.983 \\ 0.994$	$0.970 \\ 0.988$	$0.973 \\ 0.986$	$0.967 \\ 0.988$	$0.996 \\ 0.999$
Image 47 Image 48	26.60	26.60	24.44	26.64	40.76	0.989	0.988	0.980	0.988	0.998
Image 49	26.72	19.72	26.93	26.65	35.35	0.991	0.974	0.983	0.982	0.998
Image 50	26.49	26.08	24.26	26.46	39.68	0.982	0.965	0.973	0.964	0.996
Image 51	26.69	26.11	23.78	26.65	38.92	0.993	0.986	0.985	0.986	0.999
Image 52	26.53	26.42	25.85	26.68	35.85	0.988	0.975	0.979	0.976	0.995
Image 53 Image 54	26.67 26.63	25.86 26.65	22.79 25.45	$\frac{26.73}{26.55}$	$\frac{38.64}{37.02}$	$0.991 \\ 0.988$	0.982 0.977	0.983 0.980	0.982 0.977	0.999 0.997
Image 55	26.61	26.63	26.06	26.64	40.94	0.989	0.980	0.982	0.979	0.999
Image 56	26.63	26.37	24.77	26.70	39.96	0.992	0.985	0.984	0.985	0.999
Image 57	26.70	26.26	23.52	26.68	38.81	0.993	0.986	0.985	0.986	0.999
Image 58	26.51	25.16	25.15	26.55	40.82	0.985	0.969	0.977	0.971	0.998
Image 59 Image 60	$\frac{26.35}{26.70}$	26.32 26.16	$24.56 \\ 24.74$	$\frac{26.32}{26.74}$	$37.70 \\ 38.12$	$0.976 \\ 0.993$	$0.955 \\ 0.985$	$0.968 \\ 0.985$	$0.955 \\ 0.985$	$0.996 \\ 0.999$
Image 61	26.57	25.61	23.44	26.58	38.34	0.990	0.980	0.982	0.980	0.998
Image 62	26.59	18.74	25.54	26.68	39.29	0.993	0.948	0.984	0.986	0.999
Image 63	26.69	14.67	25.83	26.67	35.83	0.992	0.907	0.985	0.986	0.998
Image 64	26.65	26.10	24.36	26.71	37.95	0.991	0.983	0.983	0.983	0.999
Image 65	26.74	25.08	24.06	26.75 26.73	37.34	0.990	0.980	0.982	0.981	0.998
Image 66 Image 67	26.68 26.33	25.22 25.40	23.09 24.34	$\frac{26.73}{26.32}$	$39.78 \\ 42.47$	$0.991 \\ 0.972$	$0.982 \\ 0.953$	$0.983 \\ 0.965$	0.982 0.951	0.999 0.997
Image 68	26.64	26.47	24.96	26.72	37.17	0.993	0.986	0.985	0.987	0.999
Image 69	26.68	20.18	25.97	26.70	39.28	0.992	0.977	0.984	0.984	0.999
Image 70	26.65	16.99	25.81	26.71	37.20	0.993	0.938	0.984	0.986	0.999
Image 71	26.45	25.50	24.83 23.19	26.54	$\frac{38.15}{37.06}$	0.986	0.972	0.977	0.972	0.997
Image 72 Image 73	26.70 26.65	24.93 26.12	23.19	$26.74 \\ 26.65$	37.47	$0.988 \\ 0.994$	0.977 0.989	0.981 0.986	0.977 0.989	0.999 0.999
Image 73	26.68	26.69	25.88	26.66	37.41	0.991	0.982	0.982	0.982	0.998
Image 75	26.67	14.79	26.05	26.70	36.54	0.992	0.913	0.983	0.985	0.998
Image 76	26.71	26.46	24.83	26.65	37.76	0.993	0.987	0.985	0.986	0.999
Image 77	26.71	17.65	26.04	26.63	37.92	0.991	0.940	0.983	0.984	0.999
Image 78	26.68	26.44	23.97	26.64	38.64	0.994	0.988	$0.985 \\ 0.975$	0.988	0.999 0.996
Image 79 Image 80	$26.52 \\ 26.64$	$26.50 \\ 26.33$	25.00 24.63	26.50 26.63	$37.56 \\ 40.29$	$0.984 \\ 0.993$	$0.969 \\ 0.985$	$0.975 \\ 0.984$	$0.969 \\ 0.985$	0.996
Image 80 Image 81	26.66	25.53	25.46	26.74	36.28	0.989	0.978	0.981	0.98	0.998
Image 82	26.64	25.31	21.54	26.61	38.75	0.986	0.974	0.975	0.973	0.997
Image 83	26.75	14.14	25.12	26.67	37.40	0.993	0.911	0.984	0.987	0.999
Image 84	26.62	26.62	24.67	26.62	38.59	0.988	0.977	0.981	0.977	0.997
Image 85	26.68	25.72	22.47	26.70	38.35	0.992	0.984	0.983	0.984	0.998
Image 86 Image 87	26.58 26.50	25.23 24.99	22.01 22.62	26.69 26.48	$41.08 \\ 41.37$	$0.991 \\ 0.987$	0.982 0.975	0.981 0.977	0.982 0.974	0.998 0.998
Image 88	26.72	22.94	25.82	26.74	36.90	0.992	0.982	0.984	0.985	0.998
Image 89	26.66	26.60	25.02	26.46	43.24	0.992	0.985	0.984	0.984	0.999
Image 90	26.64	14.80	25.46	26.77	36.87	0.993	0.961	0.985	0.986	0.999
Image 91	26.72	25.26	23.65	26.66	38.79	0.990	0.981	0.983	0.981	0.999
Image 92 Image 93	26.63 26.67	26.21 25.33	25.20 22.82	26.62 26.75	$37.44 \\ 38.17$	$0.993 \\ 0.992$	$0.985 \\ 0.984$	0.985 0.983	0.985 0.984	0.999 0.999
Image 93 Image 94	$\frac{26.67}{26.72}$	26.46	24.95	26.69	39.76	0.992	0.984	0.985	0.986	0.999
Image 95	26.64	26.57	25.50	26.72	37.54	0.993	0.986	0.985	0.986	0.999
Image 96	26.57	25.38	24.01	26.56	43.27	0.989	0.980	0.981	0.980	0.999
					•					

			Tal	ole A.25 con	tinued from	previous p	age			
_			$\mathrm{SNR}\ (\mathrm{dB})$		1					
Image 97	26.46	24.06	24.62	26.39	39.21	0.980	0.960	0.972	0.964	0.996
Image 98	26.69	26.64	26.12	26.70	37.58	0.992	0.985	0.985	0.985	0.999
Image 99	26.73	19.28	25.60	26.70	39.34	0.993	0.977	0.984	0.986	0.999
Image 100	26.27	24.80	20.57	26.27	38.81	0.966	0.941	0.955	0.938	0.994
Image 101	26.65	17.38	25.37	26.67	38.56	0.992	0.970	0.984	0.985	0.999
Image 102	26.73	16.12	25.52	26.72	39.18	0.992	0.904	0.984	0.985	0.999
Image 103	26.70	17.41	27.39	26.71	35.63	0.988	0.937	0.979	0.978	0.997
Image 104	26.76	10.18	25.18	26.70	37.55	0.994	0.796	0.985	0.989	0.999
Image 105	26.39	26.41	24.22	26.31	40.49	0.971	0.950	0.963	0.948	0.995
Image 106	26.74	26.38	24.07	26.75	39.32	0.991	0.983	0.984	0.983	0.999
Image 107	26.68	26.67	24.56	26.61	42.85	0.993	0.987	0.984	0.986	1.000
Image 108	26.67	26.63	25.01	26.66	37.19	0.993	0.986	0.985	0.986	0.999
Image 109	26.78	26.31	23.43	26.67	37.54	0.992	0.984	0.984	0.984	0.999
Image 110	26.68	17.05	27.11	26.73	36.13	0.989	0.931	0.980	0.981	0.998
Average	26.63	23.11	24.66	26.63	38.52	0.989	0.964	0.981	0.979	0.998

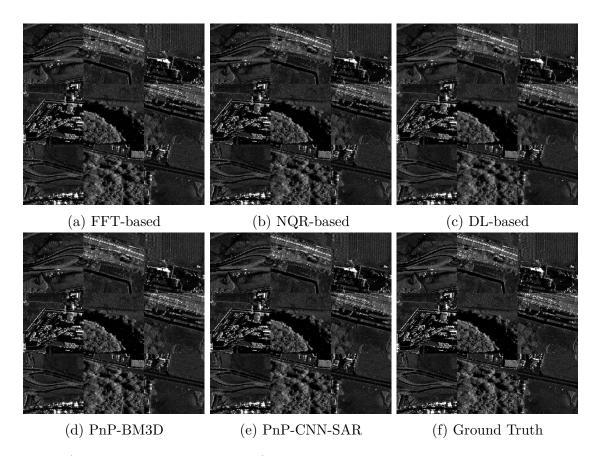


Figure A.25 Reconstruction results for Images 19, 40, 45, 46, 59, 61, 66, 96, and 109 from the Kapıkule test set for the case of 100% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$.

A.3.2 Experiment 2: 90% Data Availability and $\sigma_n = 0.1\sigma_y$

Table A.26 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR for 106 images, and best SSIM for 105 images. Fig. A.26 shows the reconstruction results for a selection of images from the Kapıkule test set from all methods for this experiment. All methods except the NQR-based produce good reconstructions for all images.

Table A.26 SNR and SSIM values for images from the Kapıkule Test Set for the case of 90% data availability and $\sigma_{\bf n}=0.1\sigma_{\bf y}.$ Best results are shown in bold, second best results are shown in red.

)		SSIM					
	FFT-based NQ	R-based D	L-based P	nP-BM3D F	nP-CNN-SAR	FFT-based	NQR-based I	OL-based I	PnP-BM3D P	nP-CNN-SAI
mage 1	13.29	8.16	15.29	15.57	16.67	0.756	0.627	0.837	0.853	0.87
mage 2 mage 3	13.22 13.10	11.45 10.33	$14.22 \\ 15.79$	$\frac{14.48}{15.48}$	16.90 16.60	$0.762 \\ 0.767$	$0.528 \\ 0.755$	$0.815 \\ 0.860$	0.814 0.839	0.90 0.86
mage 4	13.10	12.48	14.01	14.62	16.56	0.766	0.790	0.843	0.845	0.89
mage 5	13.11	11.20	11.78	14.62	15.48	0.780	0.585	0.755	0.840	0.85
mage 6	13.17	11.77	15.93	17.16	17.52	0.711	0.575	0.834	0.867	0.87
mage 7	13.39	12.08	14.42	15.74	17.36	0.744	0.779	0.835	0.855	0.88
mage 8	13.37	9.71	13.46	15.39	16.64	0.767	0.435	0.830	0.857	0.88
mage 9	12.85	8.43	11.68	15.08	17.59	0.760	0.444	0.776	0.851	0.91
mage 10	13.18	12.56	14.51	15.68	16.73	0.780	0.628	0.818	0.843	0.91
mage 11 mage 12	13.29 12.99	12.23 10.86	15.72 9.75	$\frac{16.45}{12.74}$	17.24 17.42	0.751 0.813	0.758 0.808	$0.840 \\ 0.740$	$0.854 \\ 0.838$	$0.89 \\ 0.92$
mage 13	13.11	13.03	13.51	14.22	17.00	0.756	0.768	0.740	0.810	0.92
mage 14	13.43	11.75	15.07	16.48	16.79	0.739	0.768	0.839	0.867	0.87
mage 15	13.24	13.04	15.08	15.71	16.76	0.767	0.801	0.863	0.859	0.88
mage 16	12.74	11.59	9.81	12.57	17.48	0.746	0.618	0.635	0.753	0.94
mage 17	13.35	8.53	16.07	16.75	17.33	0.727	0.583	0.838	0.861	0.85
mage 18	13.33	12.49	15.34	16.30	16.65	0.758	0.790	0.862	0.862	0.88
mage 19	13.40	13.29	13.54	14.74	16.22	0.811	0.834	0.845	0.857	0.90
mage 20	13.27	9.96	12.07	14.64	17.28	0.775	0.525	0.811	0.855	0.90
mage 21	13.10	11.41	14.19	14.46	17.27	0.749	0.750	0.801	0.814	0.90
mage 22	13.10 13.23	13.12 12.13	13.87 15.05	14.29 15.98	16.96 16.79	0.753 0.761	$0.772 \\ 0.687$	$0.790 \\ 0.831$	$0.808 \\ 0.859$	$0.91 \\ 0.88$
mage 23 mage 24	13.38	12.13	14.32	14.71	16.59	0.772	0.807	0.858	0.839	0.88
mage 24	13.38	10.15	16.89	$14.71 \\ 17.79$	17.42	0.772	0.527	0.829	0.860	0.83
mage 26	13.38	13.39	14.02	14.77	16.52	0.788	0.821	0.847	0.849	0.90
mage 27	13.27	11.72	15.91	15.37	16.83	0.764	0.766	0.862	0.837	0.88
mage 28	13.25	11.87	15.96	16.84	17.02	0.720	0.528	0.841	0.861	0.87
mage 29	13.26	13.28	14.30	14.97	16.04	0.791	0.816	0.836	0.858	0.90
mage 30	13.35	12.03	14.86	15.70	16.88	0.759	0.772	0.836	0.854	0.88
mage 31	13.27	13.30	15.03	14.92	15.79	0.789	0.819	0.862	0.853	0.88
mage 32	13.27	9.92	15.13	14.62	16.33	0.778	0.598	0.849	0.835	0.87
mage 33	13.33	12.99	14.89	15.05	17.29	0.727	0.760	0.812	0.823	0.89
mage 34	13.26	12.88	13.62	13.81	16.76	0.747	0.755	0.780	0.790	0.92 0.90
mage 35 mage 36	13.41 13.14	12.45 12.19	14.96 14.05	15.63 15.12	16.56 17.49	0.779 0.739	$0.795 \\ 0.765$	0.847 0.801	$0.864 \\ 0.826$	0.90
nage 37	13.24	9.68	15.40	15.33	16.49	0.741	0.611	0.837	0.848	0.86
nage 38	13.41	9.07	15.98	16.49	16.99	0.732	0.542	0.841	0.850	0.85
nage 39	13.27	8.19	15.66	15.38	16.69	0.743	0.594	0.843	0.847	0.85
nage 40	13.19	10.75	13.91	15.19	16.58	0.772	0.509	0.826	0.852	0.89
mage 41	13.20	10.21	15.11	16.13	16.88	0.748	0.457	0.849	0.858	0.88
mage 42	13.51	11.32	15.46	16.00	16.67	0.746	0.626	0.845	0.861	0.87
mage 43	13.19	12.83	14.46	14.74	16.86	0.744	0.755	0.797	0.806	0.91
mage 44	13.29	13.11	14.81	14.28	15.96	0.808	0.821	0.868	0.827	0.89
mage 45	13.39	10.41	12.06	14.48	16.46	0.786	0.796	0.800	0.845	0.90
mage 46 mage 47	12.76 13.27	9.80 12.12	9.30 14.65	13.56 15.20	18.52 16.48	0.776 0.779	$0.605 \\ 0.807$	0.694 0.859	$\frac{0.829}{0.858}$	0.92 0.89
nage 47	13.14	13.18	13.88	13.20 14.27	16.17	0.779	0.809	0.839 0.837	0.831	0.90
mage 49	13.15	11.61	16.77	17.83	17.80	0.700	0.503	0.842	0.862	0.87
mage 50	12.92	12.10	13.45	14.17	17.06	0.762	0.776	0.780	0.794	0.91
nage 51	13.28	12.98	13.42	13.69	16.55	0.775	0.809	0.841	0.846	0.90
nage 52	13.15	12.92	15.56	16.45	17.85	0.713	0.723	0.823	0.848	0.88
nage 53	13.26	12.00	11.91	14.40	16.61	0.774	0.798	0.812	0.850	0.89
nage 54	13.06	13.10	14.90	15.83	17.50	0.721	0.743	0.817	0.839	0.89
nage 55	13.28	13.30	15.73	16.49	17.35	0.750	0.776	0.837	0.853	0.89
nage 56	13.10	12.91	14.54	15.46	16.45	0.774	0.794	0.849	0.863	0.89
nage 57 nage 58	13.30 13.06	11.80 12.47	13.12 14.75	14.86 15.45	16.60 17.18	0.780 0.747	$0.803 \\ 0.621$	0.833 0.817	$0.853 \\ 0.835$	0.89
nage 59	12.92	12.47 12.44	13.85	14.63	17.18	0.747	0.734	0.817	0.805	0.9
nage 60	13.31	11.98	14.26	15.46	16.24	0.758	0.786	0.847	0.862	0.88
nage 61	13.34	11.26	13.38	14.87	16.20	0.783	0.795	0.816	0.850	0.89
age 62	13.22	11.40	15.25	15.72	16.63	0.760	0.665	0.841	0.856	0.8
nage 63	13.30	9.25	15.57	15.64	16.81	0.748	0.727	0.850	0.848	0.8
age 64	13.25	13.08	13.92	15.04	17.02	0.761	0.792	0.828	0.837	0.8
nage 65	13.44	10.17	13.71	15.02	17.27	0.734	0.461	0.820	0.849	0.8
age 66	13.06	8.90	12.51	15.12	16.84	0.790	0.446	0.821	0.862	0.9
nage 67	12.90 13.32	12.92 13.30	13.54 14.71	14.43 15.02	16.99 15.85	0.759	$0.761 \\ 0.814$	$0.775 \\ 0.863$	$\frac{0.811}{0.854}$	0.93 0.88
age 68	13.32	11.23	14.71 15.60	16.60	17.03	0.788 0.741	0.814 0.575	0.863 0.843	$\frac{0.854}{0.866}$	0.8
age 70	13.36	9.79	15.65	15.89	17.13	0.741	0.611	0.842	0.856	0.8
age 70	13.00	12.62	13.99	15.47	17.54	0.731	0.739	0.793	0.833	0.9
age 72	13.31	11.56	12.98	14.88	16.06	0.756	0.772	0.818	0.858	0.8
age 73	13.30	12.14	12.05	13.05	16.52	0.832	0.845	0.823	0.823	0.9
nage 74	13.33	13.31	15.57	16.99	17.69	0.720	0.743	0.835	0.863	0.8
age 75	13.28	9.83	15.84	16.13	16.72	0.747	0.737	0.850	0.854	0.8
age 76	13.37	12.92	14.42	15.21	16.59	0.768	0.802	0.851	0.846	0.8
age 77	13.28	11.45	15.84	16.90	17.30	0.714	0.590	0.836	0.869	0.8
nage 78	13.22	13.22	13.74	14.61	16.08	0.794	0.820	0.839	0.852	0.8
nage 79	13.00	12.56	14.28	15.46	17.25	0.733	0.743	0.800	0.835	0.8
nage 80	13.16	12.52	14.25	15.12	16.39	0.777	0.801	0.837	0.848	0.8
nage 81	13.17	12.02	14.65	16.99	18.53	0.685	0.438	0.831	0.860	0.8
nage 82	12.95	10.76	9.96	13.34	17.25	0.793	0.543	0.735	0.823	0.9
nage 83	13.34	10.77 13.39	14.83 14.00	14.89 15.05	$16.30 \\ 16.84$	0.777	0.781	$0.849 \\ 0.820$	$0.839 \\ 0.823$	0.80
nage 84 nage 85	13.35 13.07	11.23	14.00 11.27	15.05 14.35	16.84	0.773 0.785	$0.794 \\ 0.793$	0.820	0.823 0.846	0.89
nage 86	12.97	10.80	10.80	13.50	16.81	0.783	0.793	0.760	0.846 0.835	0.83
nage 87	13.00	10.54	11.75	14.32	16.74	0.779	0.599	0.763	0.833	0.93
	13.31	11.49	15.77	16.24	17.20		0.538	0.846	0.859	0.88

			Ta	ble A.26 con	tinued from	previous p	age			
_			SNR (dB)		1			SSIM		
Image 89	13.22	12.74	14.43	15.36	16.58	0.781	0.798	0.830	0.858	0.908
Image 90	13.23	10.00	15.28	15.35	16.60	0.754	0.582	0.849	0.847	0.879
Image 91	13.31	11.34	13.22	15.59	16.80	0.765	0.779	0.822	0.857	0.888
Image 92	13.18	12.36	15.33	15.73	17.13	0.764	0.784	0.862	0.852	0.889
Image 93	13.08	10.28	12.31	14.03	16.40	0.801	0.796	0.811	0.845	0.898
Image 94	13.36	12.14	14.71	15.64	16.58	0.763	0.797	0.847	0.863	0.892
Image 95	13.18	13.22	15.42	15.54	16.70	0.766	0.791	0.865	0.856	0.890
Image 96	13.02	11.83	13.49	14.93	17.13	0.778	0.793	0.814	0.858	0.919
Image 97	12.97	12.56	14.02	14.58	16.95	0.745	0.756	0.796	0.807	0.908
Image 98	13.21	12.84	15.93	16.73	17.00	0.729	0.749	0.850	0.862	0.877
Image 99	13.28	12.60	15.28	15.85	16.79	0.766	0.783	0.846	0.857	0.889
Image 100	12.90	11.15	8.69	12.00	17.50	0.746	0.664	0.627	0.769	0.941
Image 101	13.26	12.11	15.04	15.24	16.47	0.771	0.779	0.842	0.839	0.877
Image 102	13.29	10.68	15.20	15.57	16.86	0.762	0.752	0.843	0.847	0.882
Image 103	13.32	13.54	17.09	18.47	18.44	0.624	0.546	0.808	0.852	0.839
Image 104	13.10	8.49	14.85	14.85	16.12	0.783	0.655	0.846	0.852	0.873
Image 105	12.82	12.60	13.51	14.72	17.13	0.734	0.747	0.758	0.802	0.921
Image 106	13.29	12.38	14.04	15.40	16.47	0.766	0.791	0.846	0.858	0.890
Image 107	13.20	13.11	14.17	14.97	16.31	0.802	0.818	0.841	0.865	0.905
Image 108	13.30	13.15	15.32	15.36	16.40	0.782	0.811	0.864	0.847	0.888
Image 109	13.17	12.10	13.02	15.01	16.77	0.786	0.803	0.837	0.858	0.896
Image 110	13.33	11.39	17.05	18.41	18.32	0.657	0.524	0.828	0.867	0.852
Arranaga	12 21	11.70	14.22	15.97	16 97	0.759	0.705	0.833	0.944	0.803

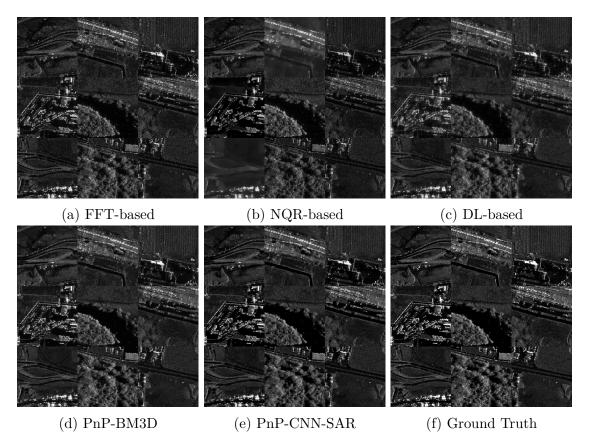


Figure A.26 Reconstruction results for Images 19, 40, 45, 46, 59, 61, 66, 96, and 109 from the Kapıkule test set for the case of 90% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$.

A.3.3 Experiment 3: 80% Data Availability and $\sigma_n = 0.1\sigma_y$

Table A.27 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR for 105 images, and best SSIM for 105 images. Fig. A.27 shows the reconstruction results for a selection of images from the Kapıkule test set from all methods for this experiment. PnP-BM3D and PnP-CNN-SAR produce better reconstructions compared to other methods.

Table A.27 SNR and SSIM values for images from the KapıkuleTest Set for the case of 80% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			SNR (dB)					SSIM		
	FFT-based NC	R-based D	L-based Pi	nP-BM3D Pn	P-CNN-SAR FI	T-based NO	QR-based D	L-based Pr	nP-BM3D PnP	-CNN-SAR
Image 1	10.24	8.48	12.11	13.02	13.96	0.649	0.610	0.705	0.746	0.775
Image 2 Image 3	10.01 10.16	10.67 10.55	11.35 12.42	11.85 12.99	$13.84 \\ 13.66$	$0.659 \\ 0.655$	$0.499 \\ 0.489$	$0.703 \\ 0.720$	$0.687 \\ 0.716$	$0.815 \\ 0.744$
Image 4	10.13	9.35	11.02	11.87	13.33	0.672	0.660	0.721	0.736	0.800
Image 5	10.08	8.48	9.05	11.04	9.39	0.707	0.457	0.642	0.732	0.704
mage 6 mage 7	10.12 10.10	9.91 8.91	12.67 11.30	$14.21 \\ 12.27$	$14.39 \\ 13.67$	$0.596 \\ 0.646$	0.537 0.385	0.698 0.704	$0.756 \\ 0.736$	0.764 0.780
mage 8	10.13	7.91	10.53	12.27	13.37	0.672	0.379	0.703	0.747	0.790
mage 9	9.90	7.17	9.22	11.39	14.26	0.674	0.384	0.663	0.737	0.834
Image 10 Image 11	10.10 10.23	11.44 11.78	11.57 12.38	12.73 13.43	$13.47 \\ 14.18$	$0.691 \\ 0.653$	$0.583 \\ 0.502$	$0.719 \\ 0.709$	$0.746 \\ 0.738$	0.838 0.794
Image 11 Image 12	9.85	8.13	7.60	9.81	13.46	0.733	0.527	0.632	0.736	0.759
Image 13	9.85	9.79	10.77	11.77	13.42	0.688	0.548	0.682	0.720	0.849
mage 14	10.24	9.70	11.74	13.35	13.57	0.626	0.368	0.700	0.751	0.769
mage 15 mage 16	$ \begin{array}{r} 10.25 \\ 9.71 \end{array} $	8.60 8.33	$\frac{12.02}{7.49}$	$\frac{12.84}{9.49}$	$13.42 \\ 13.25$	$0.664 \\ 0.704$	$0.435 \\ 0.523$	$\frac{0.741}{0.566}$	$0.737 \\ 0.663$	$0.784 \\ 0.899$
mage 17	10.08	9.77	12.70	13.97	14.16	0.600	0.540	0.699	0.746	0.732
mage 18	10.17	8.91	11.94	13.20	13.68	0.650	0.419	0.731	0.742	0.775
mage 19 mage 20	9.97 9.97	$9.39 \\ 6.71$	10.58 9.45	11.76 11.36	$13.30 \\ 14.00$	0.717 0.681	$0.701 \\ 0.402$	$0.721 \\ 0.694$	$0.739 \\ 0.744$	$0.824 \\ 0.823$
mage 20	10.01	10.77	11.13	11.88	13.72	0.663	0.602	0.687	0.692	0.820
mage 22	10.13	10.63	11.12	11.76	13.57	0.689	0.563	0.694	0.695	0.847
mage 23	10.03	11.02	11.97	13.01	13.78	0.649	0.619	0.704	0.744	0.780
mage 24 mage 25	10.17 10.15	10.31 10.52	11.43 13.34	12.10 15.16	13.56 14.32	$0.677 \\ 0.554$	$0.668 \\ 0.467$	$\frac{0.744}{0.670}$	0.720 0.744	0. 7 93
mage 26	10.04	10.41	11.11	12.14	13.23	0.703	0.580	0.731	0.739	0.816
mage 27	10.27	10.67	12.45	12.81	13.82	0.659	0.486	0.724	0.715	0.776
mage 28	10.10 10.11	10.76 9.94	12.48 11.22	13.87 12.20	14.17	0.606	0.498	0.697	0.737	0.757
mage 29 mage 30	9.99	10.41	11.22	12.20	$13.04 \\ 13.82$	$0.701 \\ 0.659$	0.534 0.571	$0.710 \\ 0.706$	$0.752 \\ 0.739$	0.828 0.785
mage 31	10.11	9.68	11.88	12.33	12.81	0.692	0.521	0.733	0.730	0.783
mage 32	10.09	10.24	11.90	12.27	13.53	0.671	0.544	0.711	0.717	0.776
mage 33 mage 34	10.07 10.09	11.40 10.66	11.98 10.85	12.39 11.43	$13.98 \\ 13.42$	$0.632 \\ 0.691$	$0.543 \\ 0.615$	$\frac{0.694}{0.685}$	0.693 0.690	$0.796 \\ 0.857$
mage 35	10.03	11.28	11.83	12.89	13.64	0.681	0.562	0.721	0.760	0.805
mage 36	9.93	11.34	11.33	12.56	13.74	0.645	0.561	0.691	0.720	0.815
mage 37	10.11	9.14	12.14	12.77	13.48	0.627	0.574	0.693	0.724	0.739
mage 38 mage 39	9.98 10.12	$7.78 \\ 7.86$	12.40 12.34	13.65 12.74	$14.03 \\ 13.72$	0.619 0.631	0.531 0.554	0.689 0.704	$0.726 \\ 0.725$	$0.735 \\ 0.744$
mage 40	10.21	8.36	10.95	11.88	13.23	0.681	0.350	0.703	0.735	0.799
mage 41	10.15	8.50	11.99	13.28	13.78	0.633	0.422	0.722	0.733	0.773
mage 42 mage 43	10.15 10.20	10.36 10.75	12.07 11.62	13.16 12.23	$13.66 \\ 13.47$	$0.633 \\ 0.670$	0.584 0.608	0.697 0.703	$0.744 \\ 0.703$	$0.754 \\ 0.837$
mage 43	10.16	9.60	11.66	11.71	12.94	0.713	0.708	0.755	0.699	0.800
mage 45	10.13	6.56	9.37	11.24	12.82	0.704	0.320	0.683	0.738	0.818
mage 46	9.66	7.56	7.19	9.98	14.40	0.718	0.513	0.593	0.732	0.877
mage 47 mage 48	10.19 9.91	10.11 9.72	11.52 11.05	$12.22 \\ 11.67$	$13.43 \\ 13.16$	0.681 0.704	$0.679 \\ 0.681$	$0.733 \\ 0.724$	$0.733 \\ 0.720$	$0.792 \\ 0.822$
mage 49	10.32	10.17	13.28	15.08	14.85	0.583	0.452	0.701	0.747	0.751
mage 50	10.11	9.62	10.61	11.52	13.39	0.696	0.541	0.689	0.692	0.851
mage 51 mage 52	9.93 10.05	9.16 11.89	10.43 12.38	10.96 13.64	$13.25 \\ 14.70$	$0.677 \\ 0.605$	0.668	$0.714 \\ 0.697$	$0.726 \\ 0.727$	0.818 0.787
mage 53	10.09	6.37	9.21	11.14	12.98	0.683	0.513 0.352	0.694	0.739	0.787
mage 54	10.10	7.82	11.80	13.19	14.10	0.618	0.383	0.687	0.728	0.790
mage 55	9.93	11.84	12.50	13.56	14.36	0.656	0.543	0.717	0.739	0.800
mage 56 mage 57	10.15 10.11	$10.71 \\ 9.03$	11.39 9.99	12.33 11.93	13.34 12.95	$0.674 \\ 0.689$	$0.540 \\ 0.683$	$0.726 \\ 0.704$	$0.739 \\ 0.742$	$0.796 \\ 0.808$
mage 58	9.95	11.49	11.71	12.58	13.77	0.652	0.574	0.703	0.721	0.814
mage 59	9.91	10.11	10.98	11.77	13.52	0.662	0.571	0.675	0.675	0.834
mage 60	10.14	7.25	11.43	12.35 11.42	$13.60 \\ 13.10$	0.651	0.405	0.720	0.742	0.798
mage 61 mage 62	10.07 10.11	7.31 10.85	10.37 12.04	12.96	13.10	$0.697 \\ 0.649$	0.324 0.598	0.692 0.699	$0.733 \\ 0.734$	$0.806 \\ 0.752$
mage 63	10.14	10.70	12.48	13.19	13.92	0.639	0.531	0.722	0.735	0.775
mage 64	10.08	10.03	11.00	12.30	13.52	0.669	0.660	0.703	0.730	0.794
mage 65 mage 66	10.23 10.14	8.13 6.83	10.65 9.68	12.08 11.70	$13.74 \\ 13.96$	$0.627 \\ 0.705$	$0.369 \\ 0.352$	$0.682 \\ 0.703$	$0.744 \\ 0.760$	0.772 0.830
mage 67	10.13	10.95	10.82	11.87	13.37	0.688	0.645	0.675	0.713	0.863
mage 68	10.11	10.14	11.81	12.19	13.10	0.684	0.671	0.745	0.734	0.787
mage 69	10.12	10.75	12.35	13.63	14.12	0.625	0.541	0.707	0.751	0.776
mage 70 mage 71	10.20 9.98	9.37 11.34	12.39 11.35	13.30 12.42	$13.84 \\ 13.79$	$0.608 \\ 0.655$	$0.576 \\ 0.561$	$0.696 \\ 0.694$	$0.737 \\ 0.717$	$0.750 \\ 0.820$
mage 72	10.15	7.57	10.05	11.50	12.68	0.662	0.355	0.686	0.745	0.784
mage 73	10.00	8.42	9.25	10.29	13.04	0.753	0.722	0.712	0.712	0.852
mage 74	10.10	12.13	12.23	13.85	14.40	0.599	0.493	0.691	0.742	0.775
mage 75 mage 76	10.18 10.05	8.17 9.81	12.60 11.61	13.41 12.25	$13.94 \\ 13.65$	$0.634 \\ 0.668$	$0.536 \\ 0.669$	$0.715 \\ 0.737$	$0.735 \\ 0.723$	0.750 0.799
mage 77	10.15	11.30	12.44	14.00	14.20	0.599	0.500	0.683	0.748	0.749
mage 78	10.04	9.94	11.01	11.70	13.20	0.703	0.697	0.727	0.731	0.814
mage 79	10.08	11.01	11.52	12.59	13.57	0.646	0.517	0.684	0.716	0.795
mage 80 mage 81	10.13 10.15	7.87 10.60	11.17 11.65	12.31 13.50	$13.45 \\ 14.48$	$0.677 \\ 0.576$	$0.315 \\ 0.404$	$0.704 \\ 0.693$	$0.727 \\ 0.736$	$0.796 \\ 0.757$
mage 82	9.89	7.69	7.70	10.03	13.42	0.730	0.692	0.635	0.720	0.855
mage 83	10.08	9.10	11.66	12.38	13.24	0.675	0.550	0.713	0.729	0.763
mage 84	9.97	9.91	11.30	12.43	13.67	0.699	0.542	0.728	0.727	0.830
mage 85 mage 86	9.71 9.88	$\frac{5.84}{7.81}$	8.93 8.47	10.90 10.28	$13.28 \\ 13.27$	$0.698 \\ 0.721$	0.281 0.483	$0.675 \\ 0.649$	$0.722 \\ 0.725$	0.807 0.856
mage 80 mage 87	10.01	6.15	9.20	11.24	13.26	0.721	0.483 0.284	0.649 0.648	$0.725 \\ 0.732$	0.851
		11.58	12.38	13.41	13.91	0.622	0.456	0.700	0.735	0.759

			SNR (dB)		1			SSIM		
Image 89	9.92	10.11	11.40	12.40	13.49	0.695	0.648	0.708	0.748	0.829
Image 90	10.13	9.40	12.25	12.64	13.30	0.641	0.554	0.724	0.724	0.765
Image 91	10.06	7.94	10.47	12.07	13.62	0.672	0.355	0.701	0.741	0.794
Image 92	10.03	8.78	11.92	13.03	13.66	0.655	0.419	0.733	0.745	0.787
Image 93	10.07	6.67	9.55	10.79	12.89	0.722	0.318	0.697	0.730	0.790
Image 94	10.06	8.02	11.62	12.74	13.44	0.660	0.371	0.718	0.746	0.799
Image 95	10.11	10.30	12.33	12.94	13.58	0.661	0.500	0.746	0.741	0.792
Image 96	10.02	9.71	10.47	11.90	13.72	0.688	0.557	0.689	0.749	0.839
Image 97	9.84	10.22	11.15	11.89	13.45	0.677	0.614	0.697	0.686	0.827
Image 98	10.19	10.97	12.65	13.79	13.91	0.621	0.466	0.718	0.739	0.766
Image 99	10.10	10.51	12.08	12.84	13.89	0.668	0.542	0.713	0.734	0.786
Image 100	9.49	6.96	6.64	8.94	13.44	0.700	0.508	0.550	0.667	0.896
Image 101	10.13	9.36	11.92	12.57	13.52	0.661	0.558	0.720	0.724	0.777
Image 102	10.09	10.11	12.06	12.82	13.81	0.654	0.653	0.715	0.734	0.780
Image 103	10.12	12.03	13.41	15.08	14.97	0.500	0.497	0.632	0.703	0.688
Image 104	10.14	8.72	11.98	12.24	13.19	0.678	0.566	0.722	0.733	0.766
Image 105	9.85	11.31	10.72	12.00	13.85	0.672	0.651	0.667	0.704	0.862
Image 106	10.03	7.97	10.74	12.08	13.48	0.662	0.401	0.717	0.748	0.800
Image 107	10.05	9.62	11.31	11.84	13.22	0.696	0.513	0.722	0.746	0.813
Image 108	10.22	8.88	11.77	12.56	13.09	0.690	0.433	0.743	0.716	0.789
Image 109	10.04	7.50	10.27	11.64	13.77	0.688	0.404	0.721	0.744	0.807
Image 110	10.17	9.76	13.27	15.35	14.82	0.532	0.515	0.663	0.737	0.707

13.60

0.662

0.518

0.699

0.729

0.796

10.07

11.22

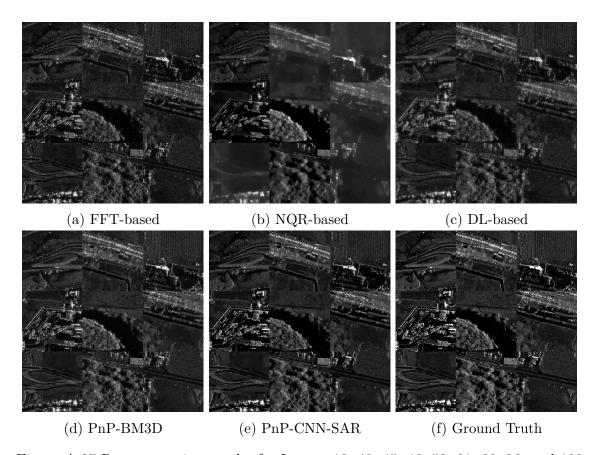


Figure A.27 Reconstruction results for Images 19, 40, 45, 46, 59, 61, 66, 96, and 109 from the Kapıkule test set for the case of 80% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$.

A.3.4 Experiment 4: 70% Data Availability and $\sigma_n = 0.1\sigma_y$

Table A.28 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR for 104 images, and best SSIM for 106 images. Fig. A.28 shows the reconstruction results for a selection of images from the Kapıkule test set from all methods for this experiment. PnP-CNN-SAR produce better reconstructions compared to other methods.

Table A.28 SNR and SSIM values for images from the Kapıkule Test Set for the case of 70% data availability and $\sigma_{\bf n}=0.1\sigma_{\bf y}.$ Best results are shown in bold, second best results are shown in red.

			SNR (dB)					SSIM		
	FFT-based NQ	R-based D	L-based P	nP-BM3D P	nP-CNN-SAR	FFT-based	NQR-based D	L-based P	nP-BM3D PnF	-CNN-SAF
Image 1	8.44	10.29	10.42	11.16	11.34	0.554	0.507	0.589	0.621	0.64
Image 2 Image 3	8.33 8.50	9.90 10.23	9.67 10.63	$10.50 \\ 11.23$	11.24 11.39	0.587 0.550	$0.462 \\ 0.468$	$0.604 \\ 0.616$	$0.574 \\ 0.574$	0.701 0.625
Image 3	8.52	5.90	9.48	10.11	10.91	0.582	0.370	0.621	0.601	0.682
Image 5	8.20	6.89	7.93	8.84	11.08	0.646	0.348	0.557	0.633	0.739
Image 6	8.52	11.59	10.78	11.84	12.03	0.506	0.482	0.579	0.606	0.624
Image 7	8.56	8.13	9.67	10.40	11.17	0.560	0.355	0.603	0.611	0.65
Image 8	8.63	7.32	8.98	10.20	10.75	0.581	0.324	0.606	0.628	0.69
Image 9	8.28	7.50	8.14	9.38	11.06	0.592	0.391	0.572	0.638	0.73
Image 10	8.54	6.92	9.96	10.45	11.17	0.614	0.321	0.635	0.648	0.752
Image 11	8.51	10.06	10.68	11.35	11.58	0.567	0.480	0.611	0.614	0.679
Image 12 Image 13	7.98 8.18	$6.32 \\ 9.41$	$6.50 \\ 9.38$	$8.01 \\ 9.98$	$10.07 \\ 10.87$	$0.666 \\ 0.620$	$0.606 \\ 0.588$	$0.547 \\ 0.614$	$0.638 \\ 0.620$	0.773
Image 14	8.53	8.45	10.13	10.94	11.24	0.536	0.337	0.593	0.606	0.63
Image 15	8.54	9.65	10.44	11.10	11.27	0.568	0.439	0.646	0.607	0.67
mage 16	8.12	6.51	6.70	7.87	10.15	0.661	0.650	0.526	0.594	0.833
mage 17	8.54	7.51	10.84	11.97	11.95	0.498	0.529	0.578	0.602	0.59°
mage 18	8.42	9.51	10.33	11.24	11.48	0.561	0.414	0.633	0.605	0.664
mage 19	8.38	5.92	9.10	10.02	10.93	0.628	0.292	0.624	0.630	0.72
mage 20	8.38	5.21	8.17	9.48	11.07	0.600	0.316	0.603	0.630	0.71
mage 21	8.44 8.35	9.97	9.54 9.57	10.39 10.16	11.18 11.22	0.589	0.545	0.595	0.596	0.720 0.76'
mage 22 mage 23	8.53	8.39 10.87	10.38	10.16	11.22 11.37	$\frac{0.621}{0.562}$	$0.471 \\ 0.488$	$0.615 \\ 0.603$	$0.606 \\ 0.617$	0.66
mage 24	8.46	8.41	9.73	10.48	11.42	0.587	0.463	0.647	0.591	0.68
mage 25	8.57	7.05	11.42	12.37	11.94	0.450	0.360	0.551	0.570	0.54
mage 26	8.40	8.64	9.68	10.24	10.95	0.613	0.612	0.638	0.616	0.71
mage 27	8.39	10.39	10.64	11.37	11.54	0.566	0.436	0.621	0.585	0.65
mage 28	8.56	11.34	10.73	11.85	11.95	0.510	0.438	0.582	0.587	0.62
mage 29	8.40	9.99	9.68	10.16	10.77	0.611	0.527	0.612	0.631	0.72
mage 30	8.52	10.37	10.21	11.00	11.52	0.566	0.514	0.602	0.620	0.67
mage 31	8.50	8.63	10.17	10.66	10.97	0.593	0.312	0.621	0.594	0.68
mage 32	8.60	7.26	10.35 10.44	$10.69 \\ 10.86$	$11.32 \\ 11.42$	0.575	0.551	0.608	0.586	0.64
mage 33 mage 34	8.35 8.34	$10.80 \\ 7.44$	9.48	10.80	10.94	0.552 0.621	$0.516 \\ 0.495$	$\frac{0.608}{0.617}$	$0.564 \\ 0.616$	0.67°
mage 35	8.31	10.48	10.18	10.79	11.64	0.590	0.534	0.616	0.632	0.70
mage 36	8.41	9.31	9.74	10.71	11.35	0.561	0.458	0.599	0.600	0.70
mage 37	8.51	7.50	10.43	10.94	11.44	0.530	0.539	0.575	0.577	0.60
mage 38	8.56	7.78	10.87	11.54	11.45	0.527	0.472	0.582	0.592	0.59
mage 39	8.48	7.98	10.60	11.10	11.43	0.539	0.491	0.587	0.598	0.60
mage 40	8.40	7.52	9.35	10.05	10.68	0.600	0.304	0.605	0.626	0.680
mage 41	8.59	8.89	10.23	11.46	11.62	0.540	0.390	0.613	0.583	0.660
mage 42	8.49	7.20	10.42	11.01	11.29	0.544	0.443	0.585	0.596	0.626
mage 43	8.34 8.58	9.27	$9.86 \\ 9.92$	10.52 10.28	11.13 10.87	0.606	0.565	0.615	0.608	0.750 0.702
mage 44 mage 45	8.32	8.86 6.82	8.14	9.22	10.24	$0.630 \\ 0.628$	$0.626 \\ 0.295$	$\frac{0.660}{0.596}$	$0.593 \\ 0.627$	0.702
mage 46	7.98	5.93	6.25	8.04	10.36	0.652	0.439	0.522	0.639	0.80
mage 47	8.57	9.05	9.83	10.41	11.16	0.588	0.403	0.633	0.600	0.68
mage 48	8.44	8.69	9.62	10.12	10.69	0.618	0.617	0.636	0.618	0.72
mage 49	8.51	9.52	11.14	12.70	12.11	0.490	0.461	0.583	0.598	0.61
mage 50	8.22	5.81	9.43	9.88	10.86	0.633	0.311	0.628	0.618	0.77
mage 51	8.22	8.45	8.93	9.72	10.73	0.584	0.586	0.616	0.607	0.70
mage 52	8.56	10.67	10.63	11.73	12.11	0.508	0.414	0.595	0.582	0.65
mage 53	8.32	5.41	8.09	9.36	10.62	0.601	0.312	0.604	0.625	0.71
mage 54 mage 55	8.38 8.44	9.96 11.05	10.12 10.72	$\frac{11.18}{11.04}$	11.58 11.89	0.526 0.563	0.419 0.505	$0.586 \\ 0.619$	$\frac{0.593}{0.608}$	0.66
mage 56	8.46	8.72	9.80	$\frac{11.04}{10.57}$	11.06	0.580	0.415	0.623	0.623	0.69
mage 57	8.39	6.41	8.66	9.59	10.89	0.602	0.298	0.607	0.616	0.70
mage 58	8.51	10.63	9.98	10.99	11.33	0.572	0.526	0.605	0.612	0.71
mage 59	8.31	9.57	9.55	10.20	11.19	0.591	0.478	0.597	0.575	0.74
mage 60	8.46	5.64	9.88	10.61	11.12	0.570	0.368	0.622	0.614	0.68
mage 61	8.49	6.45	8.91	9.63	10.57	0.617	0.264	0.592	0.622	0.70
mage 62	8.45	10.00	10.37	10.88	11.34	0.549	0.523	0.583	0.600	0.63
mage 63 mage 64	8.35 8.48	$8.93 \\ 5.23$	10.59 9.38	11.43 10.55	$11.72 \\ 11.23$	0.543	$0.512 \\ 0.352$	$0.610 \\ 0.612$	$0.596 \\ 0.603$	$0.64 \\ 0.68$
mage 65	8.44	6.44	9.38	10.33 10.17	11.23	$0.580 \\ 0.542$	0.335	0.512 0.580	0.606	0.64
mage 66	8.25	5.99	8.42	9.35	10.71	0.617	0.330	0.611	0.657	0.73
mage 67	8.39	9.63	9.32	10.11	10.95	0.609	0.551	0.599	0.621	0.78
mage 68	8.48	5.29	9.99	10.55	11.01	0.589	0.375	0.635	0.607	0.68
mage 69	8.57	10.80	10.59	11.41	11.81	0.533	0.501	0.608	0.616	0.64
mage 70	8.41	10.69	10.55	11.43	11.77	0.517	0.453	0.577	0.590	0.61
mage 71	8.38	9.24	9.67	10.74	11.45	0.577	0.430	0.599	0.614	0.71
mage 72	8.41	6.13	8.66	9.52	10.57	0.568	0.287	0.584	0.623	0.67
mage 73	8.27	4.53	8.07	8.71	10.31	0.675	0.275	0.630	0.615	0.76
mage 74	8.58	11.62	10.45	11.62	11.76	0.510	0.469	0.580	0.586	0.64
mage 75 mage 76	8.52 8.56	7.40 6.33	10.69 9.90	11.58	11.83 11.32	0.537	0.402	0.600	0.611	$0.61 \\ 0.69$
mage 76 mage 77	8.56 8.54	6.33 10.52	9.90 10.66	$10.60 \\ 11.71$	11.32 11.80	0.581 0.502	$0.396 \\ 0.512$	$\frac{0.635}{0.570}$	$0.598 \\ 0.592$	0.69
mage 77	8.38	8.33	9.32	10.10	10.95	0.622	0.390	0.624	0.615	0.80
mage 79	8.41	8.88	9.32	10.10 10.64	11.23	0.562	0.493	0.524 0.589	0.572	0.71
mage 80	8.37	6.45	9.68	10.32	10.96	0.581	0.235	0.604	0.598	0.68
mage 81	8.46	8.62	10.08	11.20	12.52	0.483	0.376	0.585	0.597	0.62
mage 82	7.98	6.07	6.73	8.38	10.37	0.654	0.368	0.561	0.631	0.78
mage 83	8.44	7.54	9.96	10.55	11.05	0.580	0.519	0.604	0.597	0.64
mage 84	8.51	9.41	9.65	10.67	10.80	0.626	0.524	0.646	0.628	0.73
mage 85	8.26	5.31	7.81	9.14	10.50	0.610	0.267	0.585	0.608	0.71
	8.20	5.79	7.36	8.69	10.46	0.646	0.281	0.567	0.634	0.77
mage 86 mage 87	8.25	7.16	7.89	9.19	10.60	0.621	0.359	0.564	0.627	0.76

			Ta	ble A.28 con	tinued from	previous p	age			
			SNR (dB)					SSIM		
Image 89	8.50	9.31	9.94	10.47	11.07	0.611	0.586	0.613	0.638	0.730
Image 90	8.44	6.95	10.51	11.08	11.52	0.556	0.524	0.625	0.603	0.659
Image 91	8.52	6.62	8.98	10.06	10.99	0.573	0.321	0.599	0.622	0.685
Image 92	8.46	7.64	10.23	11.16	11.40	0.564	0.397	0.633	0.607	0.669
Image 93	8.17	6.68	8.18	9.02	10.78	0.640	0.306	0.602	0.621	0.724
Image 94	8.47	6.97	9.88	10.84	11.16	0.572	0.320	0.616	0.620	0.690
Image 95	8.58	6.72	10.34	11.12	11.41	0.577	0.402	0.639	0.609	0.683
Image 96	8.36	5.95	9.12	9.79	11.18	0.601	0.260	0.592	0.639	0.735
Image 97	8.27	9.30	9.68	10.29	11.07	0.599	0.577	0.619	0.596	0.740
Image 98	8.51	10.71	10.82	11.78	11.60	0.528	0.476	0.616	0.593	0.644
Image 99	8.42	10.63	10.52	11.16	11.59	0.566	0.485	0.611	0.621	0.667
Image 100	7.82	6.28	5.80	7.54	10.11	0.651	0.621	0.491	0.588	0.822
Image 101	8.54	9.69	10.28	10.77	11.38	0.577	0.504	0.620	0.606	0.666
Image 102	8.50	8.82	10.21	10.84	11.35	0.560	0.472	0.603	0.606	0.656
Image 103	8.49	9.87	11.48	12.17	12.38	0.392	0.445	0.510	0.532	0.509
Image 104	8.48	7.83	10.16	10.60	11.17	0.584	0.523	0.602	0.606	0.650
Image 105	8.21	10.32	9.31	10.09	10.91	0.602	0.604	0.592	0.611	0.775
Image 106	8.31	6.48	9.38	10.24	10.97	0.578	0.371	0.622	0.623	0.686
Image 107	8.44	9.70	9.74	10.13	10.92	0.607	0.495	0.615	0.639	0.705
Image 108	8.54	3.78	10.23	10.84	11.18	0.603	0.388	0.647	0.599	0.690
Image 109	8.42	6.11	8.76	9.92	10.94	0.605	0.374	0.624	0.633	0.711
Image 110	8.55	10.65	11.47	12.85	12.56	0.430	0.471	0.542	0.576	0.562
Average	8.41	8.22	9.65	10.49	11.20	0.576	0.439	0.601	0.608	0.688

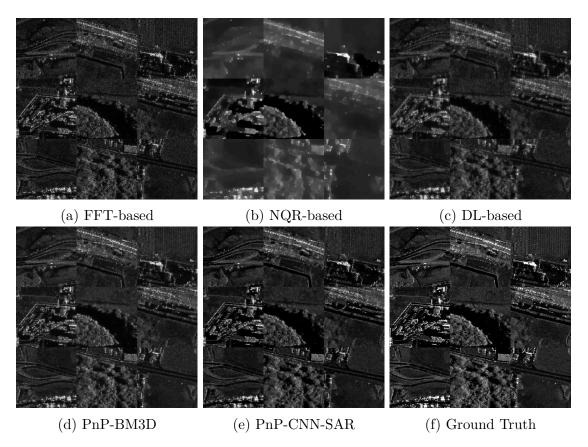


Figure A.28 Reconstruction results for Images 19, 40, 45, 46, 59, 61, 66, 96, and 109 from the Kapıkule test set for the case of 70% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$.

A.3.5 Experiment 5: 50% Data Availability and $\sigma_n = 0.1\sigma_y$

Table A.29 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR for 37 images, and best SSIM for 101 images. Fig. A.29 shows the reconstruction results for a selection of images from the Kapıkule test set from all methods for this experiment. PnP-CNN-SAR produce better reconstructions compared to other methods.

Table A.29 SNR and SSIM values for images from the Kapıkule Test Set for the case of 50% data availability and $\sigma_{\bf n}=0.1\sigma_{\bf y}.$ Best results are shown in bold, second best results are shown in red.

			SNR (dB)					SSIM		
	FFT-based NQ	R-based D	L-based Pr	ıP-BM3D I	PnP-CNN-SAR	FFT-based	NQR-based l	DL-based F	PnP-BM3D F	PnP-CNN-SAR
Image 1	6.45	8.67	8.25	9.51	8.72	0.379	0.463	0.422	0.411	0.480
Image 2 Image 3	6.29 6.46	9.11 7.29	$7.65 \\ 8.24$	$\frac{8.93}{10.13}$	8.90 8.81	0.346 0.311	$0.440 \\ 0.452$	$0.468 \\ 0.442$	$0.402 \\ 0.420$	$0.529 \\ 0.461$
Image 4	6.42	5.98	7.52	8.47	8.41	0.339	0.355	0.475	0.421	0.523
Image 5	6.05	6.29	6.42	6.92	7.88	0.459	0.333	0.432	0.452	0.621
Image 6	6.45	10.90	8.51	10.40	8.96	0.300	0.420	0.422	0.405	0.445
Image 7 Image 8	$6.36 \\ 6.28$	$7.99 \\ 6.79$	7.67 7.20	$8.66 \\ 8.08$	8.39 7.92	$0.341 \\ 0.361$	$0.329 \\ 0.296$	$0.457 \\ 0.464$	0.417 0.439	0.503 0.533
Image 8 Image 9	6.06	6.18	6.52	7.46	8.28	0.301	0.296	0.404 0.427	0.439 0.444	0.550
Image 10	6.40	8.07	7.89	8.70	8.83	0.399	0.382	0.502	0.464	0.594
Image 11	6.47	8.64	8.29	9.82	9.22	0.355	0.434	0.458	0.402	0.508
Image 12	5.88	5.08	5.27	5.91	7.52	0.466	0.497	0.412	0.442	0.635
Image 13	6.25	8.79	7.32	8.15	8.58	0.399	0.517	0.483	0.437	0.615
Image 14 Image 15	$6.42 \\ 6.34$	8.58 9.23	$7.96 \\ 8.13$	$9.18 \\ 9.68$	8.62 8.77	0.318 0.307	0.319 0.427	$0.436 \\ 0.504$	$0.404 \\ 0.446$	0.471 0.524
Image 16	5.82	4.95	5.34	5.63	8.04	0.400	0.363	0.450	0.434	0.723
Image 17	6.51	6.91	8.37	10.59	8.98	0.293	0.447	0.405	0.424	0.418
Image 18	6.49	5.70	8.04	9.76	8.63	0.302	0.372	0.484	0.426	0.498
Image 19	6.32	6.15	7.28	8.15	8.24	0.412	0.280	0.473	0.435	0.540
Image 20 Image 21	6.16 6.33	$\frac{5.72}{7.98}$	$6.60 \\ 7.63$	$7.36 \\ 8.60$	8.03 8.92	0.369 0.370	0.333 0.449	$0.476 \\ 0.466$	0.449 0.419	0.576 0.567
Image 21 Image 22	6.31	8.13	7.56	8.37	8.81	0.380	0.435	0.486	0.411	0.601
Image 22 Image 23	6.43	8.36	8.13	9.25	8.84	0.376	0.314	0.434	0.388	0.502
Image 24	6.36	7.35	7.76	9.17	8.95	0.312	0.448	0.504	0.457	0.536
Image 25	6.52	8.94	8.77	11.55	9.16	0.250	0.387	0.371	0.399	0.365
Image 26	6.33	8.61	7.53	8.60	8.49	0.387	0.459	0.485	0.441	0.560
Image 27 Image 28	$6.51 \\ 6.54$	$\frac{6.08}{9.91}$	$8.40 \\ 8.37$	$10.21 \\ 10.26$	9.05 9.18	0.327 0.291	$0.401 \\ 0.416$	$0.472 \\ 0.431$	$0.410 \\ 0.394$	$0.486 \\ 0.452$
Image 28 Image 29	6.37	7.85	7.66	8.52	8.45	0.291 0.427	0.416	0.431 0.461	0.394 0.417	0.452
Image 30	6.44	9.14	8.07	9.32	8.73	0.354	0.462	0.445	0.410	0.498
Image 31	6.45	8.99	8.00	9.04	8.58	0.370	0.374	0.455	0.363	0.501
Image 32	6.38	5.45	8.11	9.24	8.72	0.375	0.457	0.430	0.392	0.471
Image 33 Image 34	$6.40 \\ 6.44$	$9.43 \\ 8.73$	$8.06 \\ 7.42$	9.64	9.22 8.69	0.322 0.368	0.496	$0.456 \\ 0.488$	0.414 0.429	0.512 0.611
Image 34 Image 35	6.51	9.60	8.01	8.13 9.14	8.73	0.389	$0.530 \\ 0.461$	0.454	0.429	0.521
mage 36	6.32	7.58	7.73	9.06	9.12	0.356	0.429	0.465	0.416	0.556
Image 37	6.43	6.72	8.21	9.65	8.67	0.319	0.437	0.407	0.392	0.425
mage 38	6.55	8.99	8.41	10.19	9.04	0.330	0.433	0.404	0.402	0.430
Image 39	6.48	4.61	8.27	9.89	8.80	0.329	0.447	0.411	0.440	0.437
Image 40 Image 41	$6.25 \\ 6.47$	7.11 9.68	$7.42 \\ 8.08$	$8.10 \\ 10.01$	8.17 9.08	0.387 0.290	0.284 0.414	$0.446 \\ 0.474$	$0.426 \\ 0.422$	$0.534 \\ 0.503$
Image 42	6.48	5.65	8.09	9.48	8.75	0.334	0.433	0.408	0.392	0.450
Image 43	6.35	8.47	7.85	8.86	8.93	0.367	0.504	0.496	0.440	0.589
Image 44	6.50	8.10	7.86	9.07	8.56	0.366	0.457	0.528	0.462	0.559
Image 45	5.99	5.59	6.60	7.09	8.01	0.411	0.262	0.456	0.417	0.573
Image 46	5.81 6.35	$\frac{4.37}{8.66}$	$\frac{5.08}{7.85}$	5.76	7.67 8.63	0.459	0.466	$0.409 \\ 0.488$	0.448	0.670 0.525
Image 47 Image 48	6.34	8.23	7.44	8.91 8.31	8.43	0.355 0.392	0.388 0.488	0.488 0.481	$0.410 \\ 0.427$	0.564
Image 49	6.53	11.43	8.82	11.25	9.25	0.258	0.441	0.443	0.420	0.440
Image 50	6.17	5.68	7.32	8.09	8.72	0.392	0.311	0.507	0.450	0.636
Image 51	6.30	7.28	7.23	7.83	8.00	0.369	0.371	0.473	0.420	0.527
Image 52	6.41	10.60	8.31	10.31	9.43	0.286 0.390	0.453	0.452	0.402	0.494
Image 53 Image 54	6.03 6.38	$\frac{4.67}{9.28}$	$6.57 \\ 8.15$	$7.17 \\ 9.54$	8.16 9.15	0.390	0.319 0.362	$0.467 \\ 0.458$	0.434 0.394	0.576 0.506
Image 55	6.34	9.95	8.37	9.70	9.13	0.374	0.440	0.464	0.416	0.519
Image 56	6.47	7.99	7.66	8.79	8.61	0.366	0.397	0.473	0.419	0.527
Image 57	6.31	5.46	7.04	7.93	8.54	0.393	0.306	0.476	0.429	0.547
mage 58	6.35	9.16	7.90	9.29	9.11	0.356	0.492	0.471	0.428	0.560
Image 59 Image 60	$6.30 \\ 6.38$	$8.63 \\ 8.79$	$7.58 \\ 7.70$	8.73 8.85	8.85 8.55	0.343 0.337	$0.486 \\ 0.398$	$0.479 \\ 0.475$	$0.410 \\ 0.420$	$0.584 \\ 0.522$
mage 61	6.21	6.25	7.18	7.83	8.08	0.402	0.358	0.449	0.423	0.553
mage 62	6.42	9.93	8.20	9.25	8.73	0.365	0.387	0.415	0.362	0.466
mage 63	6.50	7.92	8.32	10.09	9.01	0.313	0.482	0.455	0.429	0.482
mage 64 mage 65	6.37	6.12	7.59	8.73	8.61	0.355	0.327	0.467	0.418	0.522
mage 65 mage 66	$6.25 \\ 6.07$	$7.05 \\ 5.51$	$7.30 \\ 6.59$	8.44 7.31	8.43 8.05	$0.326 \\ 0.399$	$0.306 \\ 0.285$	$0.436 \\ 0.469$	0.418 0.466	0.499
mage 67	6.25	7.87	7.36	8.22	8.66	0.397	0.413	0.403 0.472	0.430	0.628
mage 68	6.43	8.97	7.84	8.95	8.51	0.337	0.453	0.489	0.404	0.527
mage 69	6.47	10.25	8.34	9.83	9.02	0.330	0.444	0.449	0.408	0.479
mage 70	6.43	9.25	8.21	9.91	8.72	0.320	0.387	0.402	0.382	0.432
mage 71 mage 72	6.38 6.26	$\frac{9.01}{6.39}$	7.65	8.98 7.46	9.04 7.82	0.348 0.354	$0.455 \\ 0.301$	$0.469 \\ 0.429$	$0.428 \\ 0.413$	0.567 0.511
mage 72 mage 73	6.08	4.84	6.98 6.48	6.79	7.76	0.354 0.454	0.269	0.429 0.495	0.413 0.434	0.611
mage 74	6.49	9.83	8.27	10.13	9.07	0.305	0.438	0.431	0.400	0.470
mage 75	6.55	6.72	8.40	10.27	8.78	0.300	0.383	0.444	0.444	0.45
mage 76	6.37	8.66	7.68	9.05	8.62	0.322	0.412	0.489	0.429	0.528
mage 77	6.49	10.69	8.41	10.29	8.82	0.305	0.418	0.404	0.392	0.420
mage 78 mage 79	6.31 6.33	7.78 9.29	$7.34 \\ 7.79$	$8.17 \\ 9.07$	8.28 9.01	0.386 0.341	$0.356 \\ 0.436$	$0.469 \\ 0.444$	$0.403 \\ 0.373$	$0.541 \\ 0.512$
mage 79	6.36	7.37	7.79	8.71	8.55	0.341	0.436	0.444 0.444	0.373	0.493
mage 81	6.43	8.81	7.95	9.44	8.58	0.259	0.371	0.449	0.408	0.46
mage 82	5.87	4.01	5.59	6.24	7.87	0.454	0.247	0.444	0.458	0.657
mage 83	6.42	8.01	7.96	8.95	8.46	0.361	0.408	0.449	0.411	0.491
mage 84	6.27	8.08	7.62	8.65	8.75	0.376	0.456	0.526	0.475	0.598
mage 85 mage 86	$6.07 \\ 5.97$	5.27	6.32 6.01	$7.01 \\ 6.43$	7.82 7.80	0.416	$0.261 \\ 0.207$	$\frac{0.441}{0.426}$	$0.402 \\ 0.424$	0.556 0.618
Image 86 Image 87	5.97 6.11	$\frac{4.37}{6.58}$	6.01 6.50	$6.43 \\ 6.85$	7.80 8.18	$0.465 \\ 0.438$	$0.207 \\ 0.327$	$0.426 \\ 0.435$	0.424 0.408	0.600
	0.11	9.88	8.38	9.90	9.01		0.395	0.428	0.376	0.456

			Ta	ble A.29 cont	tinued from	previous p	age			
			SNR (dB)					SSIM		
Image 89	6.36	9.08	7.83	8.71	8.63	0.420	0.481	0.463	0.424	0.550
Image 90	6.39	9.36	8.25	9.71	8.69	0.319	0.473	0.470	0.446	0.494
Image 91	6.30	6.87	7.12	7.91	8.05	0.365	0.311	0.460	0.424	0.531
Image 92	6.37	8.55	8.07	9.69	8.78	0.306	0.388	0.492	0.433	0.514
Image 93	6.12	5.46	6.66	7.33	7.67	0.418	0.268	0.454	0.436	0.565
Image 94	6.41	8.33	7.87	8.93	8.67	0.361	0.376	0.469	0.410	0.516
Image 95	6.42	7.67	8.25	9.80	8.86	0.318	0.401	0.504	0.437	0.528
Image 96	6.27	7.48	7.27	8.01	8.50	0.427	0.366	0.446	0.429	0.565
Image 97	6.26	8.53	7.74	8.54	8.88	0.356	0.495	0.504	0.433	0.593
Image 98	6.49	7.22	8.50	10.56	9.11	0.292	0.407	0.466	0.403	0.482
Image 99	6.37	9.89	8.19	9.51	8.81	0.360	0.438	0.446	0.398	0.497
Image 100	5.71	4.72	4.79	5.13	7.91	0.413	0.375	0.407	0.402	0.702
Image 101	6.34	8.86	8.10	9.27	8.67	0.364	0.406	0.462	0.417	0.503
Image 102	6.34	9.19	8.08	9.44	8.68	0.349	0.419	0.449	0.424	0.491
Image 103	6.51	10.74	8.90	10.71	9.51	0.230	0.336	0.326	0.319	0.323
Image 104	6.37	6.33	8.03	9.03	8.44	0.392	0.379	0.430	0.407	0.479
Image 105	6.29	8.87	7.30	8.16	8.72	0.377	0.555	0.479	0.435	0.638
Image 106	6.22	6.04	7.31	8.42	8.41	0.340	0.356	0.475	0.437	0.536
Image 107	6.30	8.79	7.58	8.49	8.29	0.438	0.422	0.438	0.423	0.526
Image 108	6.38	8.49	8.06	9.37	8.89	0.345	0.377	0.505	0.432	0.541
Image 109	6.26	6.03	7.05	7.92	8.33	0.361	0.354	0.489	0.459	0.567
Image 110	6.54	10.21	8.87	11.64	9.53	0.247	0.394	0.371	0.368	0.368
Average	6.33	7.78	7.63	8.81	8.62	0.358	0.393	0.455	0.419	0.527

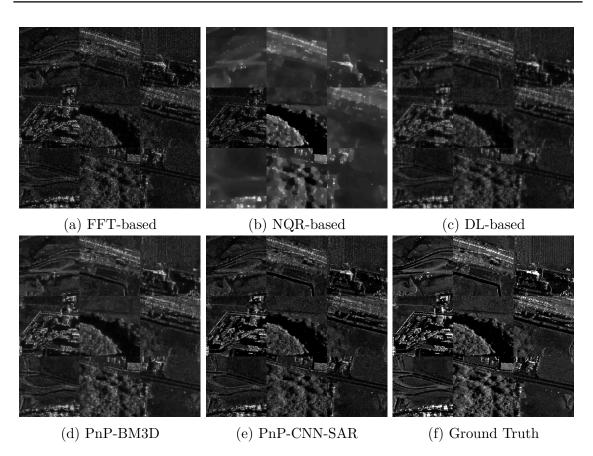


Figure A.29 Reconstruction results for Images 19, 40, 45, 46, 59, 61, 66, 96, and 109 from the Kapıkule test set for the case of 50% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$.

A.3.6 Experiment 6: 30% Data Availability and $\sigma_n = 0.1\sigma_y$

Table A.30 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR for 7 images, and best SSIM for 63 images. Fig. A.30 shows the reconstruction results for a selection of images from the Kapıkule test set from all methods for this experiment. PnP-CNN-SAR produce better reconstructions compared to other methods.

Table A.30 SNR and SSIM values for images from the Kapıkule Test Set for the case of 30% data availability and $\sigma_{\bf n}=0.1\sigma_{\bf y}.$ Best results are shown in bold, second best results are shown in red.

			SNR (dB)					SSIM		
	FFT-based NQF	R-based I	DL-based Pn	P-BM3D I	PnP-CNN-SAR	FFT-based	NQR-based	DL-based l	PnP-BM3D	PnP-CNN-SAR
Image 1	4.77	9.07	6.04	7.15	5.83	0.232	0.392	0.281	0.303	0.313
Image 2 Image 3	$\frac{4.65}{4.76}$	$8.15 \\ 6.62$	5.79 6.15	$\frac{6.79}{7.74}$	6.51 6.05	0.211 0.176	0.364 0.370	$\frac{0.367}{0.332}$	0.323 0.377	0.410 0.338
Image 4	4.77	5.67	5.63	6.67	5.78	0.204	0.330	0.370	0.347	0.394
Image 5	4.33	4.66	4.84	5.37	5.19	0.295	0.197	0.315	0.319	0.422
Image 6	4.75	9.71	6.07	7.59	6.13	0.196	0.374	0.291	0.308	0.317
Image 7 Image 8	4.66 4.68	$\frac{7.51}{5.94}$	$5.70 \\ 5.42$	$6.75 \\ 6.39$	5.62 5.53	$0.193 \\ 0.208$	0.316 0.285	$0.335 \\ 0.345$	0.333 0.342	$0.359 \\ 0.389$
Image 9	4.41	4.34	4.91	5.62	5.61	0.268	0.198	0.300	0.294	0.408
Image 10	4.71	7.57	5.75	6.84	6.64	0.253	0.348	0.381	0.354	0.439
Image 11	4.78	9.28	6.25	7.44	6.30	0.230	0.367	0.350	0.310	0.364
Image 12	4.23	4.70	4.06	4.68	4.83	0.292	0.425	0.298	0.314	0.431
Image 13 Image 14	4.63 4.82	$\frac{7.55}{6.28}$	5.47 5.95	6.26 7.11	6.73 5.97	0.265 0.186	0.397 0.298	0.378 0.324	0.342 0.301	$0.456 \\ 0.341$
Image 14	4.75	7.88	6.01	7.45	6.04	0.130	0.390	0.400	0.390	0.407
Image 16	4.08	2.77	4.07	4.47	5.42	0.281	0.179	0.377	0.360	0.562
Image 17	4.82	8.29	6.24	7.84	6.10	0.165	0.423	0.292	0.375	0.295
Image 18	4.74	8.74	6.08	7.45	5.87	0.176	0.419	0.377	0.365	0.387
Image 19 Image 20	4.58 4.44	6.82 5.05	5.39 5.05	$\frac{6.42}{5.71}$	5.82 5.54	0.254 0.234	0.338 0.298	$0.344 \\ 0.367$	0.323 0.355	$0.394 \\ 0.426$
Image 20	4.65	8.11	5.58	6.60	6.42	0.234	0.465	0.343	0.339	0.434
Image 22	4.68	4.53	5.51	6.48	6.61	0.255	0.217	0.381	0.334	0.461
Image 23	4.77	8.95	5.90	7.10	6.07	0.238	0.345	0.287	0.251	0.337
Image 24	4.77	7.62	5.93	7.16	5.95	0.188	0.420	0.412	0.404	0.420
Image 25	4.86	7.10	6.49 5.66	8.38	6.12	0.136	0.295	0.263	0.357	0.259
Image 26 Image 27	4.63 4.79	7.30 5.18	$\frac{5.66}{6.18}$	$\frac{6.61}{7.66}$	5.95 6.16	0.247 0.180	0.351 0.258	$0.366 \\ 0.357$	0.335 0.357	$0.403 \\ 0.368$
Image 27	4.79	9.95	6.32	7.75	6.22	0.173	0.238	0.327	0.318	0.328
Image 29	4.79	8.05	5.76	6.51	6.01	0.278	0.393	0.324	0.269	0.360
Image 30	4.74	8.25	5.98	7.07	6.05	0.215	0.341	0.320	0.308	0.345
Image 31	4.71	7.20	5.94	7.11	5.95	0.228	0.334	0.325	0.269	0.355
Image 32 Image 33	$4.74 \\ 4.65$	$\frac{6.81}{7.39}$	$5.94 \\ 5.98$	$7.32 \\ 7.38$	6.03 6.83	$0.226 \\ 0.193$	$0.373 \\ 0.428$	0.293 0.356	$0.327 \\ 0.362$	$0.328 \\ 0.381$
Image 34	4.76	7.54	5.47	6.31	6.59	0.193	0.428	0.386	0.339	0.483
Image 35	4.69	8.23	5.93	6.95	6.03	0.258	0.391	0.314	0.285	0.349
Image 36	4.63	7.85	5.73	6.98	6.44	0.223	0.418	0.351	0.335	0.395
Image 37	4.86	6.56	6.08	7.45	6.01	0.185	0.382	0.273	0.327	0.298
Image 38	4.82	7.27	6.20	7.76	6.01	0.178	0.319	0.281	0.330	0.300
Image 39 Image 40	$4.76 \\ 4.71$	8.86 5.98	6.13 5.48	7.52 6.29	6.00 5.57	0.192 0.246	0.436 0.256	0.289 0.319	$\frac{0.375}{0.302}$	0.304 0.379
Image 40	4.76	8.48	6.06	7.44	6.10	0.169	0.363	0.319	0.367	0.381
Image 42	4.78	7.33	5.99	7.28	6.01	0.216	0.354	0.286	0.292	0.322
Image 43	4.78	6.91	5.85	6.84	6.75	0.237	0.454	0.393	0.359	0.453
Image 44	4.75	7.35	5.85	6.99	5.85	0.205	0.460	0.411	0.392	0.433
Image 45 Image 46	4.38 4.11	$\frac{5.12}{2.86}$	$\frac{5.00}{3.94}$	5.56 4.42	$\frac{5.24}{4.78}$	$0.263 \\ 0.302$	0.223 0.180	$\frac{0.338}{0.297}$	0.303 0.314	$0.410 \\ 0.440$
Image 47	4.63	7.16	5.69	7.05	5.93	0.207	0.381	0.362	0.348	0.385
Image 48	4.67	6.98	5.55	6.43	6.07	0.235	0.332	0.360	0.318	0.424
Image 49	4.79	3.76	6.51	8.30	6.18	0.142	0.183	0.345	0.371	0.331
Image 50	4.72	7.26	5.41	6.37	6.51	0.254	0.408	0.411	0.381	0.493
Image 51 Image 52	$4.55 \\ 4.72$	$6.53 \\ 9.89$	5.37 6.13	$6.28 \\ 7.61$	5.65 6.47	0.232 0.172	$0.440 \\ 0.393$	$0.360 \\ 0.357$	0.331 0.345	$0.389 \\ 0.395$
Image 52	4.43	4.57	4.97	5.60	5.45	0.172	0.290	0.356	0.343	0.333
Image 54	4.72	8.95	5.92	7.24	6.44	0.184	0.370	0.349	0.334	0.382
Image 55	4.80	9.44	6.10	7.39	6.27	0.234	0.378	0.341	0.306	0.375
Image 56	4.74	7.67	5.79	6.82	5.90	0.223	0.358	0.357	0.334	0.390
Image 57 Image 58	$4.66 \\ 4.76$	$\frac{5.75}{8.50}$	5.21 5.78	$6.23 \\ 6.94$	5.60 6.51	0.233 0.221	0.274 0.423	$\frac{0.350}{0.359}$	0.325 0.323	0.397 0.410
Image 59	4.60	7.72	5.69	6.70	6.89	0.225	0.367	0.380	0.328	0.469
Image 60	4.74	7.83	5.73	6.96	5.80	0.205	0.344	0.367	0.346	0.386
Image 61	4.56	4.91	5.36	6.20	5.44	0.246	0.226	0.319	0.306	0.394
Image 62	4.81	7.13	6.07	7.13	6.01	0.215	0.326	0.277	0.252	0.302
Image 63 Image 64	$4.75 \\ 4.62$	$7.46 \\ 7.40$	$6.04 \\ 5.68$	$\begin{array}{c} \textbf{7.51} \\ \textbf{6.62} \end{array}$	6.11 6.08	$0.174 \\ 0.216$	$0.439 \\ 0.376$	0.343 0.353	$\frac{0.379}{0.325}$	0.351 0.389
Image 65	4.64	5.90	5.50	6.44	5.74	0.210	0.293	0.333	0.325	0.368
Image 66	4.50	5.19	4.94	5.86	5.28	0.250	0.268	0.348	0.352	0.412
Image 67	4.68	7.44	5.50	6.30	6.40	0.267	0.423	0.359	0.308	0.458
Image 68	4.69	8.36	5.89	7.11	5.77	0.187	0.420	0.377	0.351	0.398
Image 69 Image 70	4.84	8.25	6.11	7.44	6.10	0.208	0.370	0.327	0.315	0.342
Image 70 Image 71	$4.79 \\ 4.67$	$\frac{6.46}{7.39}$	$6.12 \\ 5.70$	7.56 6.79	$6.07 \\ 6.51$	$0.181 \\ 0.227$	0.354 0.396	$0.286 \\ 0.368$	0.309 0.340	0.303 0.437
Image 72	4.59	5.98	5.28	5.97	5.39	0.214	0.258	0.313	0.309	0.358
Image 73	4.42	3.11	4.85	5.43	5.37	0.281	0.243	0.371	0.336	0.454
Image 74	4.77	9.30	6.10	7.44	6.23	0.186	0.374	0.323	0.313	0.342
Image 75	4.82	7.82	6.28	7.75	6.05	0.175	0.418	0.327	0.390	0.336
Image 76 Image 77	4.68 4.80	$8.39 \\ 9.23$	5.84 6.15	$7.06 \\ 7.63$	5.93 6.15	0.195 0.189	$0.439 \\ 0.368$	0.385 0.281	$0.366 \\ 0.312$	$0.406 \\ 0.292$
Image 77 Image 78	4.59	$\frac{9.23}{7.51}$	5.62	6.58	5.81	0.189	0.342	0.281 0.352	0.312 0.317	0.397
Image 79	4.74	7.79	5.86	6.96	6.51	0.208	0.320	0.335	0.297	0.386
Image 80	4.73	7.95	5.77	6.80	6.00	0.227	0.290	0.310	0.265	0.343
Image 81	4.86	7.58	5.98	7.07	5.86	0.146	0.365	0.350	0.356	0.354
Image 82	4.27	3.08	4.21	4.88	5.10	0.288	0.218	0.330	0.328	0.473
Image 83	4.77	7.72	5.91	6.96	5.92	0.209	0.433	0.315	0.344	0.345
Image 84 Image 85	4.73 4.35	7.39 4.57	5.61 4.69	6.68 5.59	6.77 5.39	0.229 0.256	0.412 0.228	0.417 0.316	0.404 0.306	$0.468 \\ 0.407$
Image 86	4.33	3.85	4.58	$\frac{5.59}{5.10}$	5.17	0.230	0.228	0.316	0.300	0.428
Image 87	4.43	4.77	4.85	5.56	5.46	0.281	0.174	0.308	0.293	0.419
	4.88	9.21	6.15	7.52	6.10		0.362	0.309	0.286	0.315

			Tab	ole A.30 cont	inued from p	previous p	age			
			SNR (dB)					SSIM		
Image 89	4.77	8.49	5.85	6.73	6.12	0.274	0.416	0.329	0.286	0.375
Image 90	4.75	7.85	6.01	7.45	6.10	0.182	0.460	0.357	0.395	0.370
Image 91	4.72	6.38	5.40	6.13	5.56	0.220	0.280	0.342	0.323	0.398
Image 92	4.79	5.81	5.99	7.30	6.09	0.178	0.368	0.391	0.380	0.397
Image 93	4.48	5.26	4.91	5.72	5.38	0.265	0.226	0.328	0.307	0.398
Image 94	4.76	8.58	5.80	6.98	5.92	0.221	0.414	0.351	0.313	0.368
Image 95	4.77	5.60	5.99	7.44	6.01	0.190	0.417	0.390	0.375	0.403
Image 96	4.64	7.31	5.48	6.27	5.97	0.282	0.308	0.310	0.278	0.376
Image 97	4.62	7.63	5.59	6.54	6.47	0.225	0.428	0.391	0.360	0.476
Image 98	4.85	6.06	6.18	7.80	6.20	0.175	0.336	0.355	0.338	0.359
Image 99	4.72	9.11	5.98	7.24	6.09	0.231	0.379	0.310	0.293	0.341
Image 100	4.04	4.09	3.79	4.19	5.20	0.276	0.444	0.332	0.321	0.533
Image 101	4.82	7.55	5.88	7.17	6.03	0.218	0.382	0.335	0.342	0.357
Image 102	4.71	8.49	5.98	7.21	6.12	0.216	0.365	0.329	0.336	0.348
Image 103	4.90	8.14	6.57	7.88	6.29	0.128	0.274	0.227	0.241	0.224
Image 104	4.85	7.89	5.84	7.12	5.97	0.242	0.385	0.286	0.317	0.326
Image 105	4.59	7.33	5.56	6.28	6.79	0.261	0.409	0.385	0.346	0.473
Image 106	4.62	5.14	5.53	6.50	5.69	0.209	0.329	0.370	0.348	0.401
Image 107	4.73	7.40	5.58	6.49	5.84	0.281	0.249	0.291	0.250	0.357
Image 108	4.68	7.07	5.93	7.32	5.98	0.197	0.353	0.391	0.373	0.401
Image 109	4.55	5.88	5.26	6.30	5.52	0.219	0.328	0.378	0.378	0.420
Image 110	4.85	7.77	6.51	8.27	6.25	0.138	0.298	0.263	0.322	0.263
Average	4.67	7.01	5.67	6.77	5.98	0.220	0.346	0.339	0.330	0.384

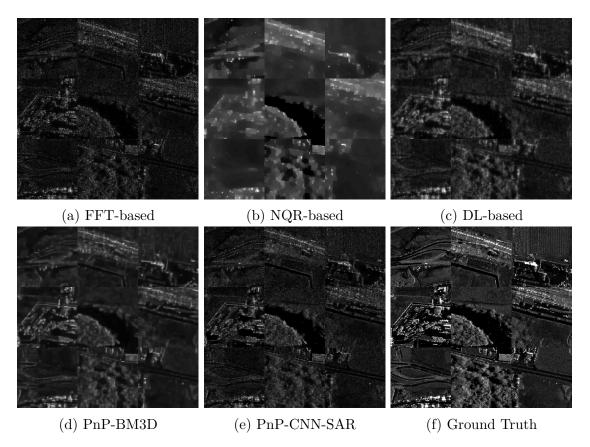


Figure A.30 Reconstruction results for Images 19, 40, 45, 46, 59, 61, 66, 96, and 109 from the Kapıkule test set for the case of 30% data availability and $\sigma_{\mathbf{n}} = 0.1\sigma_{\mathbf{y}}$.

A.3.7 Experiment 7: 100% Data Availability and $\sigma_n = \sigma_y$

Table A.31 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR and best SSIM for all 110 images. Fig. A.31 shows the reconstruction results for a selection of images from the Kapıkule test set from all methods for this experiment. PnP-CNN-SAR produce better reconstructions compared to other methods.

Table A.31 SNR and SSIM values for images from the KapıkuleTest Set for the case of 100% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			SNR (dB)					SSIM		
	FFT-based NC	R-based D	L-based Pr	P-BM3D Pn	P-CNN-SAR F	FT-based NO	QR-based D	L-based Pr	nP-BM3D PnP	-CNN-SAR
Image 1	6.68	8.88	11.36 10.68	6.76	$35.90 \\ 35.52$	0.603	0.593	0.651	0.472	0.998 0.996
mage 2 mage 3	6.45 6.93	$\frac{11.09}{10.64}$	12.18	$6.57 \\ 7.20$	33.40	$0.591 \\ 0.607$	$0.606 \\ 0.534$	$0.638 \\ 0.678$	$0.461 \\ 0.473$	0.997
mage 4	6.84	10.35	11.29	6.97	35.02	0.611	0.507	0.673	0.467	0.998
mage 5	6.07 6.90	8.51 11.60	$9.66 \\ 12.05$	$6.14 \\ 6.99$	$34.62 \\ 36.60$	$0.583 \\ 0.571$	0.500	$0.606 \\ 0.642$	0.454 0.423	0.99 0.99
mage 6 mage 7	6.82	11.00	11.44	6.97	35.33	0.554	0.552 0.513	0.642 0.617	0.415	0.99
mage 8	6.78	10.38	11.07	6.97	36.49	0.579	0.489	0.634	0.433	0.99
mage 9	6.34	9.40	9.99	6.48	37.86	0.606	0.520	0.632	0.475	0.99
mage 10 mage 11	6.24 6.57	10.61 11.45	10.46 11.28	$6.32 \\ 6.65$	$39.05 \\ 37.21$	0.584 0.585	$\frac{0.625}{0.580}$	$0.620 \\ 0.639$	0.453 0.435	$0.99 \\ 0.99$
mage 12	6.09	8.73	9.01	6.28	36.49	0.628	0.560	0.632	0.510	0.99
mage 13	5.82	11.00	9.80	5.96	37.81	0.586	0.644	0.609	0.460	0.99
mage 14 mage 15	$7.03 \\ 7.02$	10.88 10.30	11.89 11.96	$7.11 \\ 7.20$	$35.42 \\ 34.91$	$0.575 \\ 0.627$	$0.447 \\ 0.442$	$0.641 \\ 0.697$	$0.422 \\ 0.475$	0.99 0.99
mage 16	5.20	8.97	7.98	5.29	37.46	0.505	0.592	0.495	0.412	0.99
mage 17	6.91	8.14	12.17	7.05	34.63	0.563	0.523	0.636	0.436	0.99
mage 18 mage 19	$7.04 \\ 6.52$	$\frac{11.26}{9.76}$	12.02 10.80	7.23 6.80	$34.95 \\ 36.63$	$0.611 \\ 0.657$	$0.492 \\ 0.520$	$0.683 \\ 0.695$	$0.458 \\ 0.517$	0.99 0.99
mage 19	6.62	9.41	10.46	6.82	36.07	0.589	0.516	0.638	0.448	0.99
mage 21	6.36	11.04	10.65	6.39	36.49	0.587	0.640	0.633	0.462	0.99
mage 22	5.86	10.37	10.01	6.11	$37.56 \\ 37.62$	0.576	0.637	0.609	0.455	0.99
mage 23 mage 24	6.59 6.86	10.27 10.72	$11.44 \\ 11.72$	$6.80 \\ 7.11$	34.39	$0.638 \\ 0.619$	0.586 0.514	$0.678 \\ 0.704$	$0.481 \\ 0.477$	0.99 0.99
mage 25	7.05	7.40	12.52	7.16	33.36	0.511	0.450	0.588	0.392	0.99
mage 26	6.61	10.46	11.15	6.81	37.78	0.628	0.569	0.682	0.487	0.99
mage 27 mage 28	6.92 6.94	11.09 10.86	$12.00 \\ 12.15$	$6.99 \\ 7.12$	$34.22 \\ 35.45$	$0.621 \\ 0.575$	$0.555 \\ 0.473$	$0.687 \\ 0.647$	$0.472 \\ 0.423$	0.99 0.99
mage 29	6.48	10.70	10.87	6.49	40.56	0.654	0.630	0.692	0.503	0.99
mage 30	6.89	11.22	11.52	6.97	36.75	0.607	0.591	0.654	0.462	0.99
mage 31 mage 32	6.88 6.74	9.68 8.41	11.74 11.43	$7.08 \\ 6.95$	$37.49 \\ 34.52$	$0.667 \\ 0.637$	$0.447 \\ 0.573$	$0.718 \\ 0.676$	0.519 0.504	0.99 0.99
mage 32	6.43	11.20	11.45	6.71	35.15	0.582	0.618	0.640	0.457	0.99
mage 34	5.93	9.75	9.80	5.96	38.83	0.568	0.522	0.596	0.454	0.99
mage 35	6.65	10.37	11.24 10.80	6.69	$\frac{39.07}{37.27}$	0.641	0.616	0.684	0.484	0.99
nage 36 nage 37	6.35 6.84	10.39 8.94	11.98	6.59 7.10	34.54	$0.582 \\ 0.608$	0.516 0.544	$0.639 \\ 0.662$	0.447 0.483	0.99 0.99
nage 38	6.85	9.11	12.06	7.03	34.71	0.560	0.512	0.629	0.433	0.99
nage 39	7.01	7.94	11.90	7.19	34.41	0.598	0.529	0.649	0.486	0.99
mage 40 mage 41	6.73 6.92	9.91 11.37	10.92 11.82	$6.79 \\ 7.08$	$35.18 \\ 34.80$	$0.577 \\ 0.588$	$0.508 \\ 0.492$	$0.623 \\ 0.659$	$0.444 \\ 0.435$	0.99 0.99
mage 42	6.92	10.99	11.82	7.12	35.80	0.606	0.582	0.659	0.469	0.99
mage 43	5.96	11.08	10.11	6.13	37.74	0.562	0.629	0.589	0.434	0.99
mage 44 mage 45	6.95 6.50	10.62 10.12	11.54 10.22	$7.05 \\ 6.66$	35.12 37.77	$0.658 \\ 0.614$	0.537 0.490	$0.708 \\ 0.651$	$0.503 \\ 0.476$	0.99 0.99
mage 46	5.86	7.57	8.70	5.94	35.22	0.597	0.532	0.600	0.486	0.99
mage 47	6.92	11.11	11.65	7.06	35.09	0.642	0.527	0.702	0.494	0.99
mage 48 mage 49	6.55 6.94	10.48 11.83	10.80 12.41	$6.71 \\ 7.04$	$38.28 \\ 34.06$	$0.624 \\ 0.539$	$0.563 \\ 0.442$	$0.669 \\ 0.625$	0.491 0.386	$0.99 \\ 0.99$
mage 49	5.90	10.04	9.89	6.02	38.03	0.564	0.533	0.595	0.442	0.99
mage 51	6.67	10.13	10.94	6.80	36.65	0.629	0.527	0.684	0.490	0.99
mage 52	6.60	11.71 10.59	11.59	6.89	34.36	0.560	0.560	0.634	0.421	0.99
mage 53 mage 54	6.61 6.63	10.59 11.01	10.51 11.32	$6.82 \\ 6.69$	$36.89 \\ 35.48$	$0.599 \\ 0.567$	$0.532 \\ 0.556$	$0.653 \\ 0.628$	$0.457 \\ 0.417$	0.99 0.99
mage 55	6.49	11.38	11.23	6.60	37.56	0.581	0.570	0.635	0.429	0.99
mage 56	6.90	10.80	11.46	7.04	37.47	0.618	0.532	0.673	0.474	0.99
mage 57 mage 58	6.71 6.38	9.12 10.50	10.73 10.79	6.87 6.53	$36.95 \\ 38.13$	$0.624 \\ 0.591$	$0.441 \\ 0.522$	$0.670 \\ 0.636$	$0.486 \\ 0.457$	0.99 0.99
mage 59	6.02	10.93	9.95	6.03	36.17	0.550	0.642	0.581	0.430	0.99
mage 60	6.92	10.53	11.45	7.03	36.02	0.611	0.509	0.669	0.465	0.99
mage 61 mage 62	6.48 6.77	9.13	10.50	6.54 6.88	$35.48 \\ 35.98$	$0.599 \\ 0.625$	0.490	$0.634 \\ 0.673$	$0.457 \\ 0.471$	0.99 0.99
mage 63	6.89	10.71 10.66	11.68 11.90	7.12	34.41	0.595	$0.572 \\ 0.571$	0.666	0.465	0.99
mage 64	6.77	11.20	11.12	6.93	35.94	0.607	0.533	0.661	0.464	0.99
mage 65	6.86	10.13	11.14	6.98	35.57	0.561	0.464	0.626	0.423	0.99
nage 66 nage 67	6.59 5.86	$10.66 \\ 9.95$	10.47 9.81	$6.78 \\ 5.92$	$37.58 \\ 39.72$	0.584 0.593	$0.539 \\ 0.643$	$\frac{0.631}{0.613}$	$0.444 \\ 0.473$	0.99 0.99
nage 68	7.02	10.88	11.82	7.18	35.15	0.636	0.522	0.697	0.489	0.99
mage 69	6.88	9.65	11.79	7.02	36.79	0.585	0.551	0.648	0.449	0.99
mage 70 mage 71	6.96 6.24	9.04 10.55	$11.92 \\ 10.62$	$7.08 \\ 6.43$	$35.09 \\ 36.60$	$0.590 \\ 0.579$	$0.540 \\ 0.511$	$0.655 \\ 0.621$	$0.460 \\ 0.438$	0.99 0.99
nage 71	6.81	8.93	10.02	6.94	34.45	0.552	0.445	0.605	0.425	0.99
nage 73	6.40	9.01	10.12	6.55	35.99	0.673	0.514	0.705	0.531	0.99
nage 74	6.95	11.72	11.88	6.98	35.78	0.571	0.541	0.641	0.418	0.99
nage 75 nage 76	6.98 6.89	7.91 10.47	$12.04 \\ 11.61$	$7.18 \\ 7.01$	$34.43 \\ 35.91$	$0.574 \\ 0.625$	$0.501 \\ 0.505$	$0.640 \\ 0.689$	$0.448 \\ 0.467$	0.99 0.99
nage 70	6.91	11.36	12.06	7.06	35.28	0.571	0.511	0.640	0.433	0.99
nage 78	6.76	9.95	11.06	6.82	36.20	0.660	0.517	0.706	0.511	0.99
nage 79	6.35	10.50	10.79	6.51	35.97	0.584	0.485	0.629	0.450	0.99
mage 80 mage 81	$6.68 \\ 7.07$	10.56 12.19	$\frac{11.32}{11.96}$	$6.83 \\ 7.24$	$37.44 \\ 35.30$	$0.642 \\ 0.507$	$0.494 \\ 0.417$	$0.685 \\ 0.598$	0.484 0.358	0.99
mage 82	6.02	9.57	9.09	6.10	37.10	0.576	0.542	0.599	0.452	0.99
mage 83	6.81	10.66	11.46	6.96	35.13	0.620	0.588	0.665	0.490	0.99
mage 84 mage 85	6.35	10.85 10.09	10.59	$6.52 \\ 6.62$	36.69 36.55	0.579	0.565	0.631	0.450	0.99
mage 85 mage 86	$6.52 \\ 6.14$	8.13	$10.12 \\ 9.42$	6.62	36.55 38.38	$0.618 \\ 0.635$	$0.503 \\ 0.497$	$0.646 \\ 0.644$	$0.471 \\ 0.505$	0.99 0.99
mage 87	6.17	8.90	9.71	6.22	38.73	0.624	0.496	0.641	0.487	0.99
mage 88	6.93	11.56	12.05	7.16	35.55	0.596	0.480	0.660	0.448	0.9

			Tab	le A.31 con	tinued from	previous p	age			
			SNR (dB)		1			SSIM		
Image 89	6.40	10.43	10.85	6.31	39.57	0.637	0.601	0.675	0.480	0.999
Image 90	6.93	10.84	11.85	7.19	34.87	0.610	0.577	0.670	0.474	0.998
Image 91	6.76	10.63	10.89	6.88	36.79	0.578	0.488	0.629	0.433	0.998
Image 92	6.92	10.68	11.77	6.99	35.59	0.599	0.494	0.674	0.438	0.998
Image 93	6.49	9.83	10.25	6.69	36.24	0.625	0.518	0.649	0.481	0.999
Image 94	6.83	11.17	11.56	6.98	37.23	0.620	0.540	0.685	0.476	0.999
Image 95	6.99	11.57	11.97	7.14	35.51	0.613	0.539	0.688	0.468	0.998
Image 96	6.39	9.69	10.45	6.38	39.15	0.644	0.531	0.672	0.494	0.998
Image 97	6.19	10.98	10.29	6.23	37.06	0.564	0.654	0.611	0.441	0.995
Image 98	7.01	11.30	12.13	7.12	35.58	0.594	0.541	0.668	0.446	0.998
Image 99	6.71	11.07	11.56	6.94	36.77	0.622	0.571	0.668	0.474	0.999
Image 100	5.37	8.29	7.97	5.45	37.08	0.557	0.537	0.542	0.460	0.994
Image 101	6.69	10.95	11.47	6.87	35.76	0.597	0.579	0.651	0.463	0.998
Image 102	6.77	10.67	11.54	6.96	36.06	0.592	0.542	0.648	0.463	0.998
Image 103	6.98	12.59	12.67	7.08	34.44	0.483	0.384	0.569	0.335	0.996
Image 104	6.86	7.93	11.50	6.91	35.06	0.649	0.582	0.687	0.519	0.998
Image 105	5.79	10.85	9.70	5.91	38.33	0.557	0.673	0.586	0.446	0.994
Image 106	6.92	10.65	11.34	7.06	36.97	0.587	0.534	0.663	0.448	0.998
Image 107	6.65	10.53	11.04	6.64	38.98	0.678	0.621	0.707	0.521	0.999
Image 108	6.85	11.01	11.64	6.97	35.39	0.616	0.522	0.685	0.467	0.998
Image 109	6.89	10.29	10.98	6.89	35.42	0.596	0.514	0.657	0.448	0.998
Image 110	7.03	9.26	12.54	7.26	34.55	0.514	0.453	0.590	0.376	0.997
Arronogo	6.62	10.29	11.09	6.75	26 27	0.507	0.527	0.647	0.450	0.007

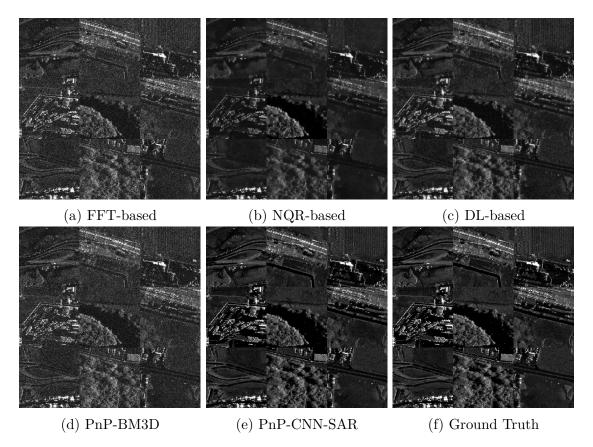


Figure A.31 Reconstruction results for Images 19, 40, 45, 46, 59, 61, 66, 96, and 109 from the Kapıkule test set for the case of 100% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$.

A.3.8 Experiment 8: 90% Data Availability and $\sigma_n = \sigma_y$

Table A.32 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR and best SSIM for all 110 images. Fig. A.32 shows the reconstruction results for a selection of images from the Kapıkule test set from all methods for this experiment. PnP-CNN-SAR produce better reconstructions compared to other methods.

Table A.32 SNR and SSIM values for images from the KapıkuleTest Set for the case of 90% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			SNR (dB)					SSIM		
	FFT-based NQ	R-based D	L-based Pn	P-BM3D Pn	P-CNN-SAR FI	T-based NO	QR-based D	L-based Pr	nP-BM3D PnF	P-CNN-SAR
mage 1	6.61	9.37	$10.88 \\ 9.94$	6.77	15.49 15.49	0.514	0.555	0.589	$0.427 \\ 0.421$	0.846
mage 2 mage 3	6.28 6.89	9.82 10.07	9.94 11.61	$6.51 \\ 7.43$	15.49	$0.508 \\ 0.523$	$0.447 \\ 0.509$	$0.584 \\ 0.613$	0.440	$0.854 \\ 0.823$
mage 4	6.71	9.55	10.32	7.06	15.48	0.531	0.470	0.614	0.441	0.851
mage 5 mage 6	5.89 6.89	7.93 11.16	$8.20 \\ 11.62$	5.17 7.19	$16.19 \\ 16.07$	$0.519 \\ 0.474$	$0.453 \\ 0.446$	$0.531 \\ 0.583$	$0.392 \\ 0.385$	0.870 0.839
mage 7	6.78	10.59	10.48	6.90	15.26	0.484	0.446 0.476	0.567	0.381	0.829
mage 8	6.69	9.74	9.70	6.87	15.22	0.510	0.449	0.576	0.397	0.838
mage 9	6.14	8.20	8.54	6.06	15.54	0.532	0.446	0.558	0.428	0.870
mage 10 mage 11	$6.01 \\ 6.40$	9.56 10.84	$\frac{9.79}{10.82}$	$6.12 \\ 6.74$	$15.51 \\ 15.79$	$0.506 \\ 0.493$	$0.474 \\ 0.531$	$0.563 \\ 0.583$	$0.406 \\ 0.399$	$0.868 \\ 0.854$
mage 12	5.83	7.85	7.18	5.66	15.66	0.570	0.498	0.548	0.464	0.89
mage 13	5.70	9.33	8.95	5.76	15.96	0.510	0.615	0.552	0.420	0.883
mage 14 mage 15	6.98 6.93	11.12 10.87	11.13 11.26	7.33 7.35	$14.65 \\ 14.91$	$0.492 \\ 0.543$	0.456 0.489	$0.580 \\ 0.642$	0.394 0.439	0.80
mage 16	5.00	8.15	6.51	4.68	16.47	0.457	0.489 0.477	0.433	0.377	0.90
mage 17	6.90	8.10	11.78	7.43	15.24	0.472	0.482	0.575	0.413	0.80
mage 18	6.93	10.77	11.44	7.44	15.01	0.526	0.460	0.632	0.433	0.83
mage 19 mage 20	6.42 6.45	9.37 9.00	$\frac{9.75}{8.90}$	$6.63 \\ 6.52$	15.23 15.78	$0.582 \\ 0.518$	$0.478 \\ 0.484$	$0.629 \\ 0.573$	$0.473 \\ 0.415$	$0.878 \\ 0.858$
mage 20	6.19	9.55	9.71	6.32	15.56	0.511	0.469	0.573	0.426	0.86
mage 22	5.81	9.20	9.18	5.88	15.82	0.506	0.447	0.547	0.410	0.88
mage 23	6.51	10.42	10.80	6.73	15.29	0.535	0.522	0.606	0.428	0.85
mage 24 mage 25	6.77 6.93	$\frac{10.89}{6.32}$	10.60 12.51	$7.14 \\ 7.63$	$15.41 \\ 15.02$	$0.542 \\ 0.415$	$0.501 \\ 0.386$	$0.637 \\ 0.542$	0.446 0.369	0.853
mage 26	6.47	9.93	10.08	6.69	15.24	0.550	0.517	0.617	0.449	0.86
mage 27	6.84	10.60	11.56	7.22	15.50	0.528	0.515	0.632	0.436	0.84
mage 28	6.91	10.85	11.72	7.25	15.28	0.481	0.504	0.588	0.388	0.81
mage 29 mage 30	6.30 6.70	9.29 10.37	9.97 10.86	$6.48 \\ 6.95$	$15.26 \\ 15.07$	$0.569 \\ 0.512$	$0.561 \\ 0.538$	$0.610 \\ 0.594$	$0.457 \\ 0.420$	0.888
mage 31	6.90	10.26	11.07	7.14	14.48	0.578	0.548	0.648	0.471	0.84
mage 32	6.66	6.50	10.90	7.01	15.03	0.543	0.530	0.607	0.466	0.848
mage 33	6.40	9.98	10.50	6.67	16.01	0.495	0.465	0.584	0.421	0.86
mage 34 mage 35	5.75 6.60	$\frac{9.96}{10.24}$	8.95 10.63	$5.74 \\ 6.79$	15.91 15.19	$0.500 \\ 0.548$	$\frac{0.650}{0.576}$	0.535 0.616	$0.409 \\ 0.442$	0.888 0.864
mage 36	6.23	10.07	9.93	6.50	16.31	0.497	0.464	0.581	0.408	0.87
nage 37	6.91	10.22	11.33	7.16	14.89	0.517	0.540	0.591	0.432	0.82
nage 38 nage 39	$6.74 \\ 6.82$	$9.40 \\ 7.07$	11.47 11.40	$7.20 \\ 7.38$	15.33 15.04	$0.466 \\ 0.513$	$0.454 \\ 0.490$	$0.559 \\ 0.585$	0.397 0.458	$0.814 \\ 0.816$
nage 39	6.60	8.68	9.93	6.60	14.93	0.513	0.364	0.560	0.403	0.849
mage 41	6.86	11.29	11.17	7.46	15.45	0.505	0.471	0.610	0.417	0.83
mage 42	6.94	10.46	11.18	7.27	15.04	0.514	0.542	0.586	0.431	0.82
mage 43 mage 44	$\frac{5.86}{6.74}$	$9.46 \\ 10.11$	$9.52 \\ 10.83$	$6.00 \\ 6.98$	$15.60 \\ 14.91$	$0.483 \\ 0.571$	$0.450 \\ 0.509$	$0.536 \\ 0.658$	$0.400 \\ 0.465$	0.872 0.859
mage 45	6.26	9.34	8.81	6.24	15.61	0.540	0.442	0.578	0.427	0.86
mage 46	5.58	6.81	6.74	5.07	16.40	0.544	0.480	0.505	0.435	0.898
mage 47	6.86	10.45	10.81	7.17	14.98	0.560	0.483	0.640	0.458	0.848
mage 48 mage 49	6.37 6.99	9.30 11.51	$9.79 \\ 12.31$	$6.47 \\ 7.52$	$15.14 \\ 16.00$	$0.548 \\ 0.450$	$0.530 \\ 0.486$	$0.601 \\ 0.576$	$0.446 \\ 0.370$	$0.872 \\ 0.817$
nage 49	5.67	9.42	8.87	5.74	16.16	0.494	0.503	0.537	0.405	0.87
mage 51	6.55	9.63	9.65	6.75	15.29	0.553	0.496	0.617	0.451	0.86
mage 52	6.60	11.01	11.00	7.05	16.05	0.476	0.534	0.577	0.392	0.83
mage 53 mage 54	$6.47 \\ 6.42$	$9.54 \\ 10.64$	8.83 10.58	6.56 6.78	$15.47 \\ 15.77$	$0.529 \\ 0.474$	$0.487 \\ 0.456$	$0.580 \\ 0.569$	$0.427 \\ 0.382$	$0.85' \\ 0.840$
mage 55	6.42	10.49	10.79	6.64	15.55	0.495	0.455	0.581	0.389	0.848
mage 56	6.73	9.06	10.61	7.06	15.17	0.537	0.458	0.616	0.438	0.852
mage 57	6.53	9.30	9.47	6.84	15.27	0.544	0.458	0.604	0.448	0.85
mage 58 mage 59	6.19 5.81	$\frac{10.08}{9.99}$	10.03 9.15	$6.49 \\ 5.94$	15.86 15.94	$0.505 \\ 0.477$	$0.580 \\ 0.609$	$0.572 \\ 0.527$	0.415 0.393	$0.86' \\ 0.87'$
mage 60	6.79	10.35	10.65	7.09	14.90	0.525	0.482	0.614	0.428	0.829
nage 61	6.31	7.38	9.52	6.45	14.93	0.527	0.336	0.573	0.423	0.85'
nage 62	6.65	10.16	11.10	$7.06 \\ 7.31$	$14.96 \\ 15.61$	0.524	0.459	0.599	0.425	0.83
mage 63 mage 64	6.79 6.69	$\frac{10.46}{9.64}$	11.34 10.15	6.93	15.61	$0.501 \\ 0.528$	$0.526 \\ 0.460$	$0.605 \\ 0.598$	$0.433 \\ 0.430$	0.840
nage 65	6.72	9.85	10.11	6.87	15.23	0.493	0.459	0.567	0.383	0.82
mage 66	6.39	9.60	8.97	6.46	15.63	0.521	0.484	0.565	0.411	0.86
nage 67 nage 68	5.68 6.91	$\frac{9.81}{10.17}$	8.98 11.09	$\frac{5.81}{7.21}$	$16.04 \\ 14.53$	0.518 0.559	$\frac{0.609}{0.478}$	$0.548 \\ 0.643$	$0.430 \\ 0.455$	$0.90 \\ 0.83$
nage 69	6.75	10.17	11.09 11.34	7.21	15.61	0.559 0.502	$0.478 \\ 0.534$	0.643 0.593	0.403	0.83
nage 70	6.86	8.88	11.44	7.25	15.27	0.494	0.458	0.586	0.415	0.82
nage 71	6.13	9.69	9.75	6.33	16.04	0.498	0.476	0.559	0.400	0.86
nage 72	6.66	8.79	9.55	6.72	14.93	0.482	0.450	0.540	0.390	0.82
nage 73 nage 74	6.28 6.78	$\frac{8.62}{11.06}$	$8.59 \\ 11.40$	$6.27 \\ 7.27$	15.09 15.47	$0.604 \\ 0.485$	$0.482 \\ 0.503$	$0.623 \\ 0.586$	$0.492 \\ 0.390$	$0.89 \\ 0.82$
nage 75	6.93	7.72	11.68	7.36	15.01	0.486	0.529	0.592	0.425	0.81
nage 76	6.80	10.57	10.68	7.18	15.24	0.537	0.491	0.625	0.439	0.84
nage 77	6.88	10.87	11.71	7.34	15.41	0.477	0.508	0.583	0.404	0.81
nage 78 nage 79	6.60 6.26	9.86 10.59	$\frac{10.04}{10.07}$	$6.84 \\ 6.61$	$15.21 \\ 15.74$	$0.577 \\ 0.501$	$0.501 \\ 0.530$	$0.634 \\ 0.573$	$0.476 \\ 0.415$	$0.87 \\ 0.85$
mage 80	6.62	9.77	10.32	6.83	15.04	0.554	0.420	0.607	0.436	0.85
mage 81	6.94	11.55	11.12	7.29	15.65	0.431	0.404	0.551	0.339	0.78
mage 82	5.72	8.10	7.24	5.43	15.91	0.521	0.477	0.520	0.413	0.89
mage 83 mage 84	6.64 6.15	$9.69 \\ 9.67$	$10.77 \\ 9.69$	6.95 6.28	$14.79 \\ 15.56$	$0.533 \\ 0.509$	$0.565 \\ 0.527$	$0.605 \\ 0.583$	$0.449 \\ 0.416$	0.83 0.86
mage 85	6.38	8.93	8.41	6.33	15.28	0.548	0.327	0.572	0.410	0.86
	5.97	7.78	7.85	5.72	15.64	0.564	0.450	0.561	0.447	0.89
mage 86 mage 87	6.01	8.81	8.27	5.90	15.57	0.552	0.462	0.558	0.445	0.88

			Tab	ole A.32 con	tinued from	previous p	age			
			SNR (dB)					SSIM		
Image 89	6.19	9.60	10.04	6.33	15.56	0.548	0.552	0.605	0.438	0.881
Image 90	6.77	10.32	11.28	7.33	15.04	0.518	0.552	0.613	0.439	0.827
Image 91	6.65	10.06	9.66	6.69	15.17	0.507	0.447	0.565	0.394	0.835
Image 92	6.85	10.92	11.01	7.21	15.38	0.517	0.483	0.623	0.420	0.837
Image 93	6.34	8.58	8.88	6.47	15.33	0.556	0.450	0.580	0.447	0.874
Image 94	6.69	9.42	10.67	7.01	15.40	0.541	0.443	0.617	0.439	0.855
Image 95	6.81	10.83	11.36	7.30	15.37	0.529	0.511	0.636	0.439	0.851
Image 96	6.16	8.13	9.36	6.24	15.73	0.558	0.426	0.592	0.448	0.892
Image 97	6.01	10.24	9.50	6.00	15.92	0.498	0.632	0.559	0.404	0.873
Image 98	6.89	11.01	11.72	7.43	15.42	0.501	0.523	0.615	0.407	0.829
Image 99	6.61	9.97	10.99	6.94	15.28	0.528	0.469	0.607	0.426	0.847
Image 100	5.05	7.22	6.10	4.91	16.42	0.504	0.461	0.454	0.422	0.903
Image 101	6.66	10.20	10.88	6.88	14.93	0.517	0.539	0.598	0.426	0.834
Image 102	6.58	10.68	10.89	6.95	15.42	0.509	0.561	0.589	0.419	0.849
Image 103	7.06	9.61	12.62	7.37	16.16	0.388	0.380	0.510	0.305	0.776
Image 104	6.69	7.37	10.90	6.90	14.84	0.564	0.526	0.622	0.475	0.850
Image 105	5.65	9.27	8.73	5.68	16.13	0.490	0.487	0.524	0.401	0.888
Image 106	6.72	9.85	10.18	7.00	15.14	0.513	0.494	0.596	0.421	0.837
Image 107	6.52	9.88	9.99	6.69	14.86	0.583	0.570	0.618	0.476	0.875
Image 108	6.77	10.41	10.96	7.10	15.13	0.539	0.491	0.628	0.439	0.849
Image 109	6.65	9.41	9.72	7.01	15.24	0.526	0.478	0.603	0.433	0.849
Image 110	7.06	10.08	12.41	7.53	15.86	0.410	0.413	0.531	0.340	0.789
Average	6.49	9.65	10.21	6.71	15.43	0.516	0.492	0.584	0.422	0.851

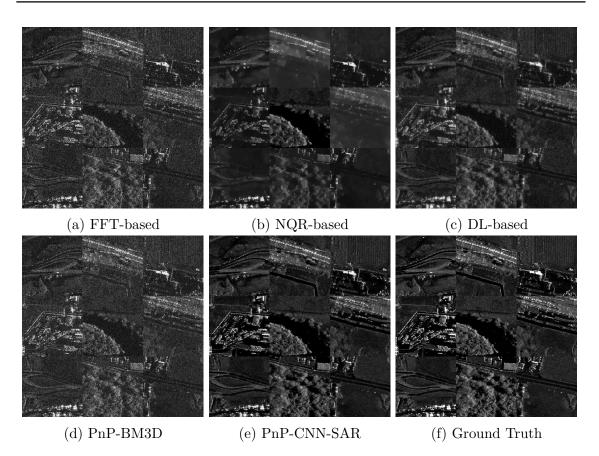


Figure A.32 Reconstruction results for Images 19, 40, 45, 46, 59, 61, 66, 96, and 109 from the Kapıkule test set for the case of 90% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$.

A.3.9 Experiment 9: 80% Data Availability and $\sigma_n = \sigma_y$

Table A.33 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR and best SSIM for all 110 images. Fig. A.33 shows the reconstruction results for a selection of images from the Kapıkule test set from all methods for this experiment. PnP-CNN-SAR produce better reconstructions compared to other methods.

Table A.33 SNR and SSIM values for images from the KapıkuleTest Set for the case of 80% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			SNR (dB)					SSIM		
	FFT-based NQI	R-based D	L-based Pı	nP-BM3D Pi	ıP-CNN-SAR	FFT-based	NQR-based I	OL-based F	PnP-BM3D I	PnP-CNN-SAR
Image 1	6.50	9.56	10.22	6.93	12.93	0.459	0.515	0.512	0.409	0.727
Image 2 Image 3	6.19 6.85	8.46 8.11	9.36 11.03	$6.59 \\ 7.69$	13.31 12.74	$0.466 \\ 0.475$	$0.400 \\ 0.432$	$0.525 \\ 0.552$	0.394 0.433	0.764 0.703
Image 3	6.71	9.41	9.42	7.10	12.64	0.487	0.444	0.546	0.418	0.749
Image 5	5.66	6.32	7.17	5.47	9.11	0.489	0.295	0.459	0.387	0.681
Image 6	6.77	9.85	11.18	7.30	13.32	0.421	0.455	0.520	0.367	0.706
Image 7	6.63	9.25	9.65	6.99	12.84	0.441	0.419	0.506	0.366	0.735
Image 8 Image 9	$6.58 \\ 5.92$	$7.05 \\ 7.14$	$8.94 \\ 7.66$	6.95 6.08	12.71 12.99	0.467 0.493	0.323 0.379	$\frac{0.523}{0.490}$	0.388 0.401	0.750 0.782
mage 10	5.98	9.04	9.18	6.17	13.17	0.454	0.420	0.503	0.378	0.782
Image 11	6.42	10.02	10.33	6.72	12.97	0.446	0.481	0.519	0.362	0.710
Image 12	5.52	5.66	6.08	5.66	12.84	0.531	0.436	0.466	0.445	0.815
Image 13	5.64	8.64	8.26	5.82	13.37	0.469	0.443	0.491	0.398	0.797
Image 14 Image 15	$6.85 \\ 6.85$	$9.05 \\ 9.65$	$10.47 \\ 10.52$	7.34 7.64	12.72 13.03	0.443 0.489	$0.338 \\ 0.422$	$0.521 \\ 0.589$	$0.370 \\ 0.430$	0.706 0.748
mage 16	4.82	5.71	5.52	4.75	13.80	0.430	0.435	0.369	0.362	0.832
mage 17	6.80	9.50	11.33	7.66	13.26	0.419	0.493	0.511	0.400	0.690
Image 18	6.87	9.91	10.73	7.64	13.06	0.472	0.455	0.574	0.413	0.740
Image 19	6.36	8.75	8.85	6.71	12.88	0.530	0.433	0.556	0.449	0.792
Image 20	$6.22 \\ 6.02$	$8.87 \\ 9.65$	7.97 9.13	$6.60 \\ 6.41$	13.27 13.17	0.484 0.476	0.467	0.525	0.400	0.750 0.779
Image 21 Image 22	5.70	9.05 8.83	9.13 8.50	6.02	13.17	0.476	$\frac{0.575}{0.430}$	0.513 0.482	0.403 0.391	0.796
mage 23	6.51	9.85	10.16	6.85	13.09	0.481	0.476	0.482 0.527	0.406	0.750
Image 24	6.69	9.99	9.85	7.45	13.36	0.496	0.472	0.584	0.433	0.765
mage 25	6.92	11.47	12.09	7.76	12.75	0.369	0.403	0.476	0.348	0.625
mage 26	6.24	9.08	9.33	6.72	13.04	0.510	0.462	0.554	0.423	0.780
mage 27	6.76	10.30 11.01	10.94 11.19	7.52 7.49	$13.28 \\ 13.47$	0.479	0.409	0.563	0.424 0.363	0.741
mage 28 mage 29	$6.86 \\ 6.22$	9.34	9.31	6.48	12.71	$0.426 \\ 0.507$	$0.408 \\ 0.520$	$0.529 \\ 0.536$	0.303 0.428	0.713 0.793
mage 30	6.58	9.89	$\frac{9.31}{10.24}$	7.10	13.17	0.460	0.500	0.528	0.428	0.750
mage 31	6.76	9.13	10.37	7.34	12.52	0.523	0.364	0.570	0.443	0.756
mage 32	6.53	6.46	10.27	7.19	12.78	0.484	0.428	0.531	0.445	0.729
mage 33	6.29	9.09	9.87	6.77	13.26	0.448	0.431	0.518	0.397	0.749
mage 34	5.62	9.35	8.26	5.75	13.19	0.457	0.615	0.476	0.383	0.816
mage 35 mage 36	6.44 6.10	$9.71 \\ 9.36$	$\frac{9.99}{9.23}$	$6.80 \\ 6.53$	12.99 13.49	$0.491 \\ 0.457$	$0.519 \\ 0.422$	$0.541 \\ 0.516$	0.413 0.385	$0.761 \\ 0.777$
mage 37	6.77	7.12	10.76	7.51	12.86	0.463	0.444	0.516	0.421	0.702
mage 38	6.78	9.06	10.94	7.36	13.22	0.423	0.472	0.491	0.381	0.692
mage 39	6.85	9.01	10.82	7.57	12.87	0.453	0.488	0.516	0.438	0.699
mage 40	6.41	7.58	9.02	6.63	12.54	0.467	0.327	0.494	0.378	0.755
mage 41	6.81	$\frac{10.82}{10.10}$	10.58	7.51 7.28	13.56 12.88	0.464	0.450	0.552	0.392	0.735 0.716
mage 42 mage 43	$6.81 \\ 5.74$	9.19	$\frac{10.57}{8.90}$	5.93	13.22	$0.461 \\ 0.441$	$0.490 \\ 0.574$	$\frac{0.516}{0.477}$	$0.403 \\ 0.372$	0.791
mage 44	6.67	9.93	9.97	7.19	12.72	0.531	0.493	0.596	0.457	0.771
mage 45	6.03	5.82	7.79	6.40	13.15	0.501	0.256	0.506	0.422	0.770
mage 46	5.27	6.05	5.74	5.26	13.76	0.511	0.421	0.442	0.426	0.835
mage 47	6.84	8.63	10.04	7.41	12.88	0.510	0.368	0.576	0.436	0.758
mage 48 mage 49	6.22 6.88	$8.80 \\ 7.72$	9.13 11.80	6.63 7.73	12.74 13.61	$0.506 \\ 0.398$	$0.443 \\ 0.439$	$0.544 \\ 0.517$	$0.431 \\ 0.353$	0.786 0.693
mage 50	5.63	7.84	8.20	5.80	13.31	0.457	0.417	0.491	0.384	0.807
mage 51	6.39	8.93	8.87	6.90	12.96	0.510	0.455	0.548	0.432	0.778
mage 52	6.68	10.23	10.56	7.25	13.87	0.421	0.452	0.521	0.368	0.736
mage 53	6.24	8.00	7.85	6.67	12.97	0.488	0.425	0.527	0.416	0.769
mage 54	6.46	9.92 10.13	$10.05 \\ 10.22$	6.90 6.66	13.62 13.27	0.445 0.442	$0.407 \\ 0.427$	$0.519 \\ 0.512$	0.366	0.747 0.739
mage 55 mage 56	6.24 6.63	9.04	9.79	7.09	13.04	0.442	0.427	0.512 0.552	$0.359 \\ 0.415$	0.768
mage 57	6.40	8.43	8.47	6.80	12.77	0.502	0.414	0.536	0.420	0.765
mage 58	6.16	9.92	9.40	6.50	13.37	0.454	0.531	0.506	0.386	0.777
mage 59	5.67	8.45	8.62	5.92	13.33	0.436	0.542	0.473	0.367	0.788
mage 60	6.65	8.37	9.84	7.21	13.10	0.482	0.380	0.557	0.413	0.753
mage 61 mage 62	6.24 6.69	$8.05 \\ 10.07$	8.58 10.44	$6.45 \\ 7.03$	11.92 12.81	$0.488 \\ 0.467$	0.387 0.481	$0.507 \\ 0.516$	0.396 0.393	$0.731 \\ 0.724$
mage 63	6.77	9.80	10.44	7.56	13.47	0.460	0.490	0.547	0.418	0.733
mage 64	6.52	9.40	9.46	7.00	13.35	0.484	0.428	0.540	0.408	0.766
mage 65	6.56	6.45	9.21	6.88	9.90	0.450	0.338	0.507	0.370	0.543
mage 66	6.32	7.46	8.05	6.48	13.13	0.484	0.348	0.520	0.398	0.782
mage 67 mage 68	5.59 6.80	8.24 9.83	$8.30 \\ 10.27$	5.83 7.53	13.22 12.54	0.477 0.510	$0.378 \\ 0.459$	$0.479 \\ 0.579$	$0.400 \\ 0.446$	0.822 0.748
mage 69	6.66	10.14	10.27	7.33 7.24	13.46	0.310 0.451	0.439	0.531	0.381	0.734
mage 70	6.79	8.15	10.95	7.51	13.06	0.449	0.465	0.523	0.394	0.706
mage 71	6.02	9.14	9.18	6.39	13.56	0.445	0.420	0.505	0.383	0.780
mage 72	6.56	7.71	8.61	6.74	12.31	0.456	0.388	0.481	0.370	0.735
mage 73 mage 74	6.09	7.32	7.55	6.32	13.02	0.562	0.423	0.552	0.473	0.825
mage 74 mage 75	$6.73 \\ 6.84$	$10.65 \\ 8.27$	10.81 11.11	7.32 7.61	13.64 13.15	$0.432 \\ 0.447$	$0.422 \\ 0.507$	$0.526 \\ 0.531$	0.364 0.409	0.728 0.708
mage 76	6.70	10.04	10.01	7.31	13.32	0.494	0.466	0.574	0.420	0.762
mage 77	6.88	10.01	11.15	7.51	13.31	0.429	0.426	0.513	0.377	0.703
mage 78	6.53	9.04	9.22	6.95	12.84	0.533	0.436	0.563	0.451	0.783
mage 79	6.20	9.88	9.49	6.57	13.11	0.458	0.500	0.503	0.386	0.749
mage 80	6.51	9.16	9.62	6.85	12.94 13.85	0.500	0.366	0.533	0.408	0.761 0.693
mage 81 mage 82	6.84 5.50	$10.84 \\ 6.68$	$10.40 \\ 6.24$	$7.40 \\ 5.47$	13.85 12.66	0.390 0.484	$0.392 \\ 0.415$	$\frac{0.507}{0.458}$	0.318 0.400	0.792
mage 83	6.61	9.05	9.96	7.11	12.81	0.490	0.529	0.438 0.531	0.432	0.739
mage 84	6.06	9.15	9.04	6.44	13.10	0.479	0.486	0.540	0.411	0.778
mage 85	6.17	7.36	7.44	6.31	12.73	0.507	0.373	0.503	0.407	0.751
mage 86	5.76	6.59	6.78	5.75	13.09	0.526	0.371	0.481	0.428	0.818
mage 87 mage 88	$5.92 \\ 6.81$	7.83	7.31	5.97	13.24		0.396	0.485	0.419	0.819
	6.81	9.51	10.94	7.46	13.11	0.448	0.405	0.526	0.384	0.71

			Tab	le A.33 con	tinued from	previous p	age			
			SNR (dB)					SSIM		
Image 89	6.12	8.76	9.43	6.42	12.85	0.499	0.387	0.530	0.408	0.779
Image 90	6.83	9.61	10.60	7.56	12.98	0.472	0.514	0.557	0.428	0.726
Image 91	6.37	6.86	8.68	6.85	12.74	0.461	0.309	0.509	0.382	0.743
Image 92	6.73	10.04	10.22	7.49	13.41	0.469	0.449	0.565	0.407	0.751
Image 93	6.21	7.80	7.99	6.35	8.33	0.512	0.404	0.515	0.418	0.598
Image 94	6.65	8.49	9.95	7.20	13.27	0.490	0.328	0.555	0.416	0.768
Image 95	6.87	9.72	10.73	7.46	12.98	0.486	0.483	0.581	0.421	0.744
Image 96	6.08	8.17	8.69	6.27	13.24	0.504	0.389	0.523	0.418	0.804
Image 97	5.93	9.24	8.91	6.15	13.16	0.459	0.451	0.504	0.391	0.782
Image 98	6.96	9.36	11.36	7.70	13.18	0.454	0.476	0.561	0.390	0.718
Image 99	6.58	8.61	10.37	6.99	13.17	0.474	0.439	0.533	0.400	0.752
Image 100	4.89	6.51	5.11	4.78	13.97	0.473	0.414	0.386	0.398	0.845
Image 101	6.51	9.08	10.05	7.05	12.89	0.471	0.524	0.528	0.408	0.740
Image 102	6.56	9.55	10.26	7.06	12.97	0.458	0.419	0.526	0.398	0.739
Image 103	7.00	9.27	12.27	7.35	13.16	0.333	0.347	0.443	0.266	0.598
Image 104	6.62	6.62	10.13	7.13	12.66	0.499	0.471	0.535	0.455	0.740
Image 105	5.51	8.68	8.13	5.67	13.67	0.450	0.464	0.469	0.382	0.824
Image 106	6.66	9.55	9.32	7.14	13.26	0.484	0.463	0.541	0.409	0.753
Image 107	6.34	8.50	9.31	6.65	12.46	0.532	0.423	0.532	0.439	0.777
Image 108	6.68	9.98	10.22	7.24	13.01	0.497	0.457	0.572	0.420	0.760
Image 109	6.54	6.15	8.68	6.98	12.79	0.488	0.366	0.544	0.420	0.764
Image 110	7.04	8.50	12.16	7.69	13.37	0.366	0.402	0.475	0.320	0.631
Average	6.39	8.79	9.50	6.82	12.98	0.471	0.434	0.520	0.401	0.749

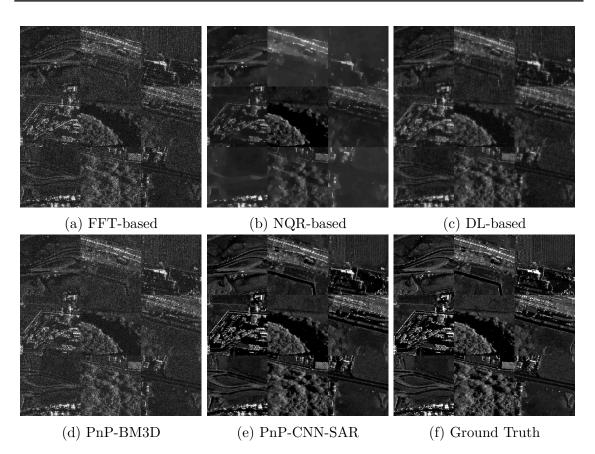


Figure A.33 Reconstruction results for Images 19, 40, 45, 46, 59, 61, 66, 96, and 109 from the Kapıkule test set for the case of 80% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$.

A.3.10 Experiment 10: 70% Data Availability and $\sigma_n = \sigma_y$

Table A.34 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR for 107 images, and best SSIM for all 110 images. Fig. A.34 shows the reconstruction results for a selection of images from the Kapıkule test set from all methods for this experiment. PnP-CNN-SAR produce better reconstructions compared to other methods.

Table A.34 SNR and SSIM values for images from the KapıkuleTest Set for the case of 70% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			SNR (dB))		SSIM					
	FFT-based NQ	R-based D	L-based P	nP-BM3D I	PnP-CNN-SAR	FFT-based	NQR-based I	DL-based I	PnP-BM3D P	nP-CNN-SAR	
Image 1	6.41	9.38	9.73	7.33	11.01	0.406	0.475	0.447	0.397	0.600	
Image 2 Image 3	6.12 6.81	$8.81 \\ 10.64$	8.76 10.53	6.76 7.96	11.09 11.20	$0.422 \\ 0.412$	$0.389 \\ 0.437$	$0.466 \\ 0.501$	$0.373 \\ 0.404$	$0.656 \\ 0.585$	
Image 4	6.54	7.83	8.80	7.38	10.90	0.451	0.365	0.504	0.405	0.656	
Image 5	5.48	6.04	6.66	5.63	5.24	0.446	0.270	0.406	0.370	0.512	
Image 6	6.77	10.11	10.59	7.85	11.48	0.376	0.357	0.460	0.355	0.586	
Image 7 Image 8	$6.56 \\ 6.44$	8.18 7.77	$9.07 \\ 8.37$	7.16 7.01	10.81 10.91	$0.402 \\ 0.437$	$0.355 \\ 0.332$	$0.470 \\ 0.478$	$0.349 \\ 0.373$	$0.628 \\ 0.652$	
Image 9	5.80	6.39	7.13	6.06	10.80	0.447	0.332	0.434	0.381	0.694	
Image 10	5.80	8.64	8.71	6.14	10.95	0.422	0.376	0.452	0.355	0.69	
Image 11	6.30	9.00	9.82	6.98	11.24	0.395	0.361	0.458	0.341	0.615	
Image 12	5.36	5.50	5.52	5.46	10.31	0.494	0.362	0.412	0.421	0.747	
Image 13 Image 14	$5.43 \\ 6.76$	$7.52 \\ 8.23$	$7.82 \\ 9.73$	5.78 7.77	11.02 10.80	$0.435 \\ 0.402$	$0.527 \\ 0.324$	$0.446 \\ 0.470$	$0.368 \\ 0.370$	0.711 0.593	
Image 14 Image 15	6.80	9.50	9.99	8.05	11.26	0.447	0.443	0.547	0.426	0.646	
Image 16	4.53	5.53	5.08	4.46	10.60	0.405	0.363	0.336	0.340	0.773	
Image 17	6.79	9.78	10.76	8.18	11.50	0.370	0.470	0.456	0.393	0.55	
mage 18	6.81	10.17	10.13	8.11	11.19	0.435	0.424	0.526	0.405	0.62	
mage 19	6.22	7.82	8.30	6.78	11.02 10.99	0.483	0.380	0.494	0.422	0.70	
Image 20 Image 21	$6.20 \\ 5.95$	$7.25 \\ 6.16$	$7.29 \\ 8.53$	6.65 6.48	10.99	$0.442 \\ 0.423$	0.417 0.409	$0.473 \\ 0.458$	$0.395 \\ 0.389$	0.672 0.683	
Image 22	5.66	8.18	7.99	5.92	10.87	0.420	0.369	0.435	0.360	0.708	
Image 23	6.41	9.28	9.63	7.07	11.03	0.431	0.331	0.456	0.371	0.636	
mage 24	6.52	9.81	9.26	7.63	11.34	0.447	0.466	0.540	0.421	0.659	
mage 25	6.91	9.28	11.51	8.49	11.13	0.316	0.412	0.417	0.333	0.50	
mage 26	6.29	8.98	8.74	6.96	11.04	0.457	0.512	0.500	0.411	0.693	
mage 27 mage 28	$6.68 \\ 6.79$	$8.70 \\ 9.48$	10.37 10.61	$7.87 \\ 8.05$	11.64 11.53	0.433 0.385	$0.374 \\ 0.372$	$0.508 \\ 0.470$	$0.399 \\ 0.356$	0.629 0.589	
mage 29	6.16	7.32	8.71	6.68	10.80	0.464	0.247	0.473	0.401	0.70	
mage 30	6.46	7.31	9.59	7.36	11.21	0.416	0.368	0.465	0.383	0.63	
mage 31	6.64	8.62	9.76	7.67	10.78	0.470	0.318	0.504	0.409	0.65	
mage 32	6.53	7.25	9.64	7.43	10.95	0.437	0.440	0.459	0.414	0.616	
mage 33 mage 34	$6.21 \\ 5.52$	$9.80 \\ 8.27$	$9.41 \\ 7.79$	7.16 5.87	11.42 10.88	$0.404 \\ 0.415$	0.447	$\frac{0.474}{0.425}$	0.384	$0.648 \\ 0.733$	
mage 35	6.29	9.12	9.39	7.02	11.12	0.415 0.444	$\frac{0.431}{0.467}$	0.425 0.477	0.371 0.386	0.75	
mage 36	6.02	8.68	8.71	6.79	11.15	0.417	0.400	0.471	0.376	0.67	
mage 37	6.71	9.61	10.26	7.89	11.03	0.404	0.462	0.457	0.399	0.580	
mage 38	6.63	9.53	10.47	7.59	11.23	0.379	0.437	0.441	0.363	0.563	
mage 39	6.76	7.79	10.14	7.90	11.20	0.406	0.489	0.452	0.414	0.583	
mage 40 mage 41	6.33 6.71	$7.53 \\ 10.45$	8.40 10.09	6.73 8.01	$10.36 \\ 11.64$	$0.426 \\ 0.421$	$0.296 \\ 0.438$	$0.452 \\ 0.510$	$0.359 \\ 0.385$	$0.651 \\ 0.623$	
mage 41	6.68	9.25	9.95	7.75	10.87	0.421	0.459	0.450	0.383	0.586	
mage 43	5.75	8.86	8.41	6.07	10.85	0.396	0.411	0.431	0.355	0.690	
mage 44	6.58	9.56	9.46	7.29	10.81	0.485	0.478	0.558	0.438	0.677	
mage 45	5.92	4.91	7.11	6.26	10.88	0.459	0.243	0.456	0.388	0.672	
mage 46	5.01	5.67	5.18	5.04	10.88	0.476	0.384	0.390	0.400	0.76	
mage 47 mage 48	6.58 6.06	9.31 8.19	$9.48 \\ 8.60$	7.69 6.54	11.04 10.88	$0.462 \\ 0.461$	0.415 0.403	$0.524 \\ 0.498$	$0.425 \\ 0.401$	0.658 0.703	
mage 49	6.89	10.21	11.35	8.30	11.83	0.353	0.390	0.480	0.347	0.568	
mage 50	5.53	8.08	7.76	5.68	10.83	0.426	0.500	0.449	0.365	0.72	
mage 51	6.28	8.09	8.34	7.18	10.88	0.462	0.413	0.500	0.418	0.684	
mage 52	6.55	10.38	9.99	7.68	11.77	0.378	0.453	0.472	0.353	0.614	
mage 53	$6.02 \\ 6.32$	$\frac{7.47}{9.49}$	7.32	$6.60 \\ 7.34$	10.78 11.72	0.449 0.392	0.409	0.483	0.398	0.67	
mage 54 mage 55	6.26	9.49	$9.57 \\ 9.79$	6.88	11.72 11.32	0.392	0.381 0.380	$0.473 \\ 0.465$	0.364 0.343	0.636 0.635	
mage 56	6.50	8.23	9.33	7.36	10.85	0.438	0.394	0.506	0.398	0.652	
mage 57	6.26	6.96	7.93	6.73	10.82	0.459	0.323	0.484	0.405	0.652	
mage 58	6.18	8.56	8.98	6.75	11.26	0.413	0.407	0.466	0.371	0.672	
mage 59	5.67	8.61	8.14	5.97	11.12	0.396	0.536	0.423	0.344	0.693	
mage 60 mage 61	$6.62 \\ 6.11$	$8.61 \\ 7.21$	$9.29 \\ 8.05$	$7.62 \\ 6.45$	11.24 10.54	$0.430 \\ 0.440$	$0.369 \\ 0.281$	$0.502 \\ 0.453$	$0.400 \\ 0.370$	0.65	
mage 62	6.55	9.35	9.78	7.38	10.93	0.410	0.352	0.433	0.370	0.59	
mage 63	6.67	8.94	10.28	8.04	11.56	0.416	0.458	0.499	0.411	0.61	
mage 64	6.42	8.80	8.80	7.32	11.35	0.435	0.398	0.486	0.397	0.658	
mage 65	6.47	7.94	8.72	7.02	10.76	0.407	0.326	0.461	0.352	0.59	
mage 66 mage 67	6.00	$\frac{7.62}{7.76}$	7.34	6.39	10.81	0.450	0.398	0.464	0.391	0.70	
nage 67	5.47 6.70	7.76 8.89	$7.88 \\ 9.73$	5.79 7.79	10.87 10.69	0.427 0.457	0.357 0.412	$0.430 \\ 0.530$	$0.378 \\ 0.421$	$0.73 \\ 0.64$	
mage 69	6.64	9.36	10.14	7.60	11.46	0.399	0.412	0.330 0.472	0.369	0.61	
mage 70	6.73	9.84	10.33	8.07	11.29	0.389	0.444	0.456	0.387	0.58	
mage 71	5.96	8.59	8.63	6.63	11.26	0.413	0.402	0.459	0.372	0.67	
mage 72	6.38	6.76	7.95	6.72	10.38	0.416	0.304	0.429	0.356	0.63	
mage 73	5.93	6.03	7.07	6.11	10.68	0.521	0.447	0.501	0.446	0.74	
mage 74 mage 75	$6.72 \\ 6.77$	$4.46 \\ 8.39$	$10.17 \\ 10.59$	7.90 7.96	11.64 11.39	0.386 0.387	$0.298 \\ 0.477$	$0.468 \\ 0.477$	$0.351 \\ 0.402$	$0.59 \\ 0.58$	
mage 76	6.67	8.22	9.35	7.56	11.11	0.357	0.394	0.477 0.524	0.402	0.64	
mage 77	6.73	9.99	10.45	8.04	11.33	0.375	0.423	0.441	0.363	0.57	
mage 78	6.34	4.41	8.63	7.01	10.74	0.482	0.274	0.501	0.421	0.68	
mage 79	6.09	9.06	8.93	6.84	11.19	0.407	0.350	0.442	0.367	0.63	
mage 80	6.49	8.85	9.04	7.26	10.95	0.439	0.335	0.473	0.397	0.65	
mage 81 mage 82	6.77 5.30	$10.36 \\ 5.69$	$9.72 \\ 5.69$	7.83 5.31	11.61 10.34	$0.356 \\ 0.458$	0.384 0.332	$0.468 \\ 0.416$	0.323 0.382	$0.57 \\ 0.72$	
mage 82 mage 83	6.44	8.70	9.37	7.35	10.34	0.458 0.436	0.332 0.476	0.416 0.464	0.382 0.411	0.72	
mage 84	6.01	9.01	8.51	6.48	10.80	0.438	0.470	0.404 0.500	0.403	0.69	
mage 85	5.84	6.37	6.97	6.35	10.59	0.463	0.281	0.448	0.393	0.680	
mage 86	5.48	5.72	6.30	5.79	10.57	0.483	0.255	0.426	0.410	0.740	
mage 87 mage 88	5.67	7.01	6.87	5.88	10.70		0.347	0.430	0.395	0.729	
	6.73	9.13	10.34	7.93	11.30	0.391	0.367	0.467	0.361	0.58	

			Tab	le A.34 con	tinued from	previous p	age			
_			SNR (dB)		I			SSIM		
Image 89	6.07	8.07	8.84	6.54	10.89	0.448	0.333	0.463	0.379	0.685
Image 90	6.72	8.65	10.03	7.87	11.30	0.422	0.444	0.501	0.415	0.625
Image 91	6.32	7.47	8.12	6.96	10.57	0.427	0.322	0.462	0.377	0.641
Image 92	6.66	9.25	9.79	7.85	11.27	0.426	0.400	0.524	0.404	0.641
Image 93	6.00	6.83	7.36	6.24	10.35	0.471	0.355	0.461	0.397	0.679
Image 94	6.50	8.02	9.45	7.44	11.13	0.444	0.315	0.501	0.400	0.658
Image 95	6.71	8.80	10.11	7.93	11.37	0.441	0.415	0.537	0.417	0.648
Image 96	6.00	8.03	8.11	6.34	10.95	0.456	0.368	0.452	0.393	0.712
Image 97	5.76	6.76	8.29	6.24	10.90	0.418	0.426	0.457	0.380	0.698
Image 98	6.84	10.41	10.67	8.28	11.62	0.406	0.391	0.498	0.378	0.614
Image 99	6.53	8.16	9.86	7.37	11.16	0.432	0.389	0.479	0.387	0.628
Image 100	4.67	5.53	4.66	4.55	10.54	0.436	0.363	0.341	0.372	0.783
Image 101	6.42	8.43	9.51	7.20	10.92	0.426	0.442	0.479	0.389	0.629
Image 102	6.47	8.75	9.64	7.36	11.20	0.423	0.479	0.465	0.388	0.632
Image 103	6.94	11.92	11.75	7.94	11.76	0.288	0.305	0.378	0.262	0.466
Image 104	6.52	8.99	9.59	7.40	10.74	0.446	0.486	0.469	0.433	0.627
Image 105	5.38	8.75	7.71	5.65	11.21	0.406	0.586	0.426	0.357	0.740
Image 106	6.49	9.11	8.66	7.23	10.98	0.434	0.446	0.497	0.390	0.643
Image 107	6.21	7.97	8.74	6.74	10.52	0.471	0.324	0.463	0.405	0.678
Image 108	6.64	8.96	9.59	7.50	11.26	0.453	0.388	0.525	0.409	0.664
Image 109	6.36	7.44	8.02	6.96	10.84	0.452	0.380	0.503	0.402	0.670
Image 110	6.98	7.60	11.70	8.38	11.77	0.313	0.315	0.420	0.313	0.503
Average	6.28	8.29	8.94	7.03	10.99	0.426	0.391	0.467	0.384	0.649

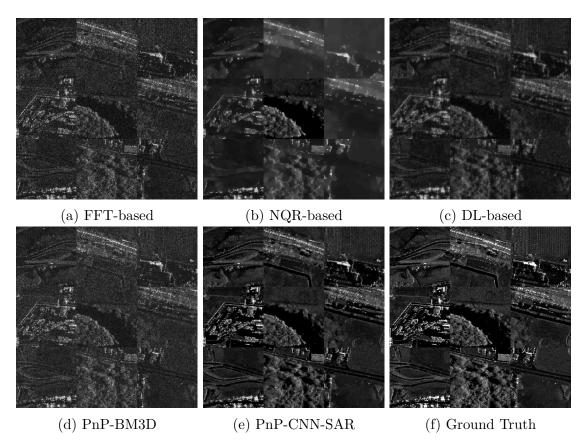


Figure A.34 Reconstruction results for Images 19, 40, 45, 46, 59, 61, 66, 96, and 109 from the Kapıkule test set for the case of 70% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$.

A.3.11 Experiment 11: 50% Data Availability and $\sigma_n = \sigma_y$

Table A.35 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR for 95 images, and best SSIM for 108 images. Fig. A.35 shows the reconstruction results for a selection of images from the Kapıkule test set from all methods for this experiment. PnP-CNN-SAR produce better reconstructions compared to other methods.

Table A.35 SNR and SSIM values for images from the KapıkuleTest Set for the case of 50% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			SNR (dB)		SSIM							
	FFT-based N	IQR-based	DL-based Pi	nP-BM3D F	nP-CNN-SAR	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAI		
Image 1	6.07	7.88	8.30	7.87	9.00	0.285	0.344	0.326	0.344	0.47		
Image 2 Image 3	5.73 6.30	8.23 4.38	$7.65 \\ 8.80$	$7.08 \\ 8.69$	$8.96 \\ 9.11$	0.256 0.241	0.387 0.391	$0.380 \\ 0.404$	0.335 0.391	$0.51 \\ 0.45$		
mage 4	6.12	8.17	7.58	7.44	8.62	0.274	0.398	0.417	0.366	0.51		
mage 5	5.00	5.83	5.80	5.32	8.15	0.335	0.256	0.326	0.320	0.61		
mage 6 mage 7	6.29 6.01	9.60 8.14	9.04 7.81	8.77 7.39	9.38 8.73	0.246 0.250	0.344 0.325	$0.358 \\ 0.392$	0.323 0.326	$0.43 \\ 0.49$		
mage 8	5.85	7.23	7.18	6.92	8.14	0.274	0.310	0.395	0.350	0.52		
mage 9	5.33	4.68	6.14	6.02	8.11	0.324	0.216	0.339	0.334	0.54		
mage 10 mage 11	5.53 6.00	$7.77 \\ 8.92$	$7.47 \\ 8.46$	$6.49 \\ 7.82$	$8.51 \\ 9.45$	$0.288 \\ 0.253$	0.293 0.326	$0.365 \\ 0.365$	0.333 0.320	$0.54 \\ 0.49$		
mage 12	4.85	4.20	4.78	4.75	7.68	0.366	0.250	0.321	0.336	0.62		
mage 13	5.21	6.19	6.80	5.97	8.41	0.300	0.363	0.359	0.326	0.58		
mage 14 mage 15	6.29 6.23	$9.07 \\ 9.29$	8.40 8.65	8.28 8.63	$8.95 \\ 9.07$	0.252 0.241	0.345 0.439	$0.385 \\ 0.471$	0.328 0.406	$0.46 \\ 0.51$		
mage 16	4.25	4.57	4.43	3.87	7.73	0.289	0.415	0.281	0.283	0.67		
mage 17	6.42	8.71	8.98	9.14	9.39	0.234	0.414	0.348	0.375	0.40		
mage 18	6.41	9.76	8.66	8.81	8.91 8.43	0.246	0.411	0.442	0.390	$0.48 \\ 0.54$		
mage 19 mage 20	5.79 5.59	6.83 6.30	$7.13 \\ 6.38$	6.85 6.33	8.17	0.324 0.288	0.284 0.342	$0.398 \\ 0.410$	0.365 0.366	0.54		
mage 21	5.61	8.37	7.47	6.71	8.99	0.278	0.435	0.378	0.352	0.54		
mage 22	5.34	8.01	6.98	6.16	8.77	0.287	0.404	0.353	0.317	0.58		
mage 23 mage 24	6.03 6.09	8.25 8.05	8.31 8.00	$7.72 \\ 8.12$	9.10 9.11	0.304 0.244	0.277 0.408	$0.343 \\ 0.469$	0.317 0.411	$0.48 \\ 0.52$		
mage 25	6.54	9.55	9.89	9.89	9.62	0.178	0.376	0.326	0.335	0.32		
mage 26	5.84	7.60	7.55	7.17	8.71	0.303	0.381	0.419	0.367	0.55		
mage 27	6.27	5.08	8.88	8.63	9.34	0.253 0.229	0.381	0.420	0.375 0.331	0.48		
mage 28 mage 29	$6.45 \\ 5.77$	10.25 6.98	9.01 7.50	8.95 7.01	9.46 8.60	0.229	0.367 0.214	$0.386 \\ 0.356$	0.342	$0.44 \\ 0.54$		
mage 30	6.09	5.97	8.33	7.93	9.00	0.271	0.356	0.378	0.352	0.49		
mage 31	6.27	7.73	8.33	8.23	8.79	0.297	0.306	0.391	0.353	0.49		
mage 32 mage 33	6.11 5.88	8.05 8.39	8.38 8.01	7.90 7.58	$9.00 \\ 8.95$	$0.302 \\ 0.240$	$0.418 \\ 0.413$	$0.355 \\ 0.375$	0.359 0.349	$0.47 \\ 0.48$		
mage 34	5.27	7.45	6.84	6.00	8.68	0.288	0.374	0.350	0.321	0.59		
nage 35	6.01	8.36	8.15	7.59	9.03	0.310	0.342	0.369	0.336	0.5		
nage 36	5.70	8.28	7.55	7.18	9.06	0.270	0.375	0.390	0.344	0.54		
nage 37 nage 38	6.27 6.29	8.21 9.03	$\frac{8.64}{8.93}$	8.59 8.63	$8.96 \\ 9.39$	0.264 0.239	$0.397 \\ 0.365$	$0.340 \\ 0.342$	$0.358 \\ 0.350$	0.4: 0.4:		
nage 39	6.31	5.67	8.81	8.68	9.09	0.260	0.418	0.354	0.400	0.43		
nage 40	5.76	7.16	7.32	6.77	8.42	0.280	0.268	0.366	0.328	0.53		
mage 41 mage 42	6.29 6.33	$\frac{9.27}{8.54}$	8.61 8.66	$8.74 \\ 8.42$	9.35 8.98	0.231	0.421 0.391	0.439 0.348	0.379	$0.49 \\ 0.43$		
mage 42	5.45	8.39	7.33	6.41	8.97	0.273 0.272	0.391	0.356	0.341 0.319	0.56		
mage 44	6.13	8.17	8.07	7.68	8.76	0.276	0.461	0.477	0.423	0.56		
mage 45	5.37	6.25	6.32	6.02	8.30	0.319	0.270	0.369	0.343	0.57		
mage 46 mage 47	$\frac{4.58}{6.12}$	3.99 8.88	4.48 8.11	$\frac{4.48}{7.97}$	7.72 8.87	0.351 0.288	$0.320 \\ 0.400$	0.303 0.434	0.336 0.378	0.66		
mage 48	5.71	7.53	7.37	6.87	8.53	0.306	0.385	0.408	0.365	0.5		
mage 49	6.49	10.47	9.77	9.65	9.63	0.200	0.405	0.406	0.348	0.43		
mage 50	5.13	7.14	6.71	5.72	8.50	0.288	0.392	0.385	0.332	0.60		
nage 51 nage 52	5.87 6.13	7.28 6.58	$7.12 \\ 8.67$	$7.06 \\ 8.49$	$8.28 \\ 9.47$	0.303 0.218	0.376 0.385	$0.411 \\ 0.401$	0.377 0.338	0.53 0.4'		
nage 53	5.49	6.87	6.35	6.26	8.44	0.287	0.378	0.406	0.364	0.56		
nage 54	5.98	8.48	8.21	7.92	9.26	0.244	0.323	0.387	0.334	0.48		
mage 55 mage 56	5.88 6.11	$\frac{8.61}{7.54}$	$\frac{8.48}{7.93}$	7.53 7.66	$9.34 \\ 8.80$	$0.270 \\ 0.282$	$0.385 \\ 0.334$	$0.368 \\ 0.415$	0.308	0.50 0.52		
nage 57	5.78	6.84	6.94	6.85	8.78	0.282	0.305	0.413	0.361 0.373	0.53		
nage 58	5.70	7.84	7.77	7.28	9.26	0.272	0.453	0.381	0.339	0.54		
mage 59	5.27	8.27	7.10	6.19	8.72	0.254	0.392	0.342	0.304	0.5		
nage 60 nage 61	$6.14 \\ 5.56$	7.91 6.60	$7.92 \\ 7.03$	7.89 6.41	$8.85 \\ 8.24$	$0.270 \\ 0.293$	0.385 0.248	$0.426 \\ 0.364$	0.375 0.315	0.5 0.5		
nage 62	6.16	9.09	8.54	8.03	9.08	0.284	0.293	0.334	0.316	0.4		
nage 63	6.30	8.29	8.79	8.79	9.32	0.244	0.432	0.410	0.398	0.4		
nage 64 nage 65	5.99 5.92	7.87 7.40	7.66	7.50	8.81 8.86	0.284	0.335 0.308	0.407	0.361 0.332	0.5 0.4		
nage 65	5.92 5.43	6.30	$\frac{7.43}{6.28}$	7.38 6.13	8.86 8.33	$0.250 \\ 0.302$	0.308 0.304	$0.381 \\ 0.393$	0.332 0.358	0.4		
nage 67	5.18	7.50	6.78	5.97	8.62	0.308	0.344	0.332	0.326	0.6		
nage 68	6.28	7.93	8.22	8.27	8.74	0.269	0.404	0.444	0.397	0.5		
nage 69 nage 70	6.25 6.38	9.76 7.75	$8.74 \\ 8.77$	8.33 8.75	$\frac{9.35}{9.10}$	$0.265 \\ 0.242$	0.368 0.345	$0.385 \\ 0.348$	0.332 0.343	0.4° 0.4°		
nage 70	5.64	8.25	7.60	6.87	9.12	0.242	0.345	0.348	0.332	0.5		
nage 72	5.82	4.89	6.95	6.57	8.09	0.269	0.269	0.361	0.317	0.5		
nage 73	5.39	5.44	6.07	5.70	7.89	0.339	0.284	0.409	0.382	0.6		
nage 74	6.37 6.39	8.52 8.33	$8.77 \\ 8.95$	$8.76 \\ 8.88$	$9.34 \\ 9.17$	0.239 0.231	0.354 0.425	$0.384 \\ 0.387$	0.325 0.398	$0.4 \\ 0.4$		
nage 76	6.18	8.34	8.12	8.10	8.85	0.259	0.382	0.453	0.392	0.5		
nage 77	6.49	7.42	9.03	8.90	9.13	0.250	0.356	0.347	0.330	0.4		
nage 78	5.94 5.73	7.14	$\frac{7.52}{7.70}$	7.14	8.46	0.311	0.290	0.412	0.363	0.5		
nage 79 nage 80	5.73 5.95	7.82 8.27	$7.70 \\ 7.84$	$7.20 \\ 7.64$	8.98 8.77	$0.262 \\ 0.298$	0.319 0.303	$0.349 \\ 0.366$	0.306 0.343	0.4		
nage 81	6.29	9.54	8.28	8.32	9.09	0.193	0.372	0.408	0.328	0.4		
nage 82	4.76	5.58	5.01	4.91	7.85	0.329	0.348	0.349	0.340	0.6		
nage 83	6.06	8.81	8.19	7.72	8.70	0.277	0.403	0.381	0.381	0.4		
nage 84 nage 85	5.55 5.39	7.96 6.02	$7.24 \\ 6.03$	$6.67 \\ 5.92$	8.60 7.96	0.274 0.306	$0.434 \\ 0.314$	$0.433 \\ 0.362$	0.386 0.339	0.5 0.5		
mage 86	5.03	5.04	5.44	5.25	7.91	0.357	0.220	0.332	0.339	0.60		
nage 87	5.20	5.59	5.99	5.57	8.26	0.331	0.282	0.338	0.332	0.58		
nage 88	6.39	8.39	8.86	8.88	9.31	0.243	0.358	0.365	0.323	0.4		

			Tab	le A.35 con	tinued from	previous p	age			
			SNR (dB)					SSIM		
Image 89	5.75	7.29	7.69	6.97	8.89	0.322	0.331	0.362	0.323	0.550
Image 90	6.21	8.05	8.52	8.57	8.91	0.249	0.438	0.420	0.412	0.484
Image 91	5.89	7.11	6.97	6.74	8.30	0.272	0.311	0.383	0.343	0.522
Image 92	6.25	8.65	8.37	8.30	9.09	0.238	0.379	0.450	0.390	0.512
Image 93	5.42	5.91	6.39	6.03	7.90	0.323	0.251	0.370	0.339	0.550
Image 94	6.08	7.92	7.98	7.76	8.96	0.279	0.299	0.405	0.356	0.512
Image 95	6.30	9.45	8.70	8.63	9.10	0.240	0.426	0.462	0.399	0.522
Image 96	5.59 5.44 6.42 6.15	7.31	7.03	$6.51 \\ 6.37$	8.73	0.333 0.265 0.235 0.289	0.306 0.473 0.380 0.341	0.348 0.387 0.419 0.366	0.340 0.351 0.353 0.334	0.561
Image 97		7.97 7.38 5.05	7.27		8.95					0.578
Image 98			9.11	9.23	9.30					0.472
Image 99			8.42	8.02	9.12					0.490
Image 100	4.16	4.20	4.05	3.78	7.81	0.312	0.294	0.265	0.283	0.679
Image 101	5.99	7.94	8.20	7.74	8.93	0.266	0.391	0.385	0.355	0.499
Image 102	6.04	8.96	8.38	7.89	8.94	0.268	0.377	0.384	0.362	0.480
Image 103	6.58	9.45	10.02	9.52	10.04	0.175	0.268	0.288	0.242	0.303
Image 104	6.09	6.91	8.14	7.82	8.75	0.309	0.393	0.344	0.374	0.479
Image 105	5.16	7.69	6.69	5.85	8.55	0.289	0.440	0.347	0.318	0.609
Image 106	5.95	7.40	7.52	7.46	8.70	0.262	0.355	0.427	0.370	0.532
Image 107	5.92	7.77	7.53	7.20	8.55	0.339	0.296	0.345	0.348	0.524
Image 108	6.19	8.97	8.23	7.99	9.13	0.262	0.417	0.449	0.389	0.535
Image 109	5.78	6.37	7.00	6.93	8.57	0.281	0.344	0.435	0.382	0.561
Image 110	6.58	10.05	9.92	9.83	9.93	0.184	0.335	0.332	0.290	0.351
Average	5.86	7.59	7.70	7.39	8.81	0.275	0.355	0.380	0.348	0.516

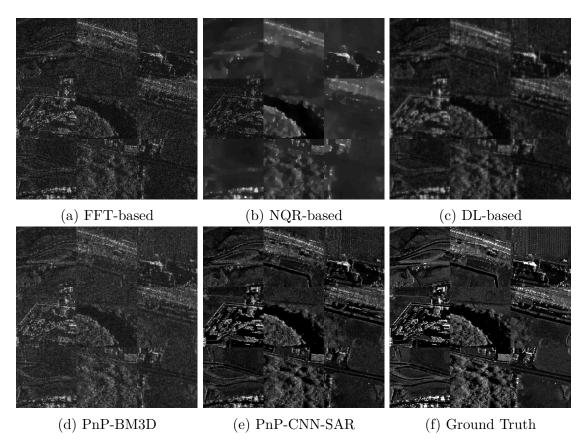


Figure A.35 Reconstruction results for Images 19, 40, 45, 46, 59, 61, 66, 96, and 109 from the Kapıkule test set for the case of 50% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$.

A.3.12 Experiment 12: 30% Data Availability and $\sigma_n = \sigma_y$

Table A.36 shows the SNR and SSIM values for this experiment. PnP-CNN-SAR gives the best SNR for 13 images, and best SSIM for 59 images. Fig. A.36 shows the reconstruction results for a selection of images from the Kapıkule test set from all methods for this experiment. None of the methods are able to produce acceptable reconstructions for most of the images.

Table A.36 SNR and SSIM values for images from the KapıkuleTest Set for the case of 30% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$. Best results are shown in bold, second best results are shown in red.

			SNR (dB)					SSIM		
	FFT-based NQ	R-based D	L-based Pn	P-BM3D P	nP-CNN-SAR	FFT-based	NQR-based	DL-based	PnP-BM3D	PnP-CNN-SAR
Image 1	5.07	4.89	6.33	7.61	6.56	0.197	0.304	0.231	0.293	0.279
Image 2 Image 3	4.90 5.30	$7.34 \\ 8.80$	$\frac{5.87}{6.70}$	$6.94 \\ 8.77$	5.87 7.08	$0.170 \\ 0.147$	$0.363 \\ 0.397$	$0.306 \\ 0.317$	0.292 0.385	0.302 0.336
Image 4	5.03	6.91	5.91	7.19	6.52	0.167	0.338	0.350	0.348	0.375
Image 5	4.25	5.00	4.61	4.92	5.41	0.234	0.207	0.243	0.254	0.369
Image 6	5.39	9.44	6.73	8.46	7.12	0.161	0.323	0.270	0.303	0.273
Image 7 Image 8	5.10 4.89	$\begin{array}{c} \textbf{7.82} \\ \textbf{6.67} \end{array}$	$5.97 \\ 5.52$	$\begin{matrix} 7.11 \\ 6.72 \end{matrix}$	6.50 6.26	0.160 0.188	0.312 0.288	0.319 0.324	0.324 0.333	$0.349 \\ 0.366$
Image 9	4.55	5.50	4.93	5.57	6.00	0.224	0.217	0.253	0.264	0.356
Image 10	4.73	7.46	5.70	6.51	6.25	0.201	0.306	0.289	0.276	0.313
Image 11	5.11	5.61	6.32	7.76	7.19	0.174	0.278	0.281	0.268	0.309
Image 12	4.12	4.37	3.95	4.41	4.94	0.250	0.276	0.244	0.270	0.375
Image 13 Image 14	$4.52 \\ 5.21$	$\begin{array}{c} 7.12 \\ 8.26 \end{array}$	$\frac{5.28}{6.42}$	$\frac{5.81}{7.80}$	5.70 6.99	0.209 0.159	0.390 0.279	0.289 0.305	0.264 0.301	$0.319 \\ 0.315$
Image 15	5.33	9.17	6.44	8.43	6.82	0.142	0.423	0.388	0.399	0.409
Image 16	3.63	4.00	3.64	3.67	4.37	0.199	0.261	0.236	0.222	0.437
Image 17	5.32	6.83	6.80	8.79	7.24	0.144	0.338	0.273	0.377	0.268
Image 18	5.26	8.86	6.62	8.46	6.87	0.150	0.365	0.369	0.376	0.379
Image 19 Image 20	$4.85 \\ 4.58$	$\begin{array}{c} 7.07 \\ 6.18 \end{array}$	5.57 5.00	$\frac{6.58}{5.80}$	6.33 5.85	0.218 0.191	0.329 0.369	$0.305 \\ 0.334$	0.307 0.339	$0.350 \\ 0.390$
Image 20	4.86	7.83	5.69	6.60	6.77	0.191	0.368	0.334	0.310	0.356
Image 22	4.58	6.91	5.43	6.07	6.31	0.202	0.306	0.276	0.260	0.369
Image 23	5.04	8.39	6.39	7.47	6.88	0.200	0.248	0.236	0.240	0.296
Image 24	5.25	8.45	6.09	7.77	6.70	0.153	0.427	0.391	0.404	0.412
Image 25	5.50	7.41	7.21	9.76	7.30	0.108	0.359	0.254	0.364	0.232
Image 26 Image 27	$5.01 \\ 5.22$	7.54 2.35	5.86 6.71	$\frac{6.93}{8.48}$	6.45 7.13	0.202 0.160	$0.372 \\ 0.179$	0.332 0.339	0.335 0.355	0.375 0.337
Image 27	5.32	7.11	6.69	8.73	7.13	0.146	0.331	0.302	0.325	0.306
Image 29	4.92	7.15	5.80	6.71	6.41	0.234	0.277	0.258	0.248	0.326
Image 30	5.09	6.69	6.31	7.68	6.97	0.189	0.324	0.287	0.308	0.316
Image 31	5.23	8.20	6.35	7.82	6.99	0.196	0.294	0.294	0.282	0.341
Image 32	5.08	2.09	6.30	7.84	6.87	0.188	0.303	0.253	0.320	0.302
Image 33 Image 34	$4.93 \\ 4.42$	4.91 6.80	$6.15 \\ 5.37$	7.56 5.86	6.53 4.51	0.163 0.198	0.339 0.307	$0.306 \\ 0.283$	$0.329 \\ 0.263$	0.298 0.314
Image 35	5.14	7.81	6.10	7.43	6.93	0.219	0.295	0.259	0.275	0.314
Image 36	4.84	6.98	5.79	7.03	6.34	0.186	0.348	0.308	0.312	0.335
Image 37	5.29	4.81	6.51	8.40	7.08	0.165	0.339	0.244	0.333	0.282
Image 38	5.28	3.92	6.55	8.41	7.29	0.165	0.176	0.254	0.330	0.279
Image 39	5.28	7.75	6.55	$8.45 \\ 6.46$	7.23	0.173	0.404	0.263	0.380	0.280
Image 40 Image 41	$4.83 \\ 5.35$	$5.55 \\ 8.18$	$\frac{5.63}{6.38}$	8.44	6.14 7.13	$0.196 \\ 0.131$	0.223 0.356	$0.276 \\ 0.362$	$0.281 \\ 0.374$	0.337 0.380
Image 42	5.26	6.92	6.44	8.02	7.00	0.173	0.316	0.249	0.297	0.282
Image 43	4.72	5.59	5.74	6.36	5.61	0.182	0.320	0.288	0.270	0.307
Image 44	5.04	7.48	6.05	7.57	6.47	0.173	0.430	0.394	0.400	0.430
Image 45	4.59	5.97	4.92	5.55	$\frac{5.61}{4.64}$	0.216	0.260	$0.283 \\ 0.233$	0.275	0.361
Image 46 Image 47	3.76 5.19	3.83 8.15	$\frac{3.58}{6.06}$	$\frac{4.01}{7.71}$	6.87	0.248 0.174	0.220 0.380	0.233	$0.265 \\ 0.349$	0.373 0.376
Image 48	4.88	7.17	5.63	6.53	6.49	0.207	0.352	0.316	0.305	0.388
Image 49	5.45	11.04	7.13	9.45	7.51	0.119	0.390	0.337	0.372	0.329
Image 50	4.56	6.86	5.20	5.77	5.60	0.199	0.368	0.318	0.299	0.341
Image 51	4.88	6.68	5.64	6.74	6.14	0.195	0.368	0.330	0.335	0.361
Image 52 Image 53	$5.21 \\ 4.57$	9.27 5.70	6.58 5.00	$\frac{8.26}{5.78}$	7.65 5.61	0.140 0.197	0.377 0.337	0.331 0.325	$0.326 \\ 0.336$	$0.369 \\ 0.372$
Image 54	5.06	8.32	6.20	7.68	7.15	0.166	0.309	0.315	0.310	0.345
Image 55	5.07	8.46	6.51	7.49	6.91	0.183	0.259	0.284	0.268	0.307
Image 56	5.18	7.49	6.04	7.42	6.65	0.184	0.314	0.328	0.336	0.365
Image 57	4.95 4.88	6.31	5.37 6.00	6.49	6.55 6.74	0.202 0.181	0.286 0.374	0.315	0.315	0.378
Image 58 Image 59	4.62	$8.14 \\ 7.00$	5.60	$7.01 \\ 6.23$	3.98	0.181	0.374 0.335	$0.299 \\ 0.278$	$0.290 \\ 0.252$	$0.341 \\ 0.267$
Image 60	5.21	7.95	6.14	7.57	6.76	0.170	0.348	0.350	0.345	0.376
Image 61	4.56	6.67	5.45	6.18	5.89	0.211	0.283	0.278	0.280	0.351
Image 62	5.14	6.52	6.40	7.83	6.98	0.185	0.257	0.229	0.253	0.291
Image 63	5.25	7.36	6.70	8.40	7.19	0.156	0.401	0.335	0.386	0.330
Image 64 Image 65	5.04 4.91	$7.63 \\ 7.50$	5.85 5.69	$7.08 \\ 6.85$	6.97 6.56	0.178 0.170	0.358 0.339	0.324 0.308	0.324 0.314	$0.375 \\ 0.346$
Image 66	4.69	5.05	4.93	5.91	5.85	0.170	0.339 0.257	0.308	0.314 0.334	0.346
Image 67	4.48	7.18	5.32	5.90	5.14	0.218	0.323	0.250	0.246	0.279
Image 68	5.11	7.94	6.38	7.96	6.50	0.164	0.374	0.363	0.367	0.398
Image 69	5.19	6.06	6.50	8.10	7.07	0.174	0.299	0.294	0.315	0.308
Image 70	5.30	9.18	6.60	8.57 6.75	7.10	0.153	0.339	0.261	0.320	0.272
Image 71 Image 72	$4.74 \\ 4.90$	$7.61 \\ 6.57$	5.73 5.31	$6.75 \\ 6.25$	6.73 6.09	$0.177 \\ 0.177$	$0.336 \\ 0.276$	0.311 0.284	0.298 0.300	$0.364 \\ 0.344$
Image 72 Image 73	4.64	5.38	4.77	5.30	5.55	0.243	0.270	0.234	0.307	0.407
Image 74	5.21	9.38	6.61	8.30	7.70	0.151	0.331	0.308	0.310	0.331
Image 75	5.31	0.97	6.73	8.67	7.25	0.142	0.336	0.313	0.403	0.325
Image 76	5.21	8.50	6.16	7.85	6.80	0.159	0.408	0.373	0.377	0.391
Image 77	5.37	9.13	6.81 5.74	8.73 6.95	7.13 6.54	0.143	0.333	0.263	0.320	0.267 0.357
Image 78 Image 79	$4.93 \\ 4.81$	$7.18 \\ 7.95$	$5.74 \\ 5.98$	$6.95 \\ 7.10$	6.54 6.29	$0.207 \\ 0.171$	$0.336 \\ 0.332$	$0.309 \\ 0.275$	$0.310 \\ 0.262$	0.357
Image 19 Image 80	4.98	7.81	6.03	7.10	6.82	0.171	0.332	0.273	0.264	0.312
Image 81	5.38	7.46	6.31	7.89	6.76	0.119	0.356	0.355	0.356	0.334
Image 82	4.01	5.12	3.97	4.49	5.30	0.218	0.323	0.270	0.276	0.423
Image 83	5.07	7.42	6.26	7.51	6.69	0.179	0.378	0.287	0.347	0.324
Image 84	4.75	7.39	5.68	6.55	6.38	0.178	0.410	0.366	0.361	0.401
Image 85 Image 86	$4.49 \\ 4.21$	4.16 5.23	$4.70 \\ 4.26$	5.45 4.85	5.75 5.33	0.212 0.247	0.215 0.269	0.276	0.279	$0.349 \\ 0.363$
image ou		$\frac{5.23}{5.16}$	$\frac{4.26}{4.72}$	$\frac{4.85}{5.33}$	5.64	0.247	0.269	0.233 0.252	$0.251 \\ 0.258$	0.346
Image 87	4.32									

			SNR (dB)		1		SSIM				
Image 89	4.92	6.55	5.90	6.85	6.42	0.224	0.287	0.262	0.262	0.310	
Image 90	5.31	8.27	6.41	8.30	6.97	0.160	0.429	0.344	0.403	0.347	
Image 91	4.89	6.44	5.44	6.38	6.18	0.189	0.279	0.309	0.307	0.360	
Image 92	5.21	8.28	6.32	8.08	7.06	0.157	0.375	0.385	0.386	0.382	
Image 93	4.52	6.19	5.01	5.66	5.76	0.226	0.307	0.283	0.288	0.363	
Image 94	5.13	8.03	6.15	7.49	6.86	0.195	0.336	0.317	0.308	0.353	
Image 95	5.28	6.83	6.52	8.35	6.89	0.149	0.409	0.380	0.384	0.404	
Image 96	4.75	5.31	5.52	6.28	6.34	0.229	0.298	0.249	0.254	0.332	
Image 97	4.70	7.15	5.59	6.32	6.09	0.175	0.420	0.316	0.309	0.364	
Image 98	5.45	5.31	6.86	8.93	7.38	0.144	0.342	0.341	0.341	0.345	
Image 99	5.12	4.93	6.39	7.96	7.01	0.188	0.286	0.275	0.296	0.309	
Image 100	3.54	3.51	3.36	3.54	3.82	0.210	0.317	0.205	0.215	0.345	
Image 101	5.06	7.45	6.26	7.61	6.92	0.182	0.353	0.305	0.335	0.333	
Image 102	5.13	8.04	6.30	7.76	7.12	0.172	0.366	0.295	0.332	0.321	
Image 103	5.52	8.87	7.27	9.13	7.45	0.102	0.223	0.201	0.231	0.193	
Image 104	5.11	6.08	6.19	7.63	6.78	0.202	0.349	0.245	0.320	0.292	
Image 105	4.48	6.84	5.27	5.75	5.31	0.200	0.327	0.276	0.261	0.301	
Image 106	4.98	6.72	5.82	7.08	6.76	0.171	0.367	0.348	0.359	0.388	
Image 107	4.96	7.23	5.80	6.84	6.54	0.230	0.250	0.235	0.250	0.314	
Image 108	5.20	8.72	6.40	7.87	6.78	0.166	0.402	0.378	0.373	0.396	
Image 109	4.77	5.00	5.32	6.59	6.24	0.185	0.314	0.353	0.361	0.404	
Image 110	5.49	7.94	7.16	9.59	7.54	0.108	0.324	0.254	0.322	0.240	
Average	4.94	6.84	5.89	7.14	6.48	0.182	0.323	0.298	0.312	0.339	

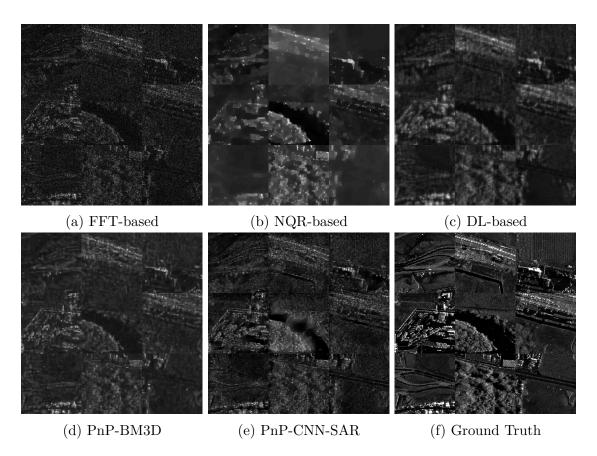


Figure A.36 Reconstruction results for Images 19, 40, 45, 46, 59, 61, 66, 96, and 109 from the Kapıkule test set for the case of 30% data availability and $\sigma_{\mathbf{n}} = \sigma_{\mathbf{y}}$.

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